

Dynamic fibrils in DOT H α observations

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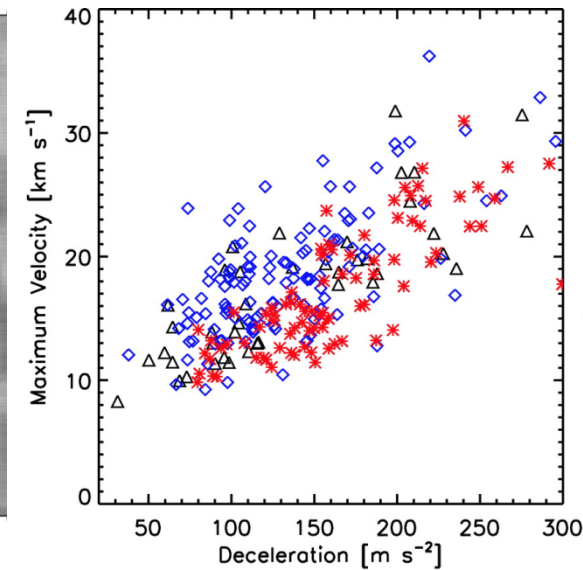
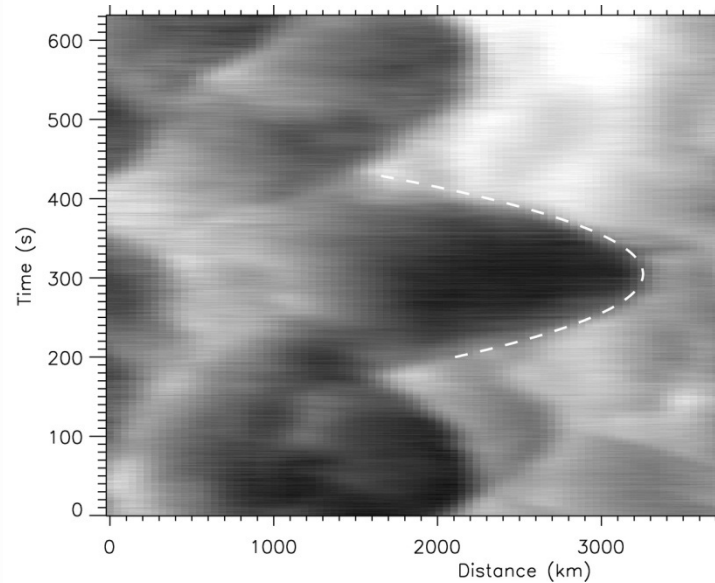
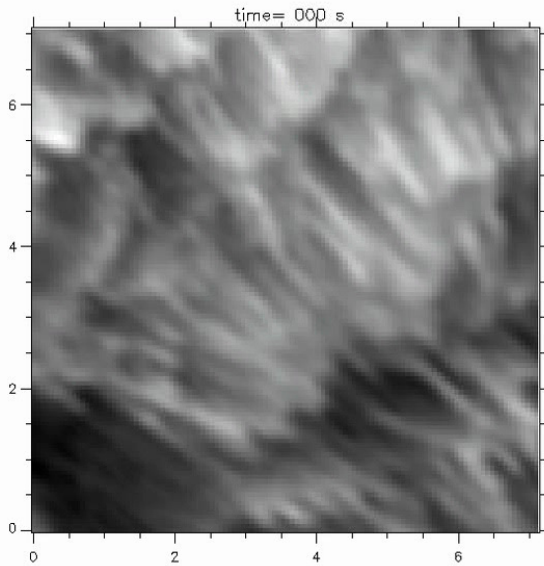




Searching an order in the mess

Dynamic fibrils in H α

- extensions and retractions
- parabolic top trajectories
- positive correlation of maximum velocity and deceleration
- field-aligned magnetoacoustic shock excitation



