

Statistical modelling of prominence fine structure using the 2D multi-thread nonLTE model

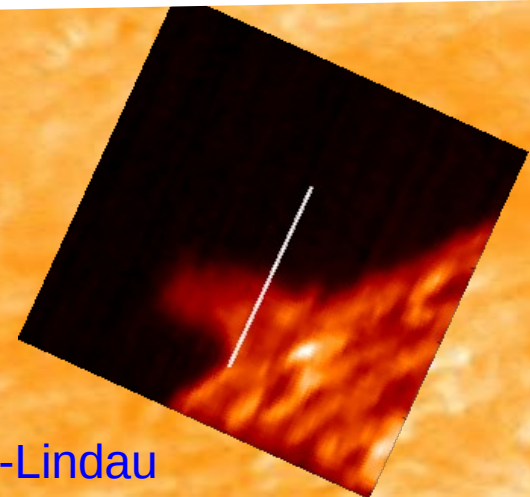
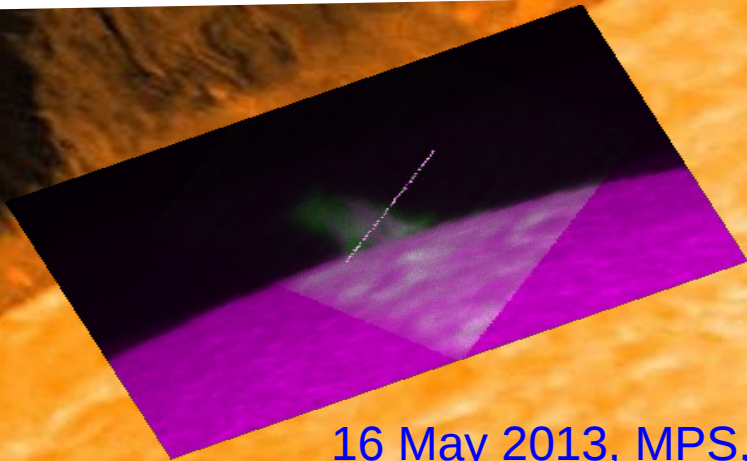
P. Schwartz^{1,2}, S. Gunár², P. Heinzel², B. Schmieder³, and U. Anzer⁴

¹ Astronomical Institute, Slovak Academy of Sciences

² Astronomical Institute, Academy of Sciences of the Czech Republic

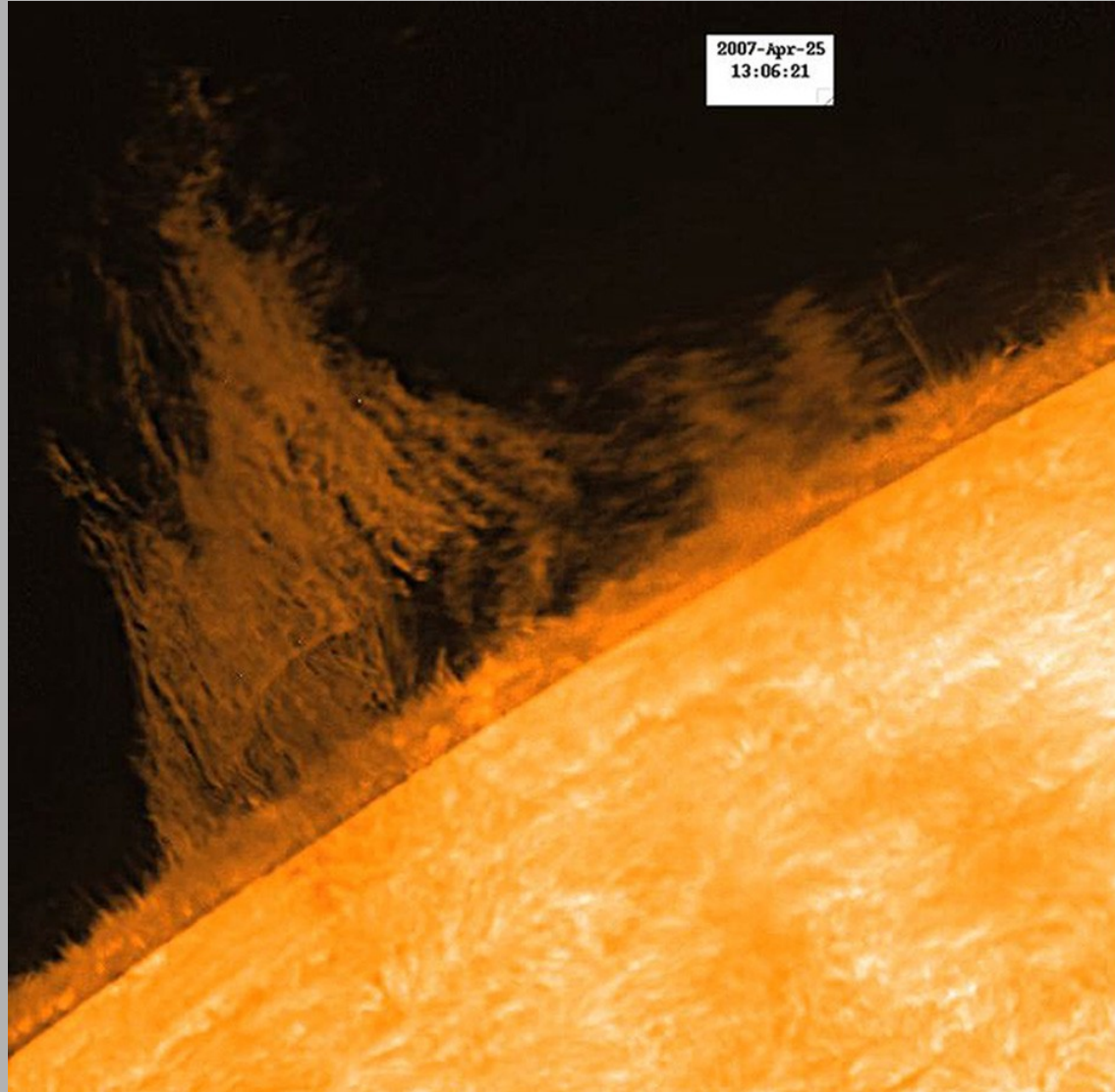
³ Observatoire De Paris, Section Meudon, France

⁴ Max-Planck-Institut für Astrophysik



16 May 2013, MPS, Katlenburg-Lindau

Vertical threads in the quiescent prominence observed by SOT in $H\alpha$



2D prominence fine-structure models

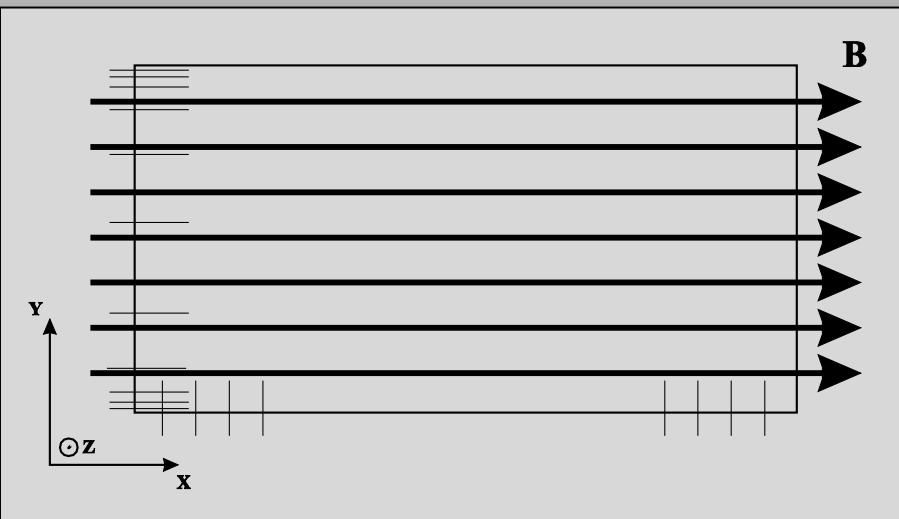
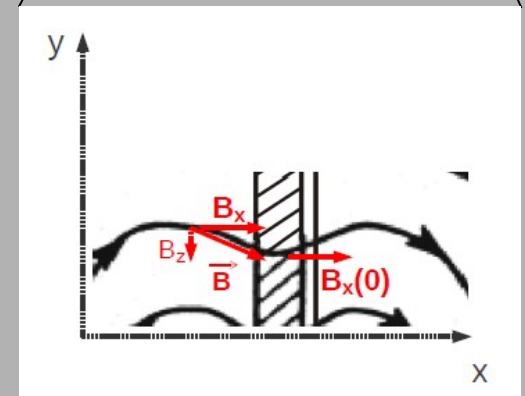
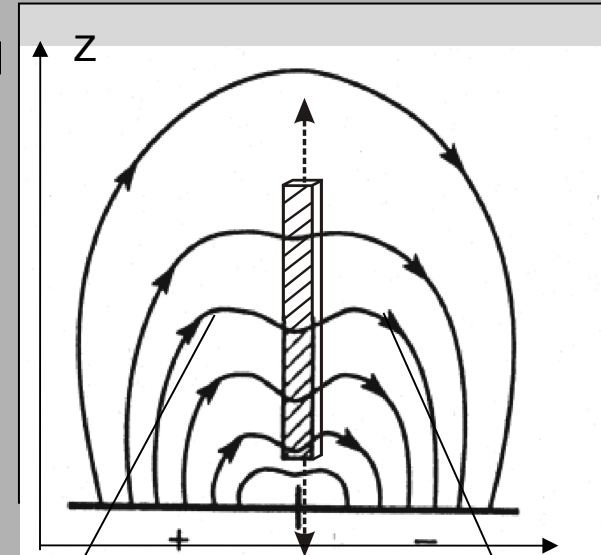
Vertically infinite fine-structure thread embedded in a dipped magnetic field

2D MHS equilibrium of Kippenhahn-Schlüter type (Heinzel & Anzer 2001).

