

Exoplanetary research in Ondřejov

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Astronomical Institute of CAS

- 1898 Ground purchased in Ondřejov
- 1928 Observatory donated to Czech state
- 1954 Observatory became a part of the Czech Academy of Sciences

Scientific cooperation: IAU, ESO (2007), ESA (2008)

~160 employs

Departments:

- Solar physics (spots, oscillations, eruptions, ALMA, Solar orbiter, GREGOR, EST)

- **Stellar physics** (hot stars, stellar atmospheres and winds, astroinformatics, exoplanets)

- **Interplanetary matter** (European and desert firebal network, asteroids)

- Galaxies and planetary systems (dynamics and kinematics of galaxies, interstellar matter, formation and fade of stars)



Science, public outreach, teaching activities

www.asu.cas.cz

Exoplanet group

P. Kabáth (head, petr.kabath@asu.cas.cz), T. Klocová (postdoc), M. Skarka (postdoc), E. Plávalová (postdoc), M. Blažek (PhD student), D. Dupkala, J. Dvořáková (BSc students)

- **RV spectroscopic follow-up of K2/TESS/PLATO candidates** •
- Exoplanetary atmospheres photometry, transmission • spectroscopy
- Stellar activity and exoplanets
- PLATO space mission ground based support ٠





European Southern Observatory



+ close cooperation with Tautenburg observatory

www.eso.org

R_p/R_{\star}

Kabáth et al., in prep.



Perek telescope – Coude focus



Perek telescope – command center





- Installed in 2007 (Koubský et al. 2007)
- R~44000, 360-950 nm
- Commercial photo lens
- Nitrogen-cooled
- Environment at ~21 Celsius degree
- Aluminium assembly
- ThAr lamp for calibration





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- Magnitude limit at 13 mag







Radial velocity determination (Doppler shift of the lines)





$$\frac{\Delta\lambda}{\lambda} = \frac{v}{c}$$

To reach sub-km/s accuracy of RVs we adopted these steps:

- Scientific spectrum is extract using narrow aperture
- Use only one ThAr calibration spectrum taken at the end of night is used
- For each scientific frame a calibration spectrum is extracted from ThAr frame
- Spectra are shifted using telluric lines in a narrow region
- All spectra are corrected of the motion of the Earth (29.7 km/s) and Earth rotation (460 m/s)
- With such approach we are able to get
 - RMS~80 m/s over one night
 - RMS~110 m/s over one month
 - RMS~350 m/s over one year



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Radial velocity standard HD 109358, G0V, V=4.3 mag



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Kabáth et al. 2018, submitted to PASP



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Future plans

- Employ iodine cell
- Employ OPERA
- Get involved in an international obs program(s), PLATO consortium
- Build a spectrograph at La Silla (PLATOSpec)
- Establish an exoplanet community in Czech Republic

Thank you for your attention