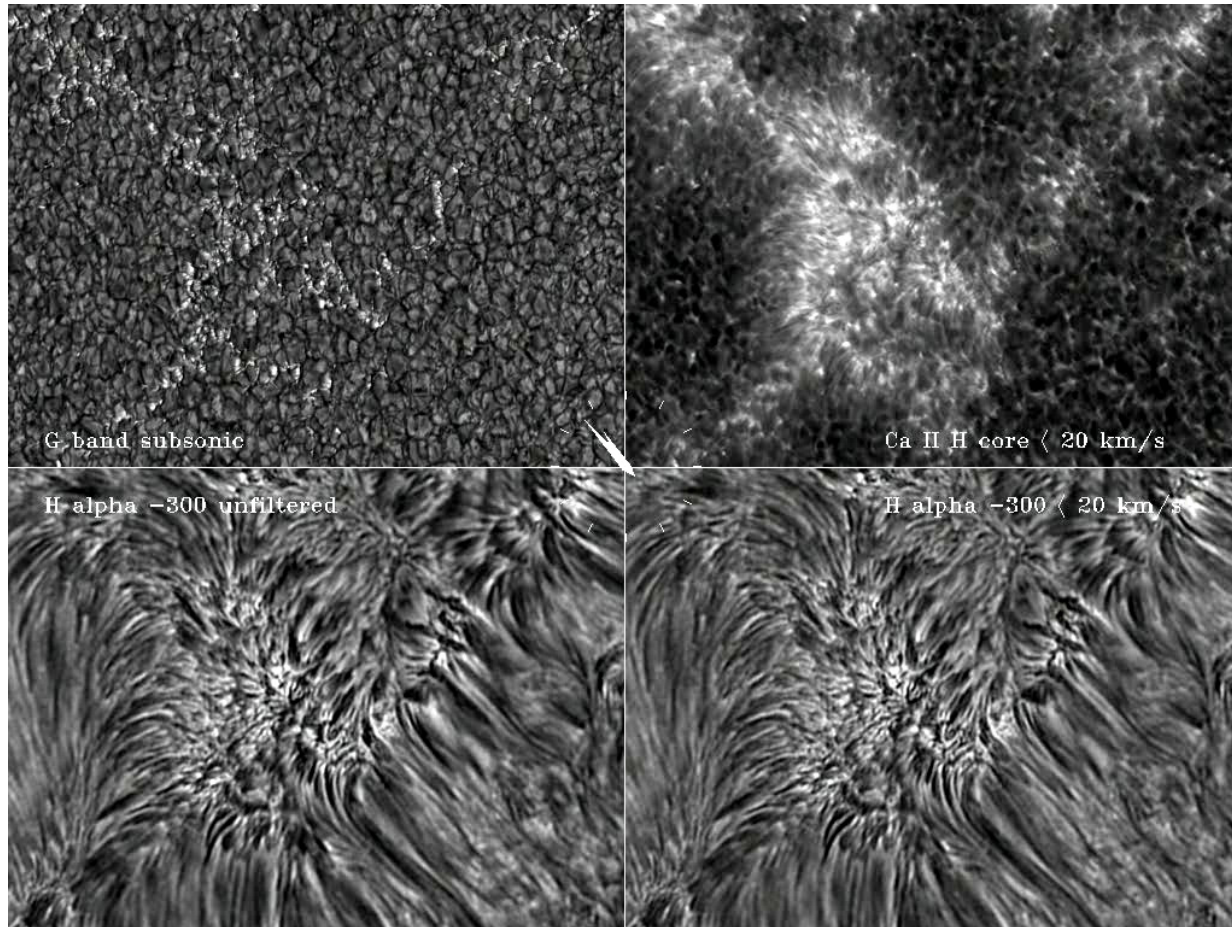
Swedish 1-m Solar Telescope

[Hansteen et al.: 2006, ApJ, 647, L73](#) , [De Pontieu et al.: 2007, ApJ, 655, 624](#)

An aim of this study

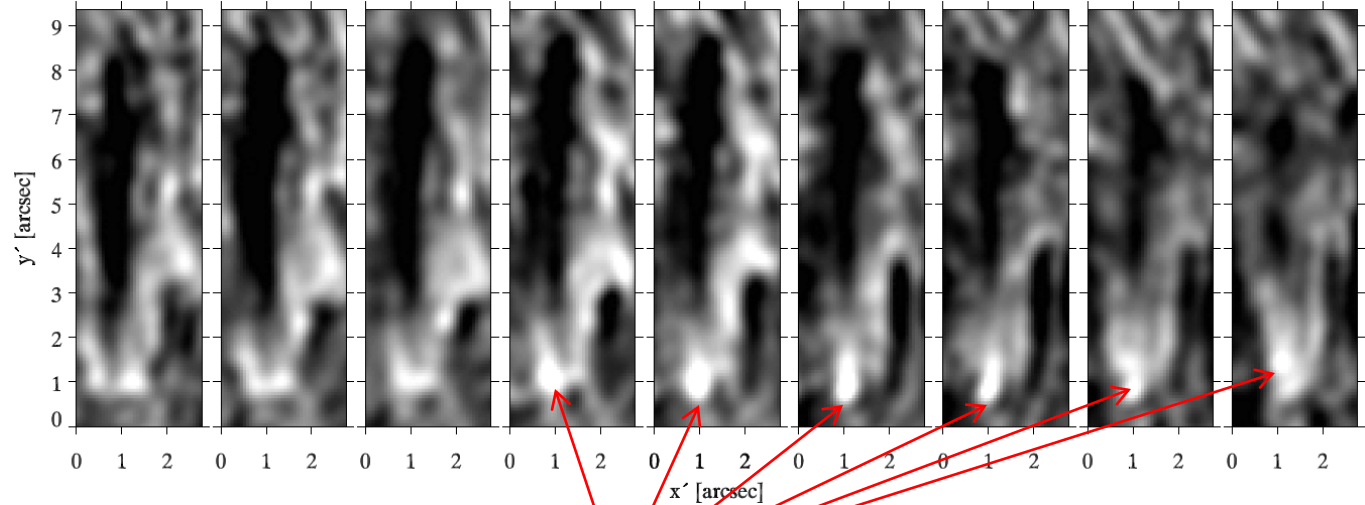
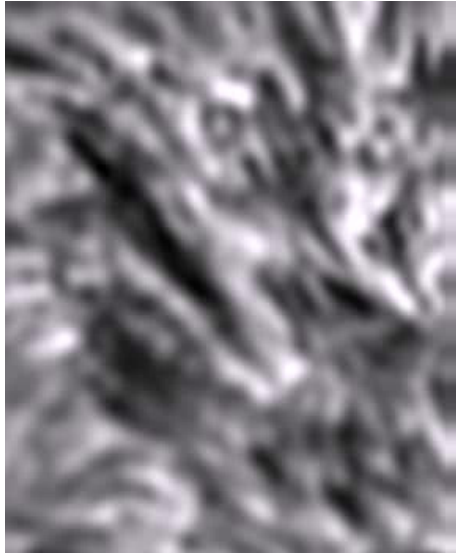
To find and study dynamic fibrils in H α image sequence obtained by the Dutch Open Telescope (DOT).