Solution of 2D radiative transfer by MALI method (Auer & Paletou 1994) with usage of the short characteristics (Kunasz & Auer 1988)

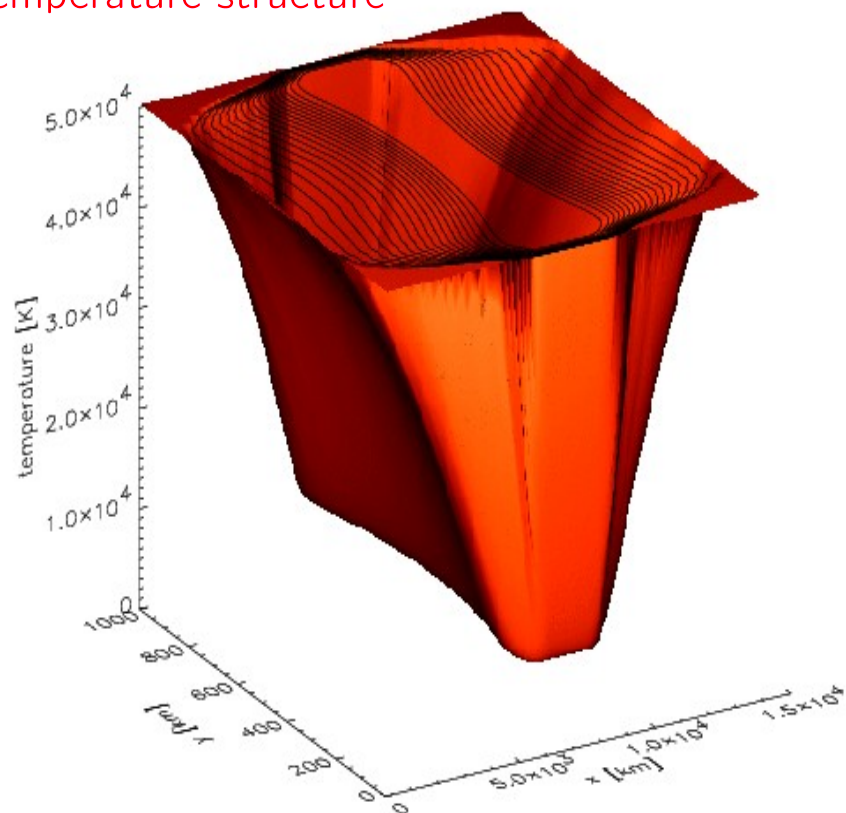
12 level plus continuum hydrogen atom

Partial-frequency redistribution (PRD) is assumed for Lyman- α and Lyman- β lines



- Anzer, U. & Heinzel, P.: 1999, *A&A* **349**, 974.
Auer, L.H. & Paletou, F.: 1994, *A&A* **285**, 675.
Heinzel, P. & Anzer, U.: 2001, *A&A* **375**, 1082.
Kunasz, P. & Auer, L.H.: 1988, *JQSRT* **39**, 67.

Temperature structure



Temperature and density structure

$$T(m, y) = T_{\text{cen}}(y) + [T_{\text{tr}} - T_{\text{cen}}(y)] \left\{ 1 - 4 \frac{m}{M(y)} \left[1 - \frac{m}{M(y)} \right] \right\}^{\gamma_1}$$

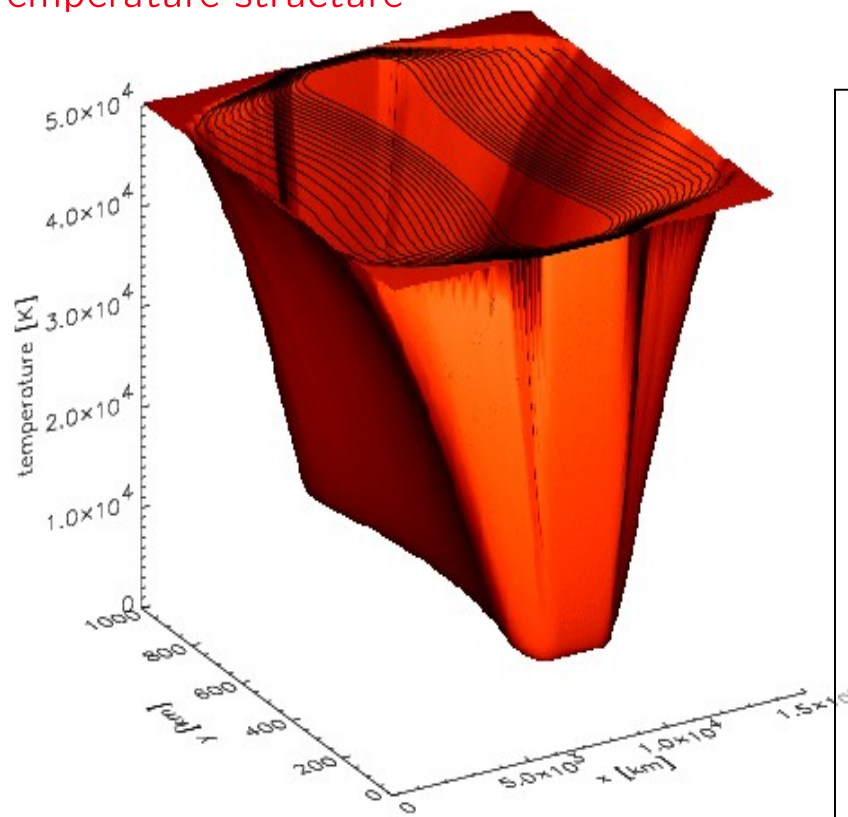
The central temperature $T_{\text{cen}}(y)$

$$T_{\text{cen}}(y) = T_{\text{tr}} - (T_{\text{tr}} - T_0) \left(1 - \left| \frac{y}{\delta} \right|^{\gamma_2} \right), \quad \text{for } |y| \leq \delta$$

$$T_{\text{cen}}(y) = T_{\text{tr}}, \quad \text{for } |y| > \delta$$

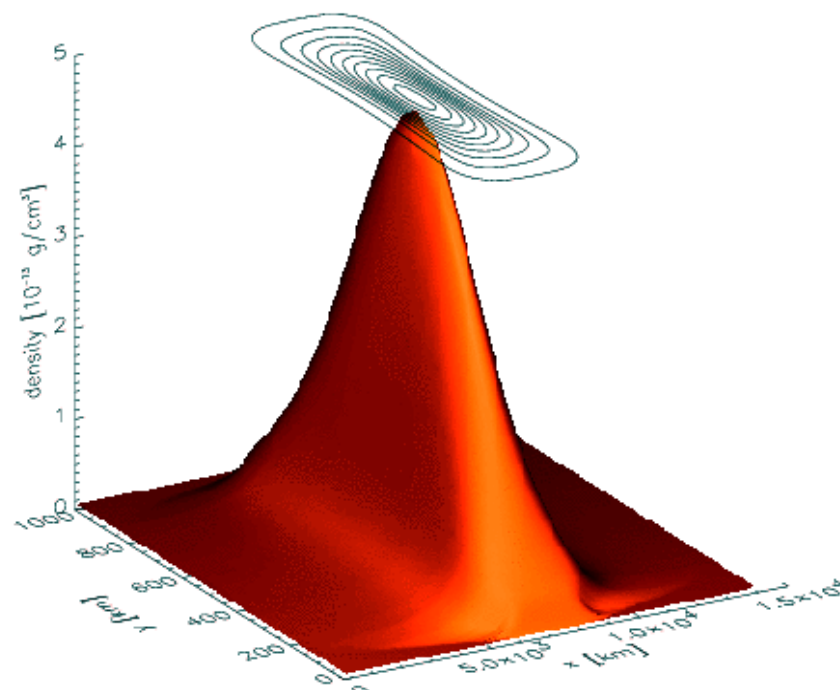
γ_1 and γ_2 are chosen parameters.