observations on 24 April 2006



- target: a small plage and network in quiet-Sun area 40° off the disk center
- observations: 10-min speckle-reconstructed image sequences of 50 images
- field of view: $83 \text{ arcsec} \times 66 \text{ arcsec}$
- key diagnostics: $H\alpha$ images taken by a Lyot filter (FWHM 0.25 \AA) at -0.3 \AA off the line center
- time resolution: 12 s

H α dynamic fibrils (DFs) in DOT observations



A possible **brightening** at the fibril's base?
 $\Delta t = 24$ s, fibril's lifetime ≈ 3.5 min

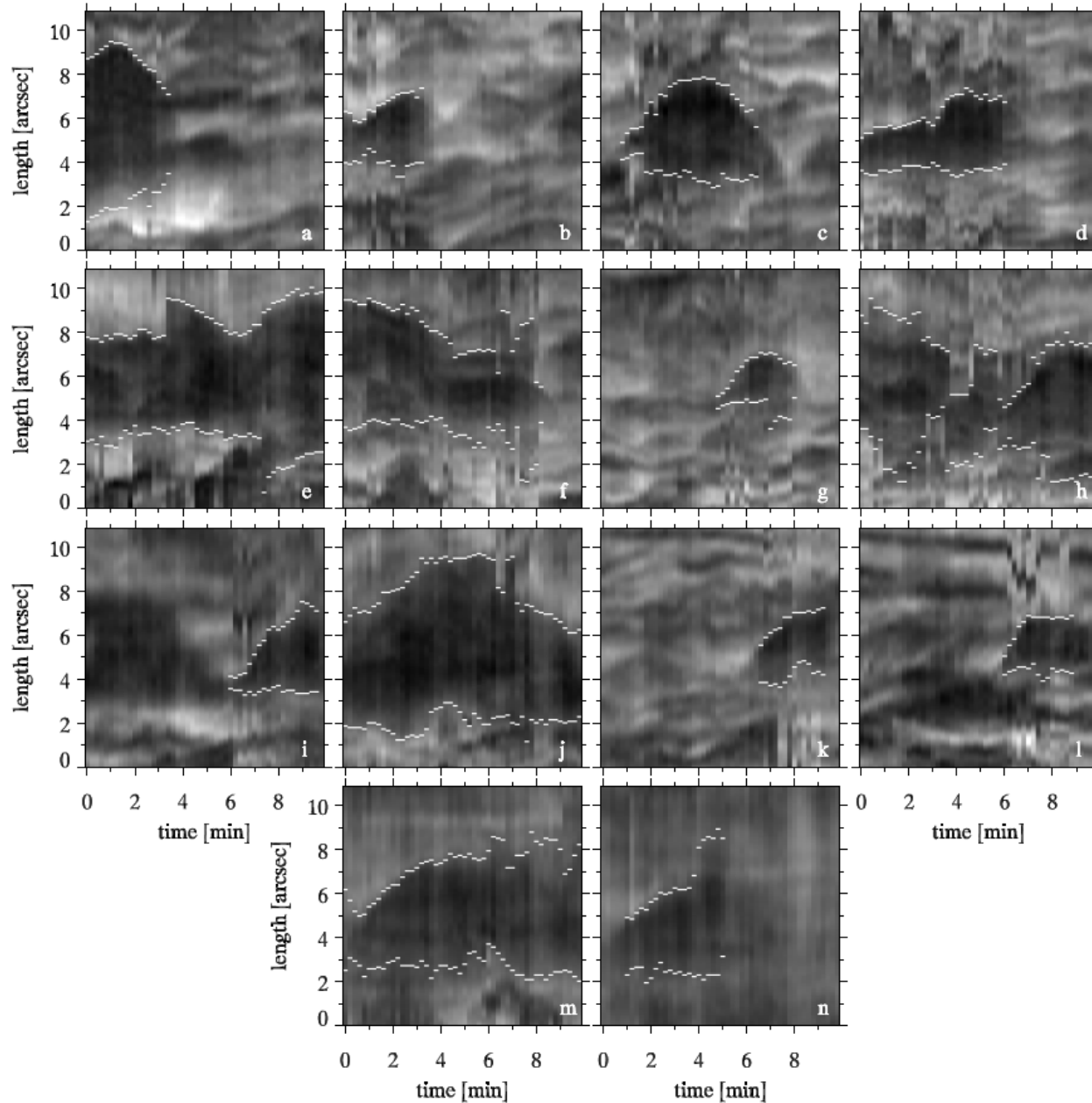
Measurements of DFs:

1. top trajectories
2. temporal variations in orientation



A parallel evolution of two DFs?

Top trajectories of DOT H α dynamic fibrils



N-shaped magnetoacoustic shocks

acoustic cutoff period

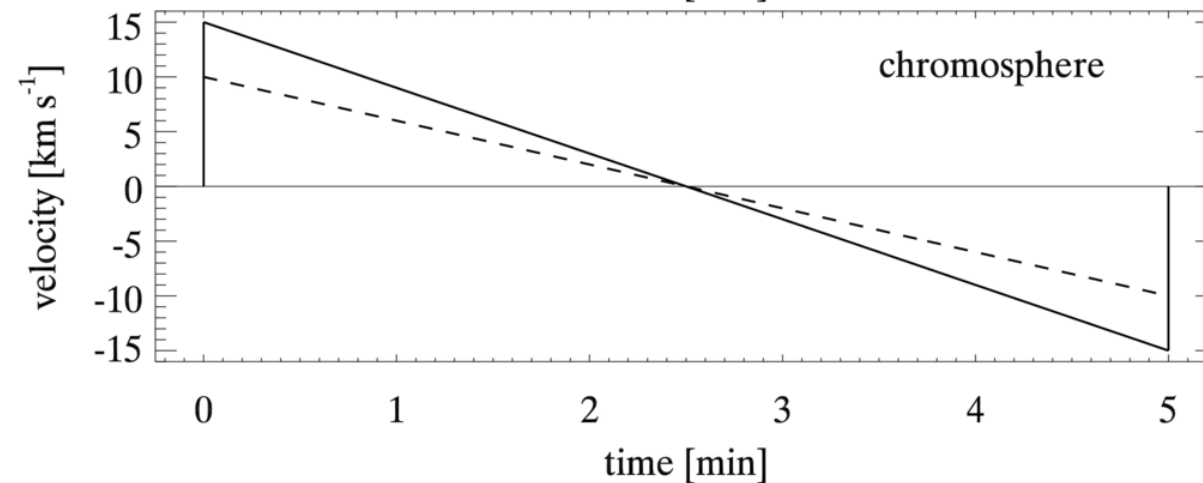
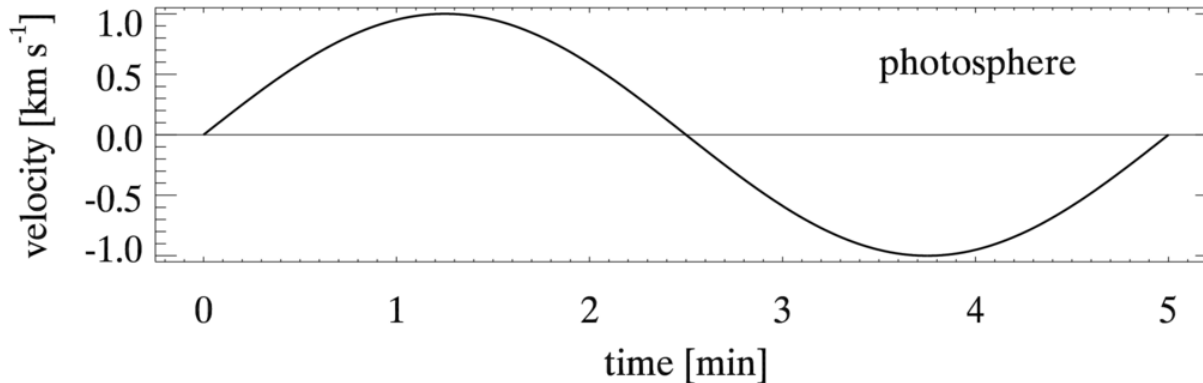
$$P_{ac} = \frac{4\pi}{g \cos \theta} \sqrt{\frac{RT}{\gamma\mu}}$$