Temperature structure



Temperature and density structure

Density structure



$$T(m, y) = T_{\text{cen}}(y) + [T_{\text{tr}} - T_{\text{cen}}(y)] \left\{ 1 - 4 \frac{M(y)}{M_0} \right\}$$

$$M(y) = \begin{cases} M_0 \left(1 - \left| \frac{y}{\delta} \right|^{\gamma_3} \right), & \text{for } |y| \leq \delta \\ 0, & \text{for } |y| > \delta \end{cases}$$

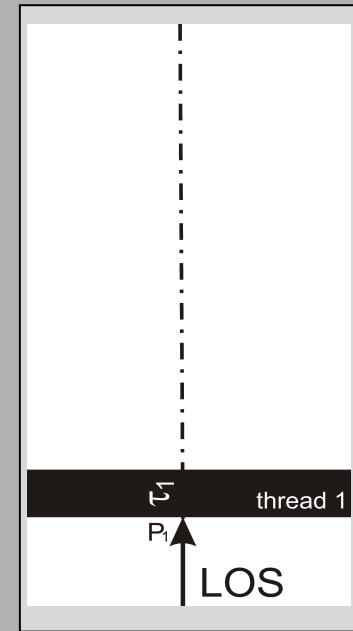
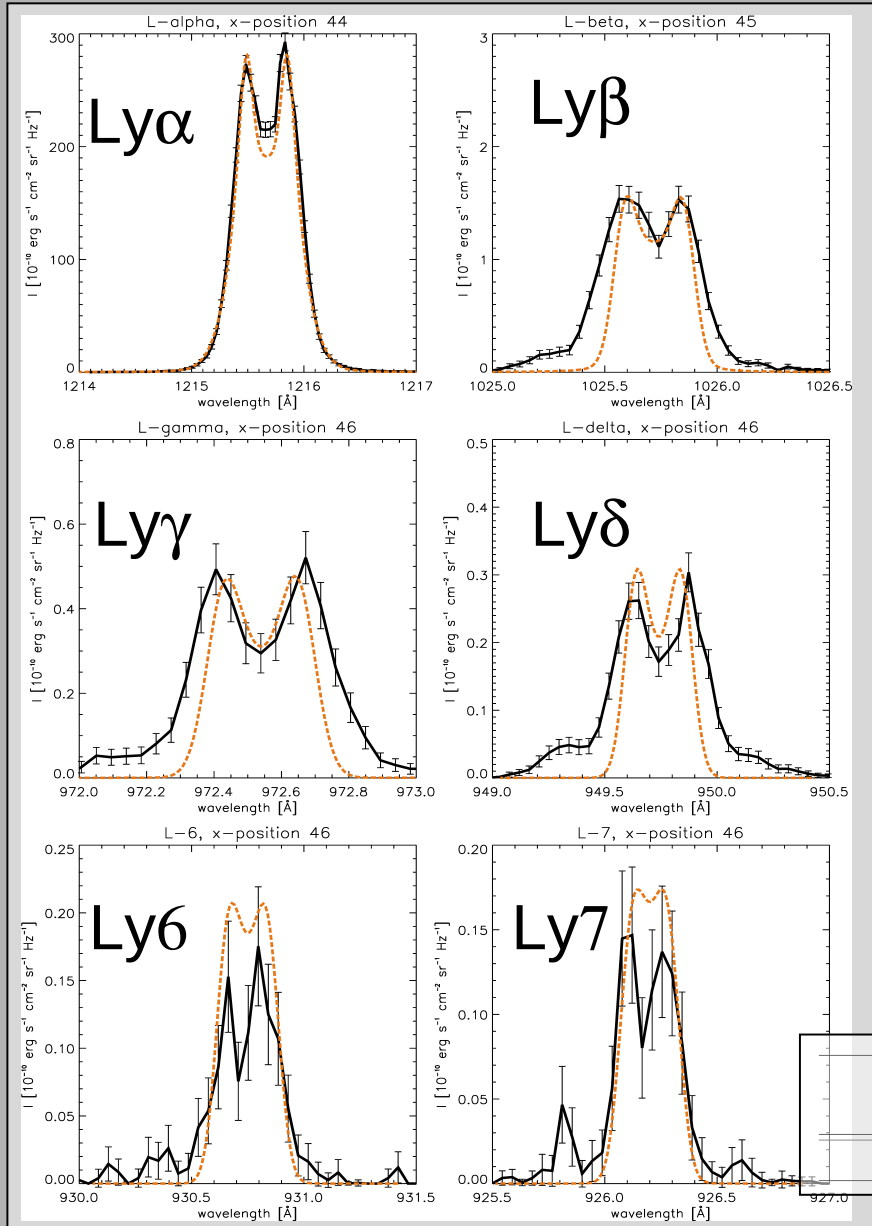
The central temperature $T_{\text{cen}}(y)$

$$T_{\text{cen}}(y) = T_{\text{tr}} - (T_{\text{tr}} - T_0) \left(1 - \left| \frac{y}{\delta} \right|^{\gamma_2} \right), \quad \text{for } |y| \leq \delta$$

$$T_{\text{cen}}(y) = T_{\text{tr}}, \quad \text{for } |y| > \delta$$

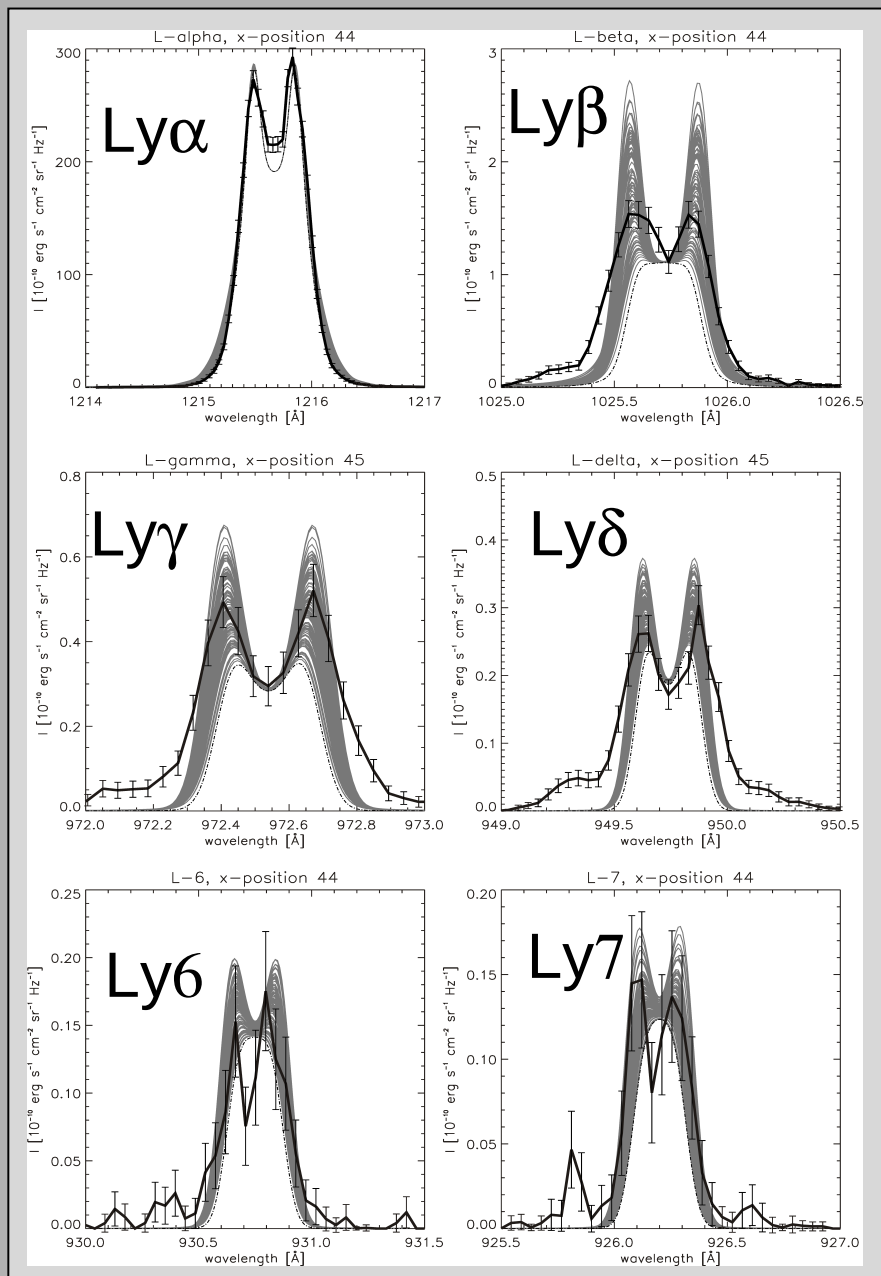
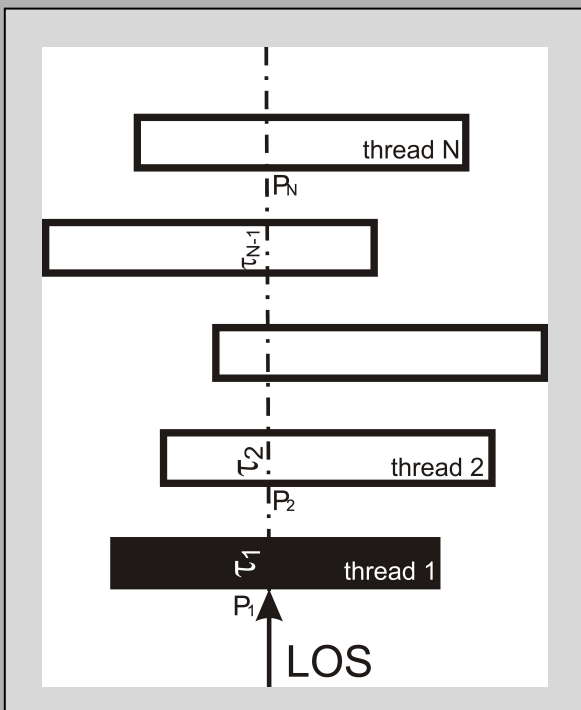
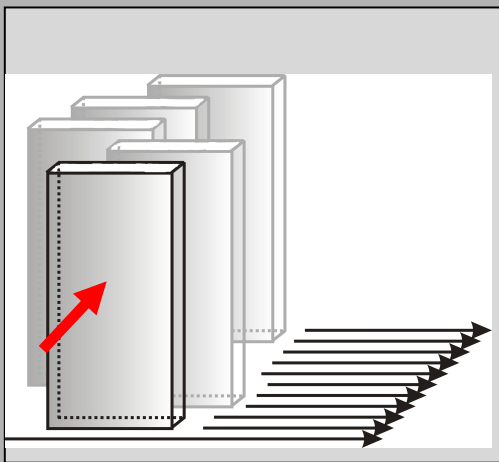
γ_1 and γ_2 are chosen parameters.