from shock discontinuity to parabola

$$v = v_{max} - at$$

$$y = y_0 + v_{max}t - \frac{a}{2}t^2$$

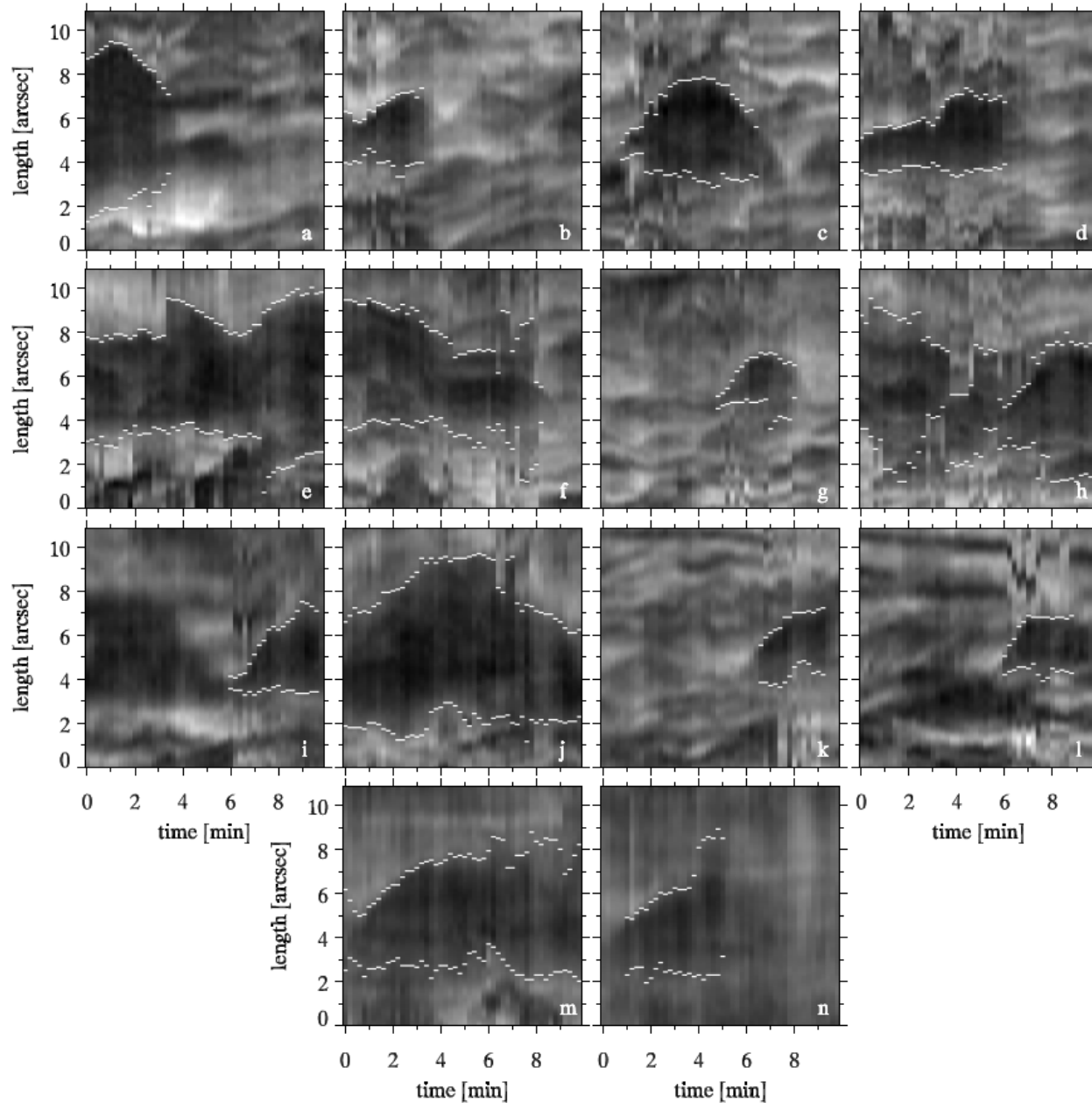
$$v_{max} = \frac{P}{2}a$$



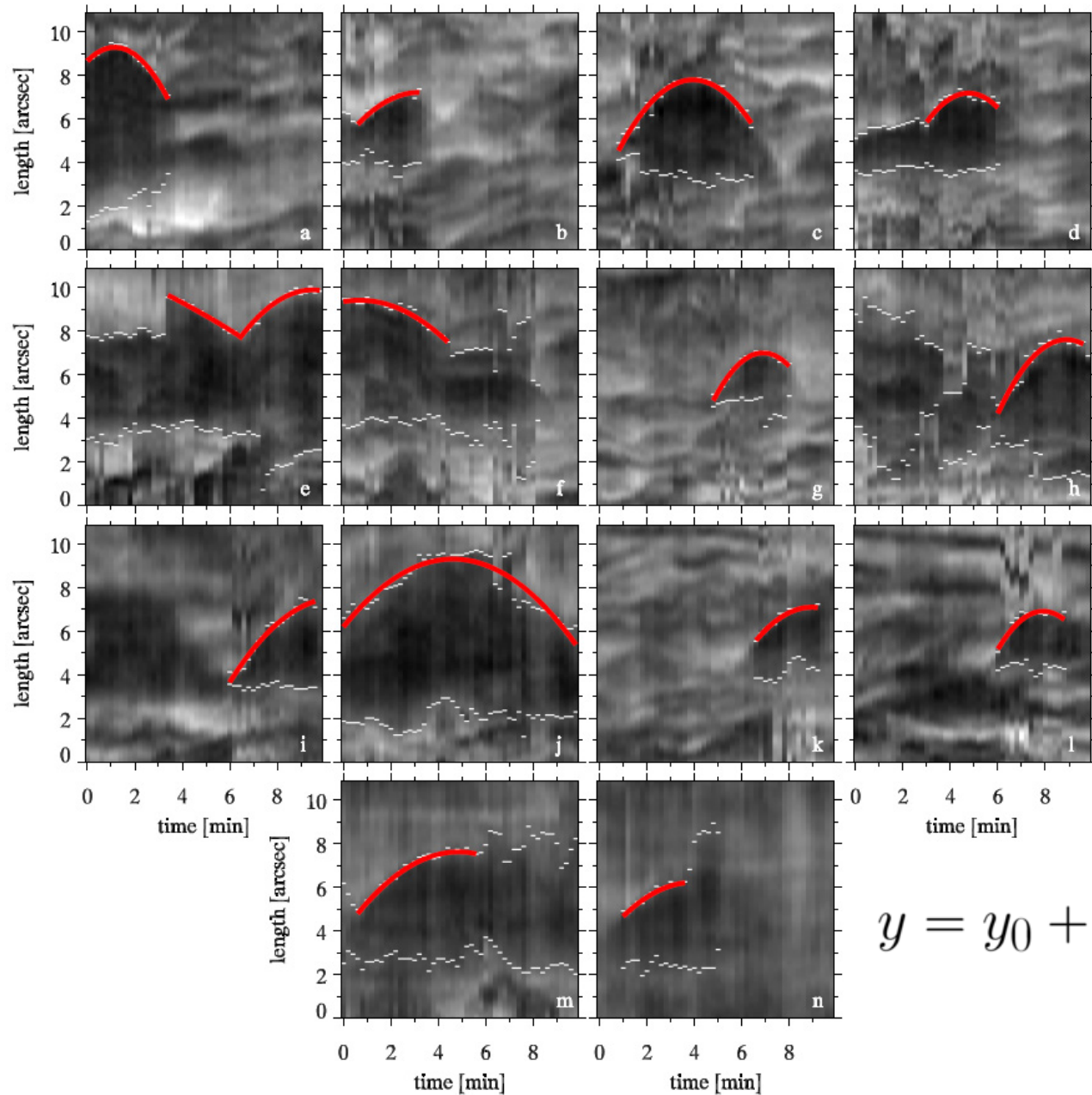
Reduction of the effective gravity $g.\cos\theta$ along inclined magnetic flux tubes:

- ⇒ increasing of the acoustic cutoff period P_{ac} , *i.e.*, lowering of the cutoff frequency
- ⇒ propagation of p-modes into the chromosphere as N-shaped shocks
- ⇒ lift of the chromosphere-transition region interface seen as a fibril

Top trajectories of DOT H α dynamic fibrils



Parabolic fits of top trajectories

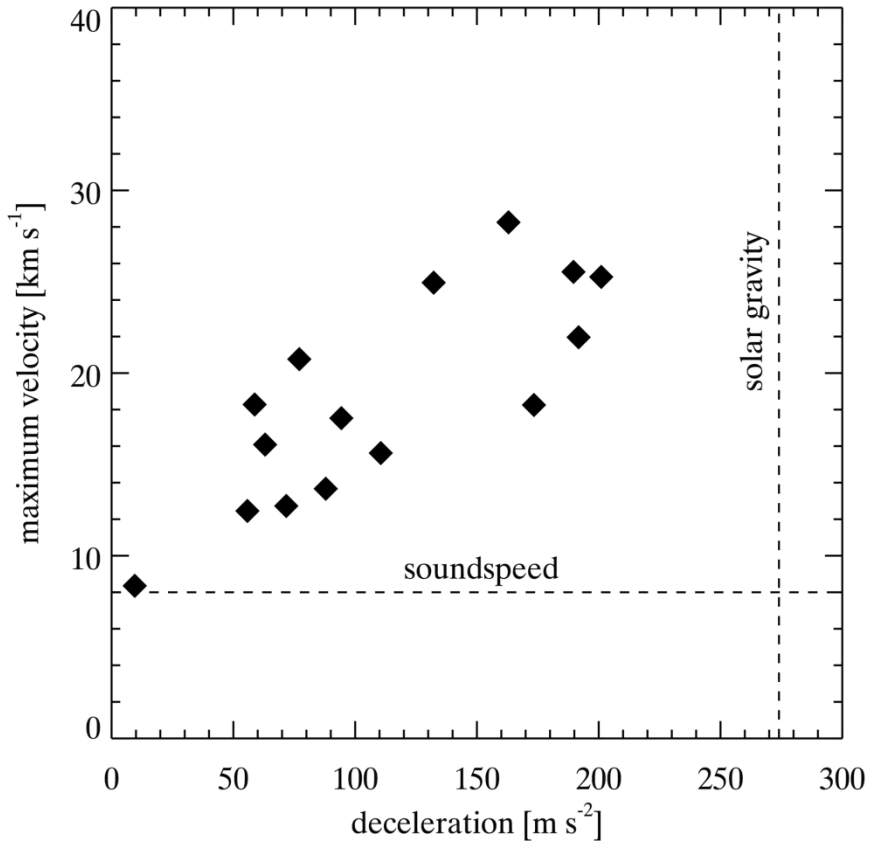


$$y = y_0 + v_{\max}t - \frac{a}{2}t^2$$

Kinematic characteristics

Positive correlation of the maximum velocity and the deceleration

No correction for projection effects.



$$v_{\max} = \frac{P}{2}a$$

The linear relationship is free of projection effects.