Comparison with observed spectra



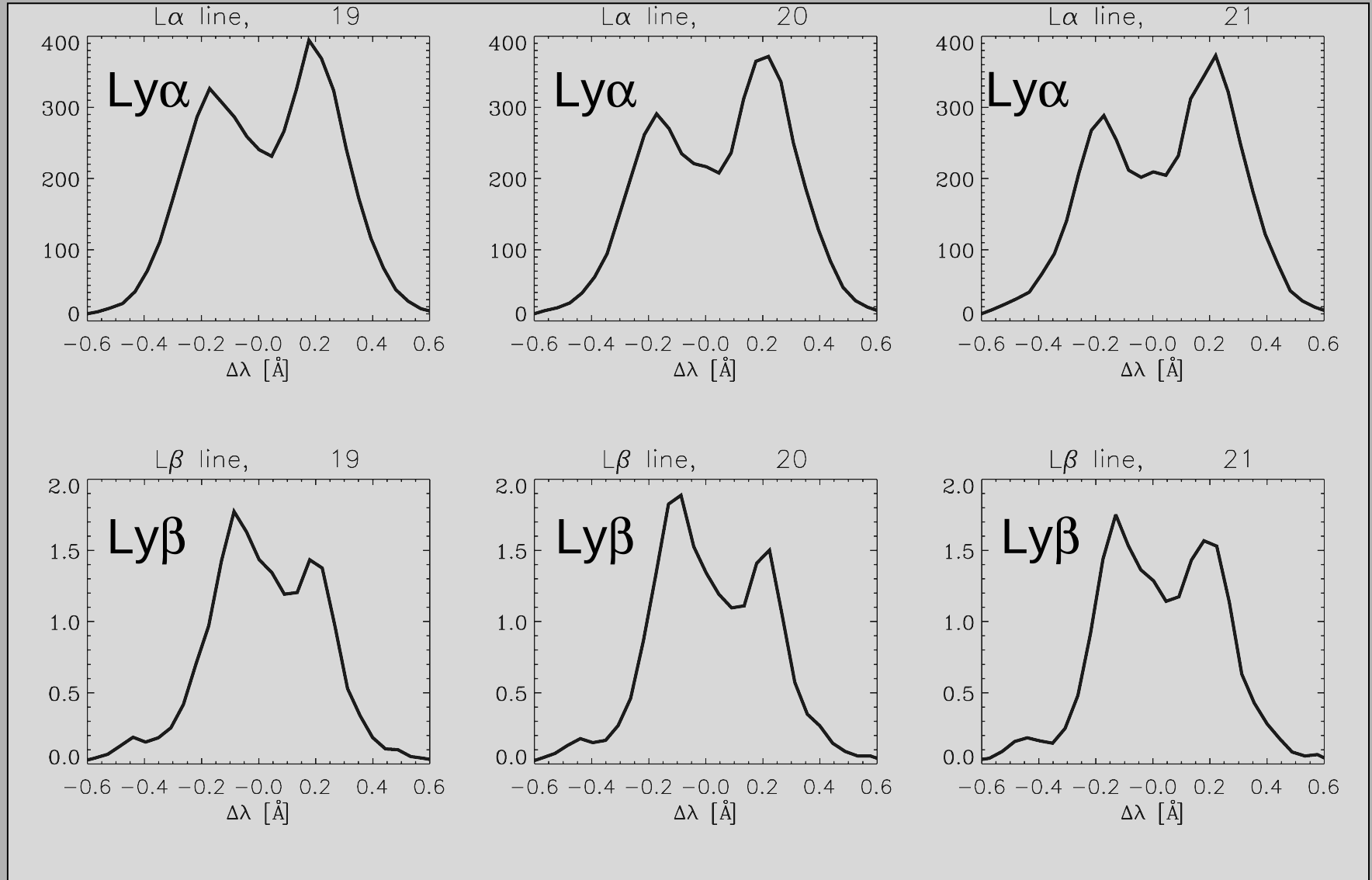
T_0 [K]	M_0 [g cm $^{-2}$]	$B_x(0)$ [Gauss]	p_0 [dyn cm $^{-2}$]	v_t [$\times c_s$]
7000	1.1×10^{-4}	6	0.015	0.5

Comparison with observed spectra



Lyman line asymmetries

Observed profiles – asymmetries of Ly β sometimes opposite to those of Ly α



2D multi-thread models with LOS velocities

The total emerging intensity

$$I_{\text{total}}(\lambda) = \left[\dots \left[\left[I_{\lambda(1)} \times \exp(-\tau_{\lambda(2)}) + I_{\lambda(2)} \right] \times \exp(-\tau_{\lambda(3)}) + I_{\lambda(3)} \right] \times \dots + I_{\lambda(N-1)} \right] \times \exp(-\tau_{\lambda(N)}) + I_{\lambda(N)} ,$$

where $I_{\lambda(n)}$ and $\tau_{\lambda(n)}$ are defined as

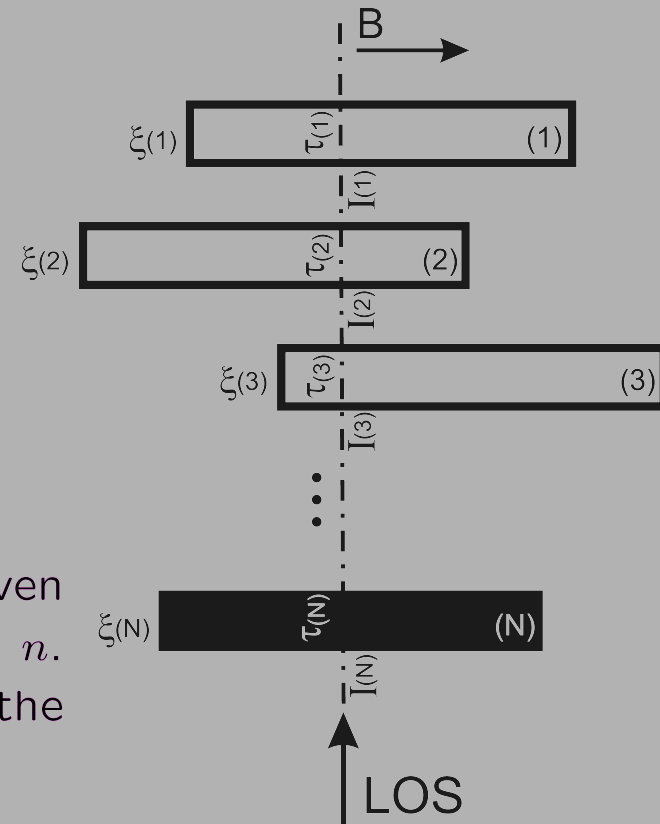
$$I_{\lambda(n)} \equiv I(\lambda - \Delta\lambda^{(n)})$$

$$\tau_{\lambda(n)} \equiv \tau(\lambda - \Delta\lambda^{(n)}) .$$

$\Delta\lambda^{(n)}$ is the Doppler shift for the thread n

$$\Delta\lambda^{(n)} = \lambda_0 \frac{\xi^{(n)}}{c} ,$$

where λ_0 represents the line-centre wavelength of a given spectral line and $\xi^{(n)}$ is the LOS velocity of thread n . Positive values of $\xi^{(n)}$ represent velocities towards the observer.