a deceleration

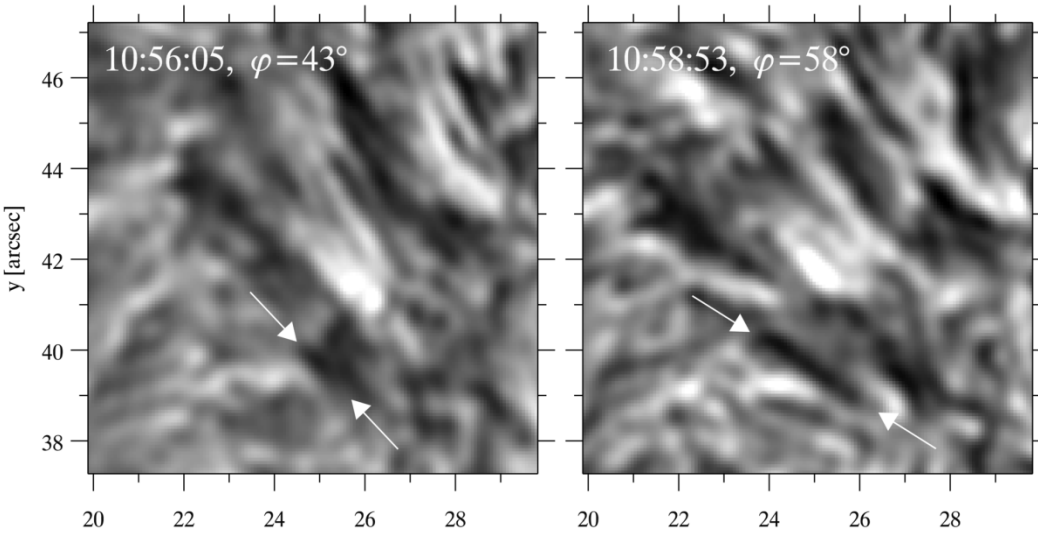
v_{\max} maximum (or onset) velocity

P duration of parabolic motion, i.e.,
lifetime of dynamic fibril

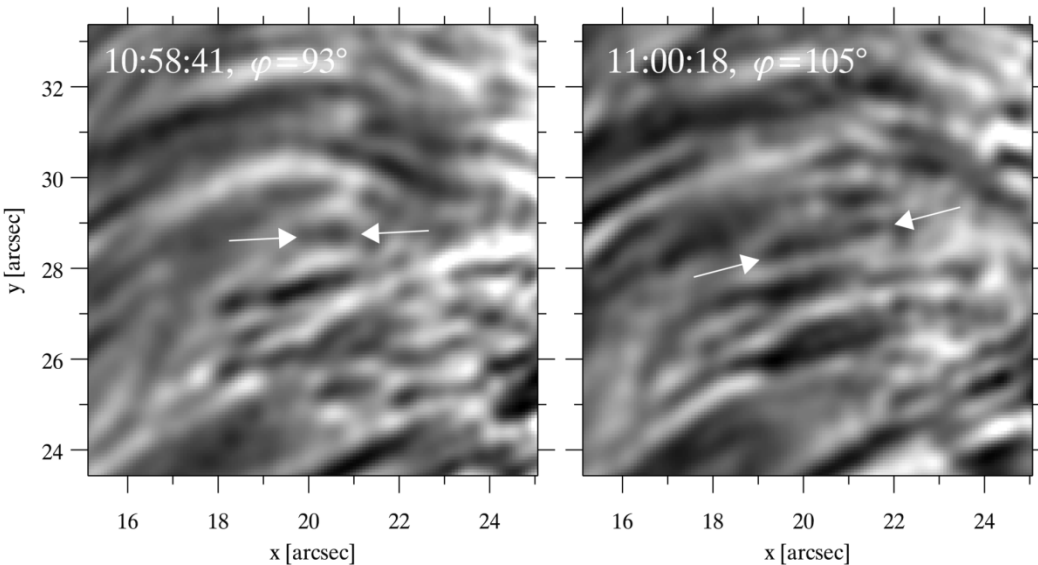
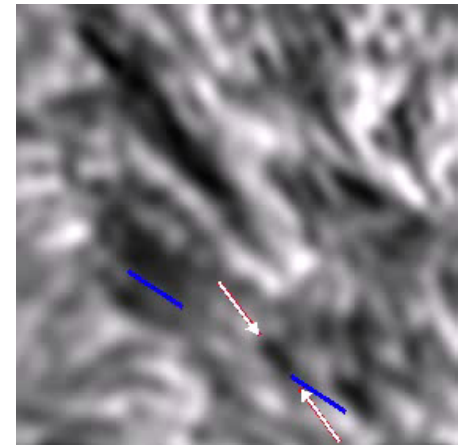
Dynamic fibrils

- supersonic, average max. velocity $\approx 19 \text{ km s}^{-1}$
- sub-ballistic, average deceleration $\approx 110 \text{ m s}^{-2}$

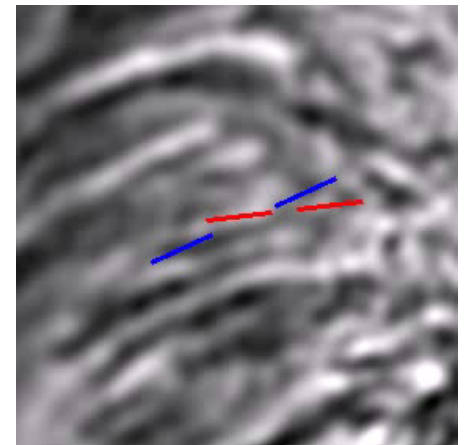
Temporal variations in fibril orientation