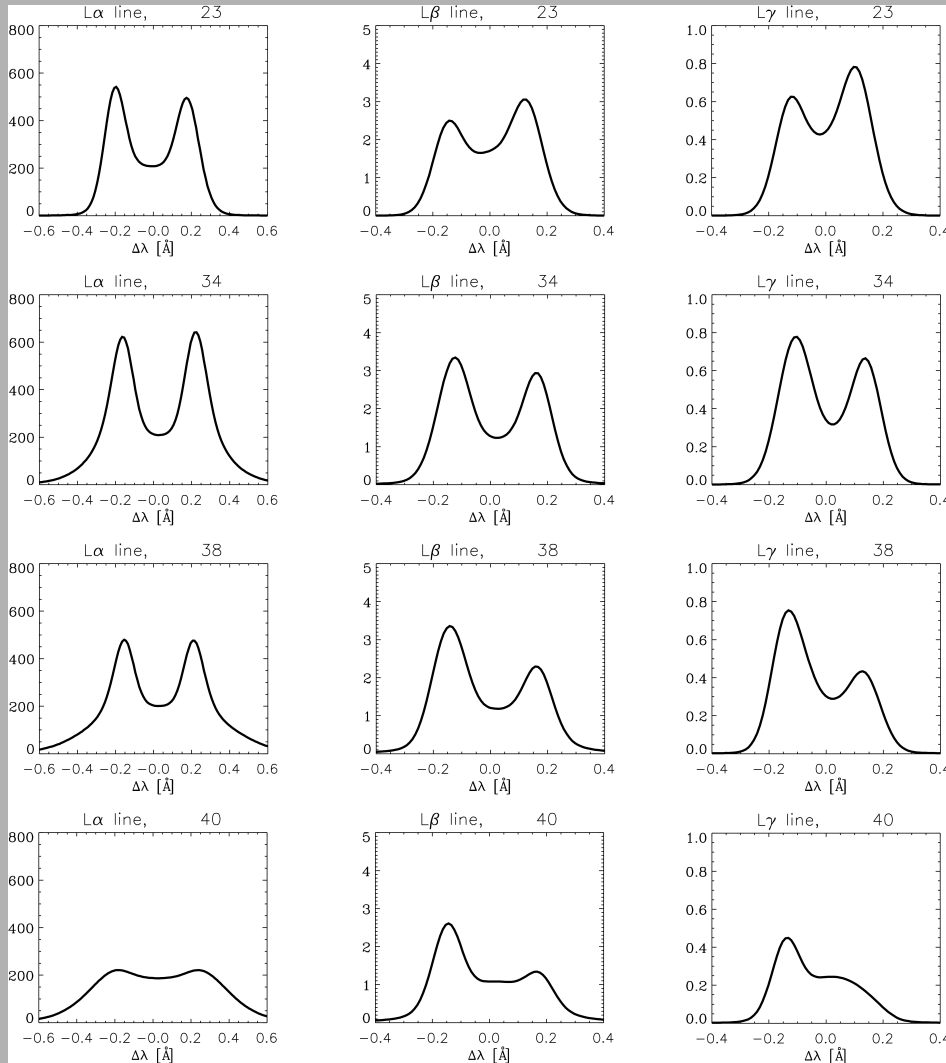
Lyman line asymmetries

synthetic profiles

Ly α

Ly β

Ly γ

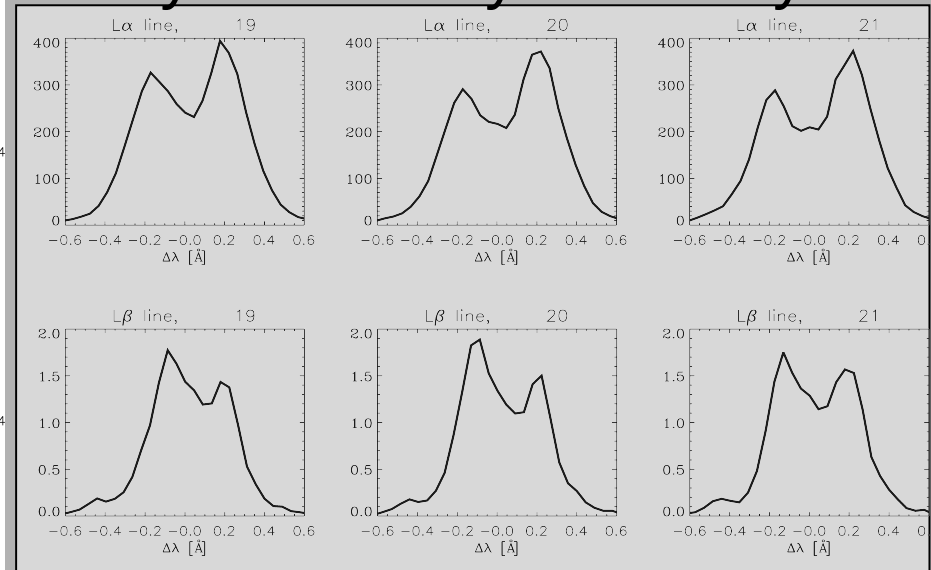


profiles observed by SUMER

Ly α

Ly α

Ly α



Ly β

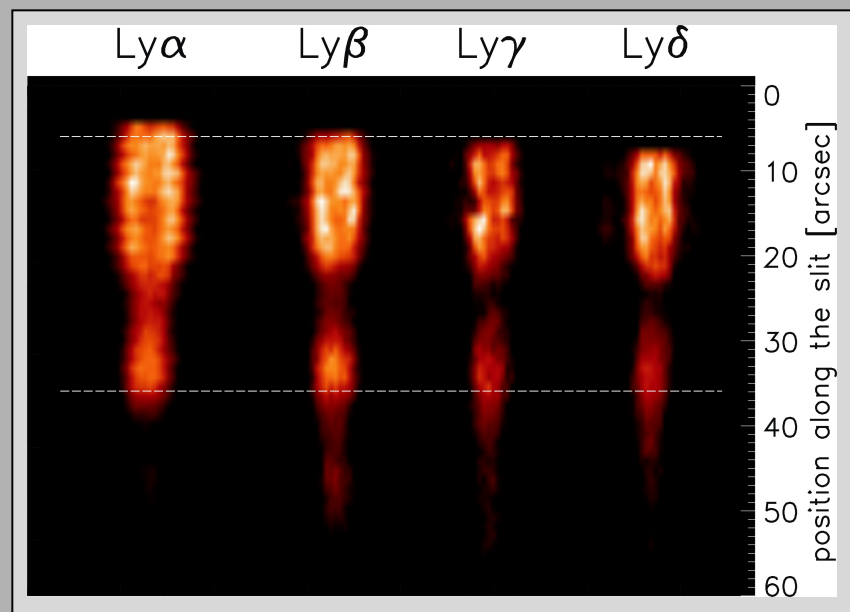
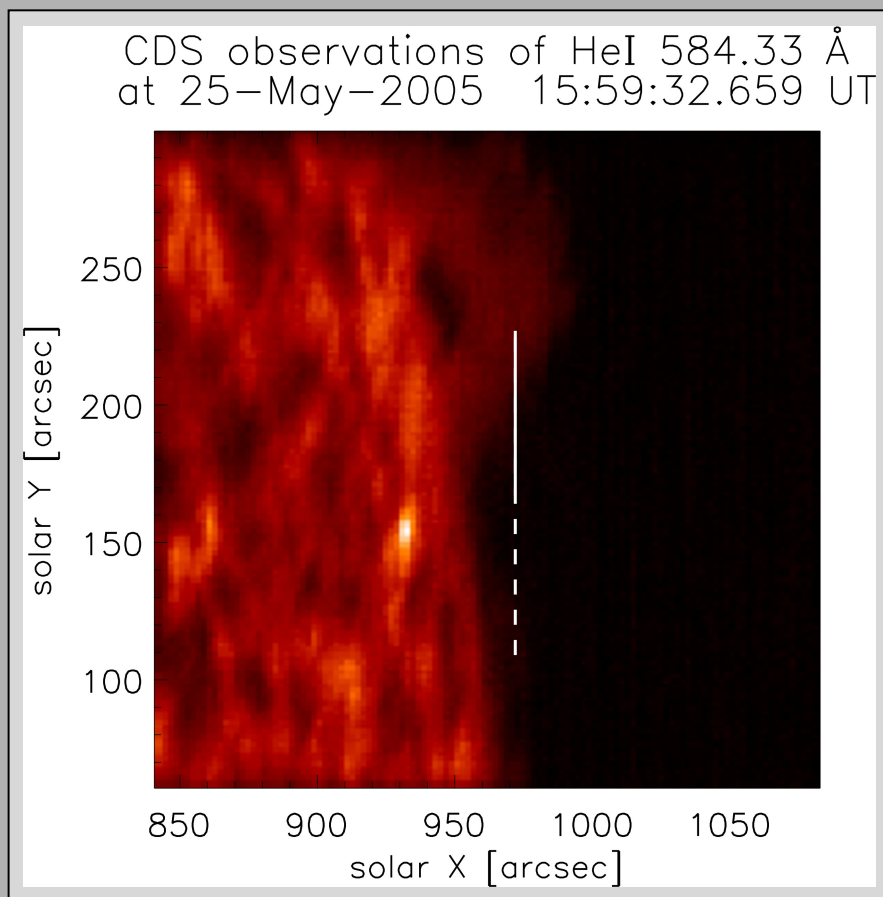
Ly β