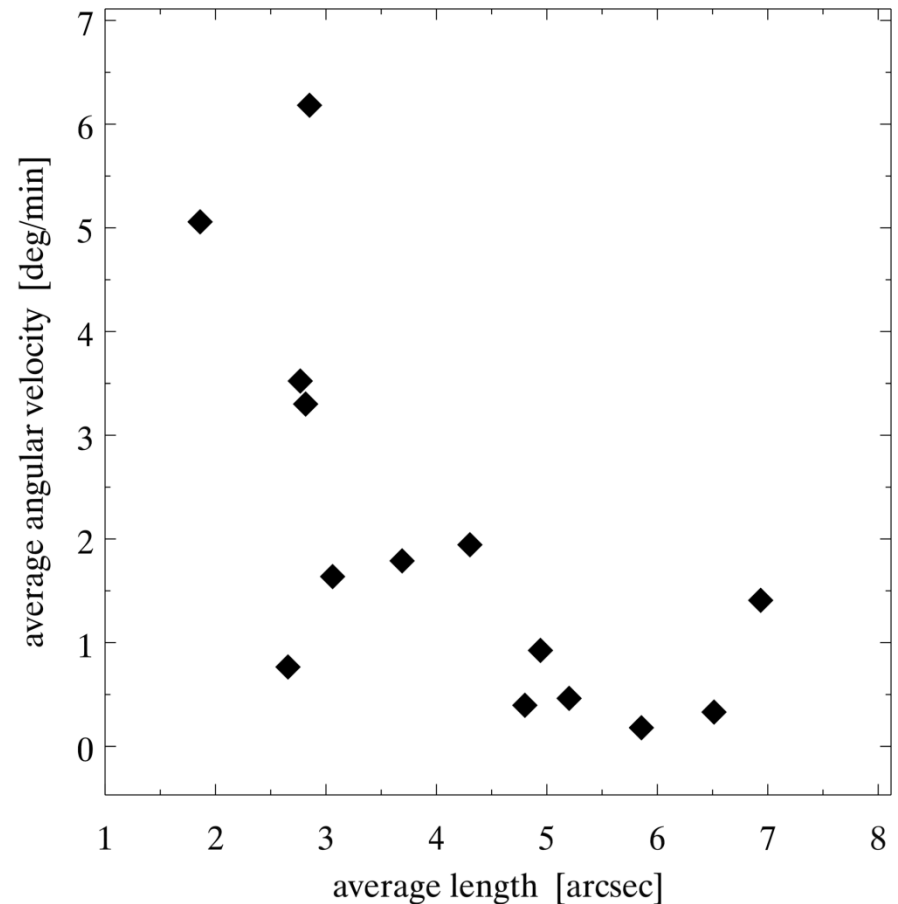
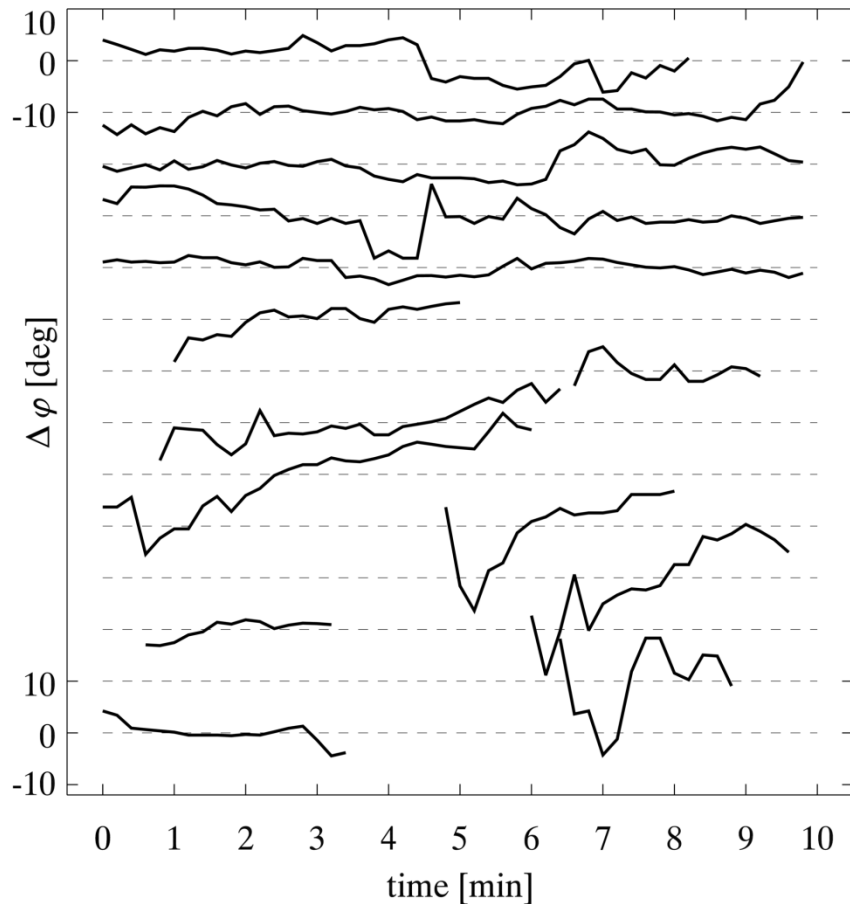
$\Delta t = 2,8 \text{ min}$
 $\Delta\varphi = 15^\circ$
 $\omega = 5,4^\circ/\text{min}$



$\Delta t = 1,6 \text{ min}$
 $\Delta\varphi = 12^\circ$
 $\omega = 7,5^\circ/\text{min}$



Temporal variations in fibril orientation



- the variations indicate turning motions with angular speeds of the order of $1^\circ/\text{min}$
- shorter fibrils tend to turn faster than longer ones

Conclusions

- tops of dynamic fibrils follow parabolic trajectories with:
 - supersonic maximum velocities of $\approx 19 \text{ km s}^{-1}$ in average
 - sub-ballistic decelerations of $\approx 110 \text{ m s}^{-2}$ in average
- the positive correlation of maximum velocities and decelerations supports the idea of the magnetoacoustic shock excitation
- temporal variations in orientation indicate turning motions with angular speed of the order of 1 deg min^{-1}
- shorter dynamic fibrils tend to turn faster than longer ones

What next

- to identify counterparts of the $H\alpha$ dynamic fibrils in the Ca II H images
- a need of new fibril measurements both in $H\alpha$ and Ca II H due to different data formats