Ly β

Statistical comparison

Number of observations within each block, May 25 to 26, 2005

Spectral window	16:03:15 UT to 20:40:59 UT	20:46:27 UT to 23:01:26 UT	23:04:54 UT to 01:21:51 UT
Lyman- α	25	12	13
Lyman- β	40	20	20
Lyman- γ and Lyman- δ	39	20	20
Lyman-5 to continuum	10	10	10

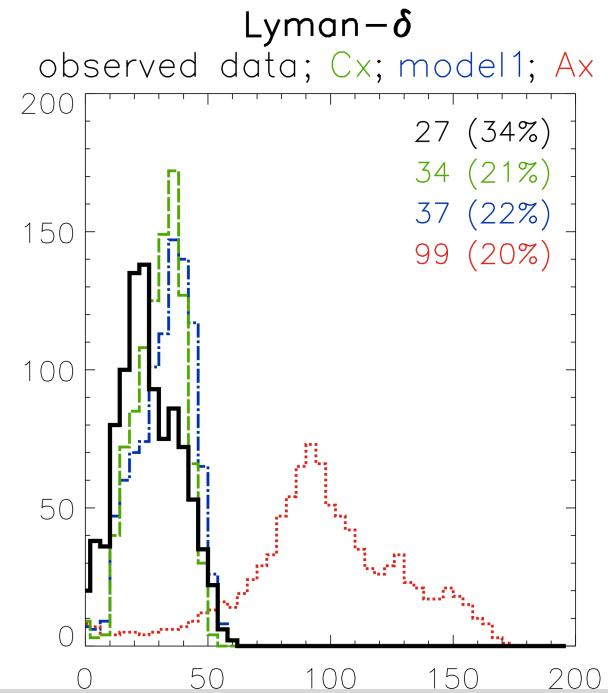
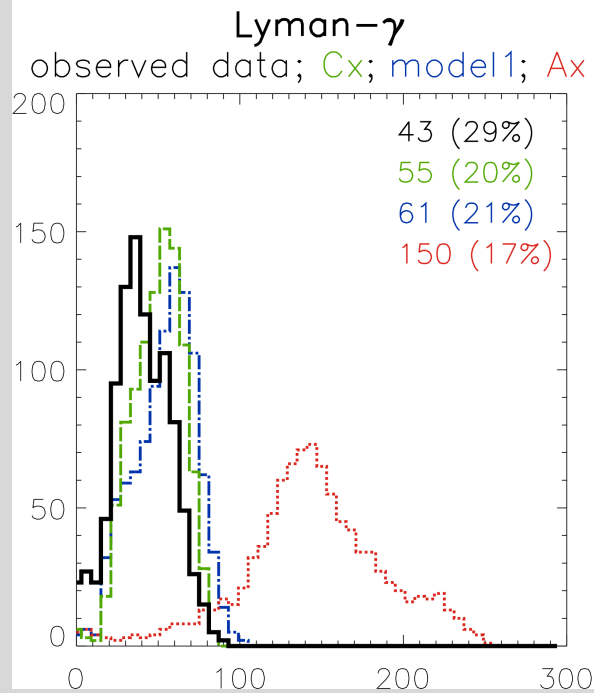
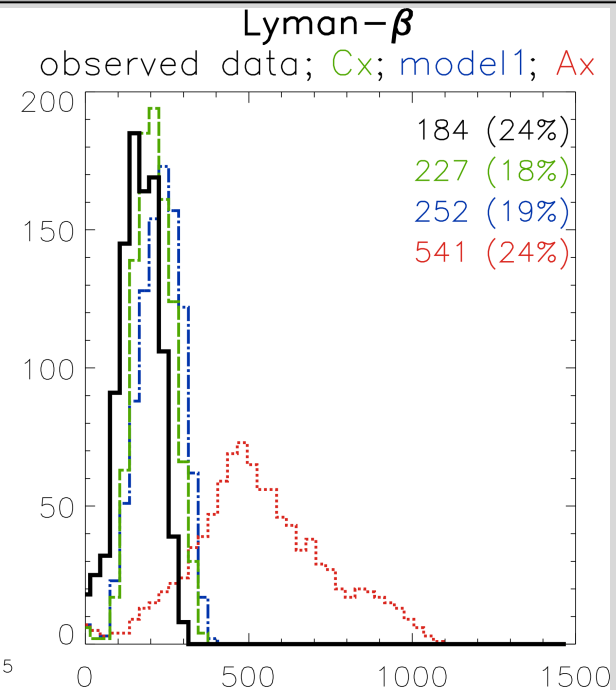
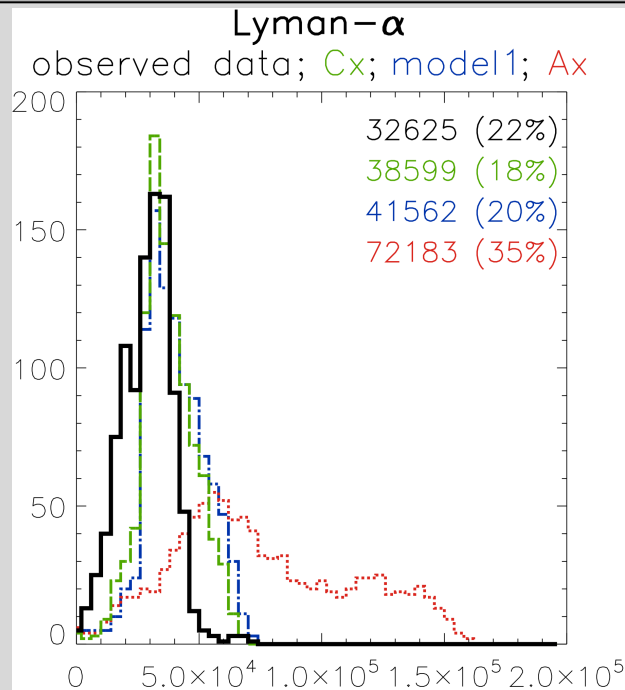
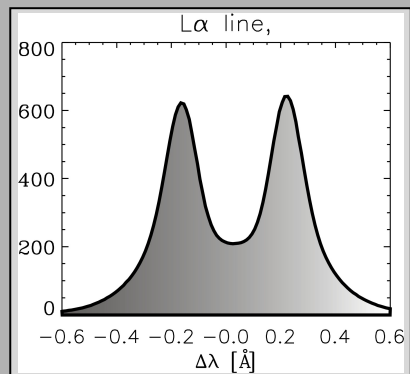


Models used for one thread of multithread model

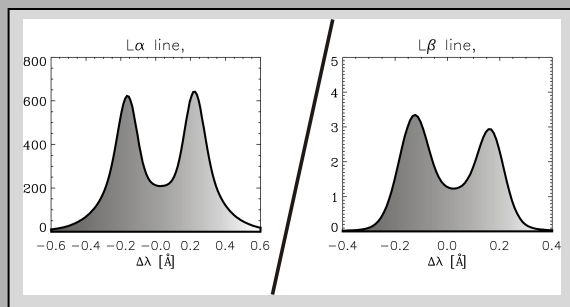
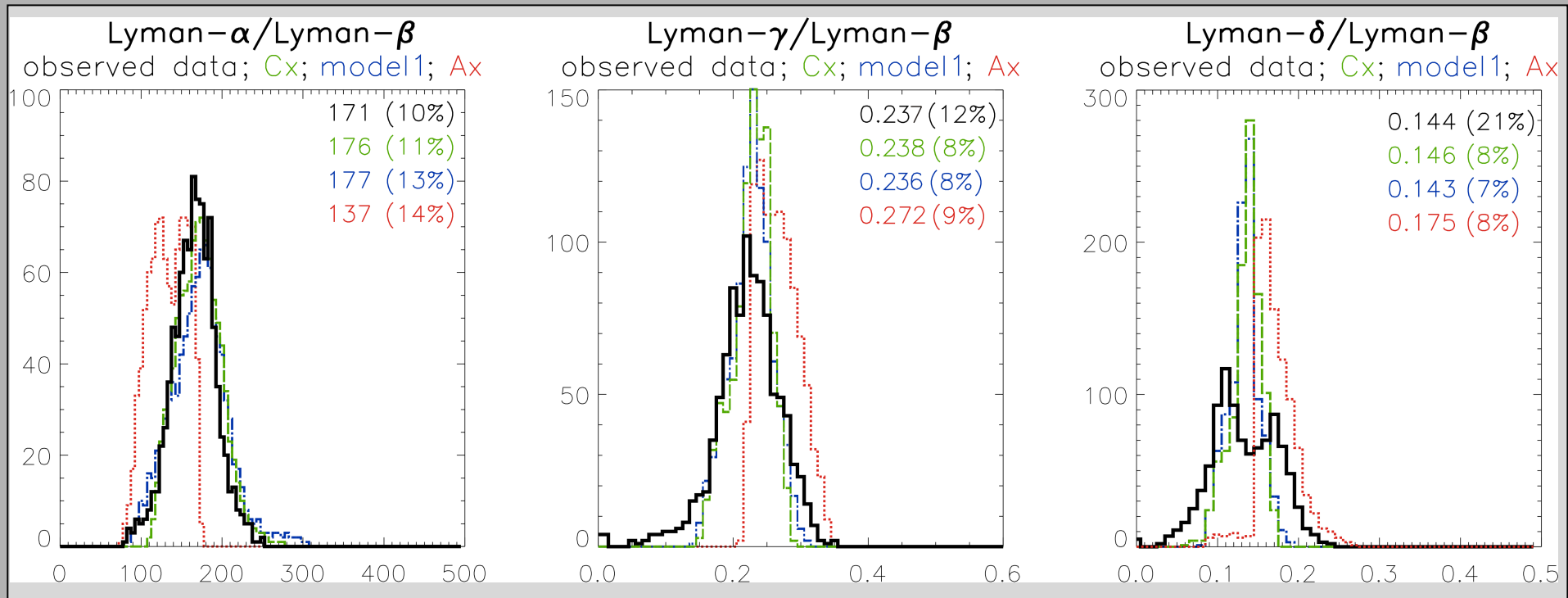
Label	Set of input parameters
model1	$T_0=7000$ K; $T_{tr}=100\ 000$ K; $B_x(0)=6$ Gauss; $M_0=1.1 \times 10^{-4}$ g cm $^{-2}$; $p_0=0.015$ dyn cm $^{-2}$; $\gamma_1=10$; $\gamma_2=60$
Ax	$T_0=7000$ K; $T_{tr}=100\ 000$ K; $B_x(0)=6$ Gauss; $M_0=1.1 \times 10^{-4}$ g cm $^{-2}$; $p_0=0.015$ dyn cm $^{-2}$; $\gamma_1=5$; $\gamma_2=30$
Cx	$T_0=7000$ K; $T_{tr}=100\ 000$ K; $B_x(0)=6$ Gauss; $M_0=1.1 \times 10^{-4}$ g cm $^{-2}$; $p_0=0.015$ dyn cm $^{-2}$; $\gamma_1=10$; $\gamma_2=30$

the LOS velocities from interval between -10 and 10 km/s

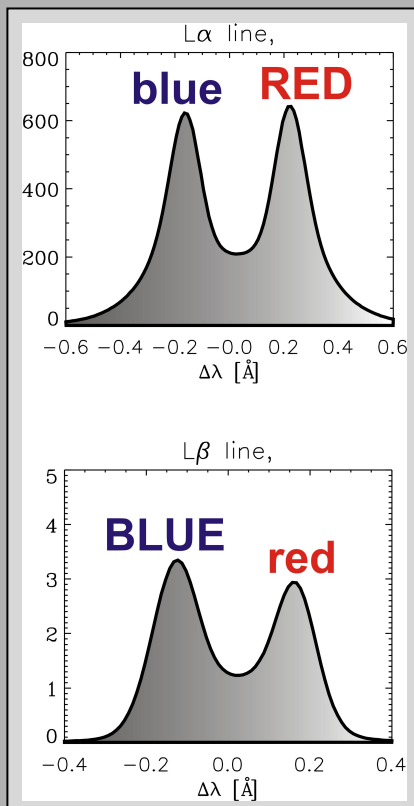
Integrated Intensities



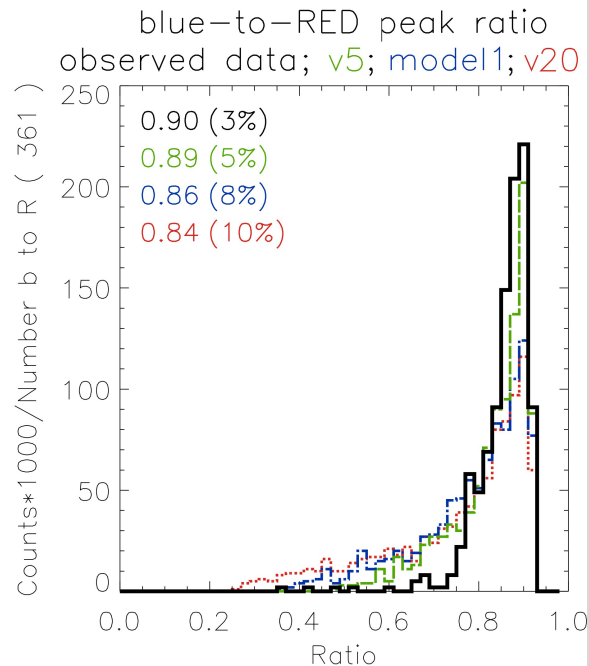
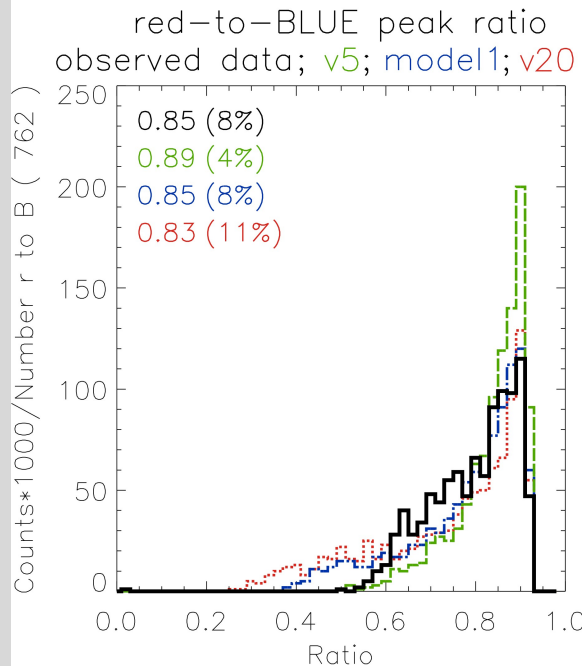
Lyman decrement ratios



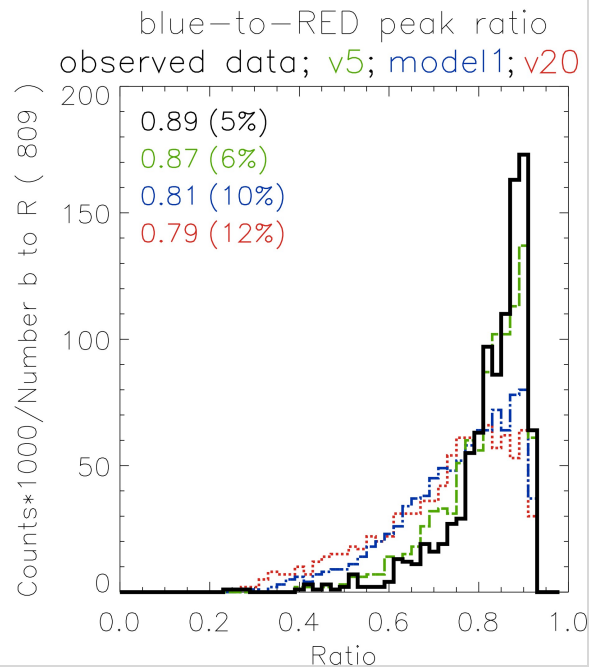
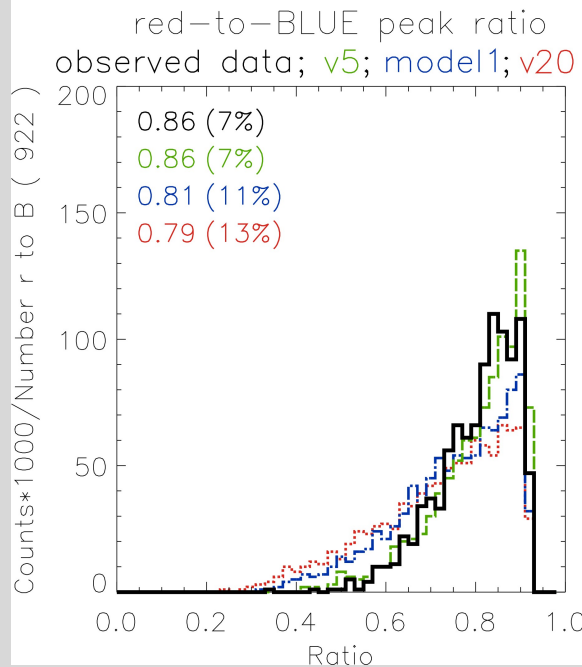
Peak asymmetries



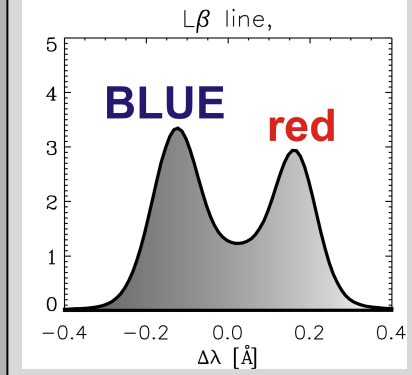
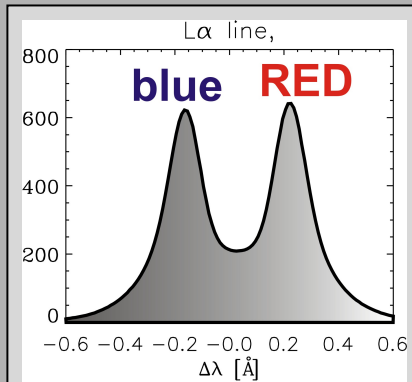
Lyman- α



Lyman- β

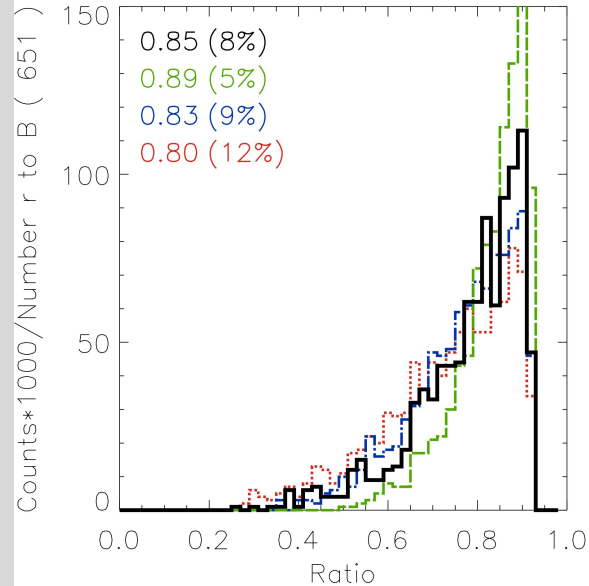


Peak asymmetries

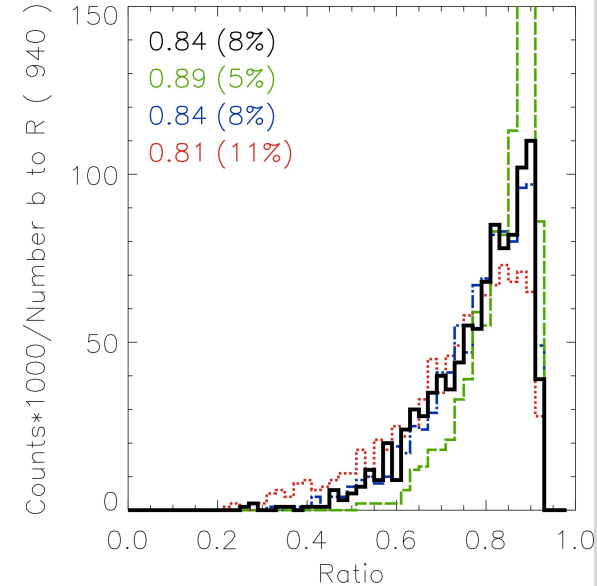


Lyman- γ

red-to-BLUE peak ratio
observed data; v5; model1; v20

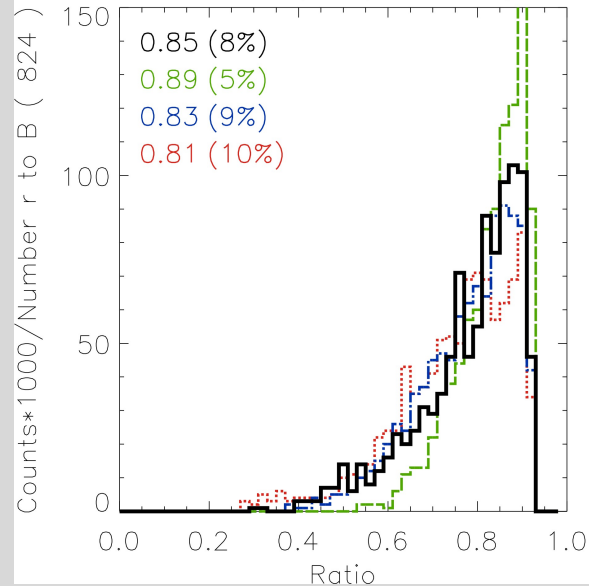


blue-to-RED peak ratio
observed data; v5; model1; v20

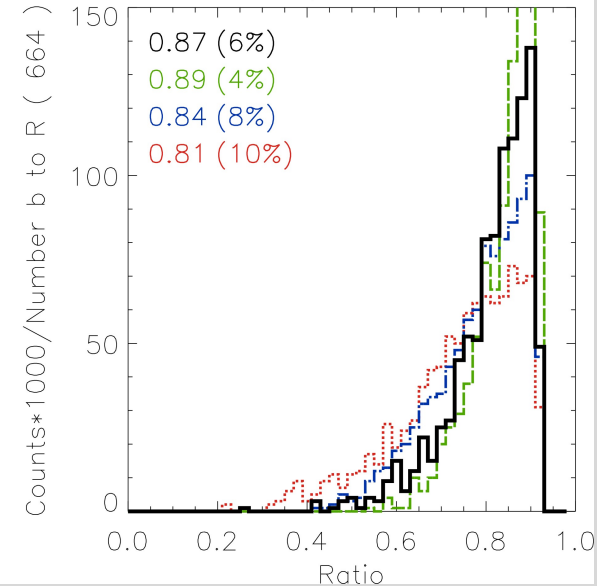


Lyman- δ

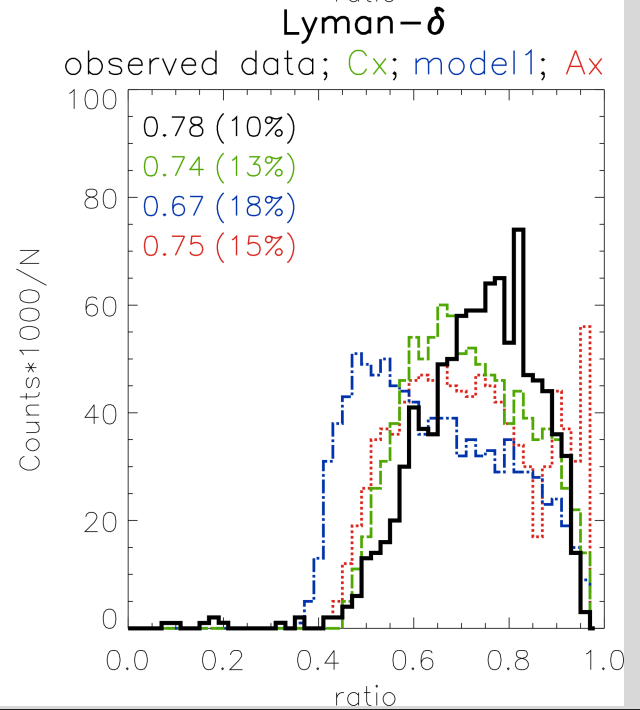
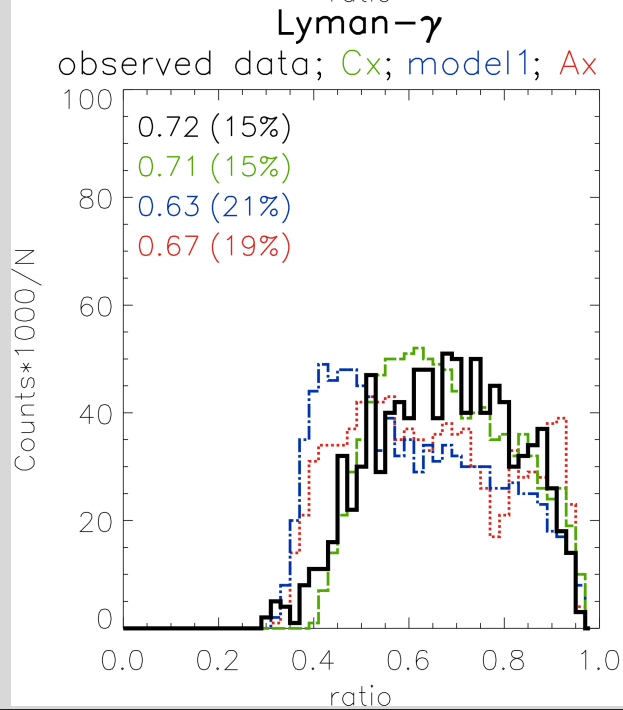
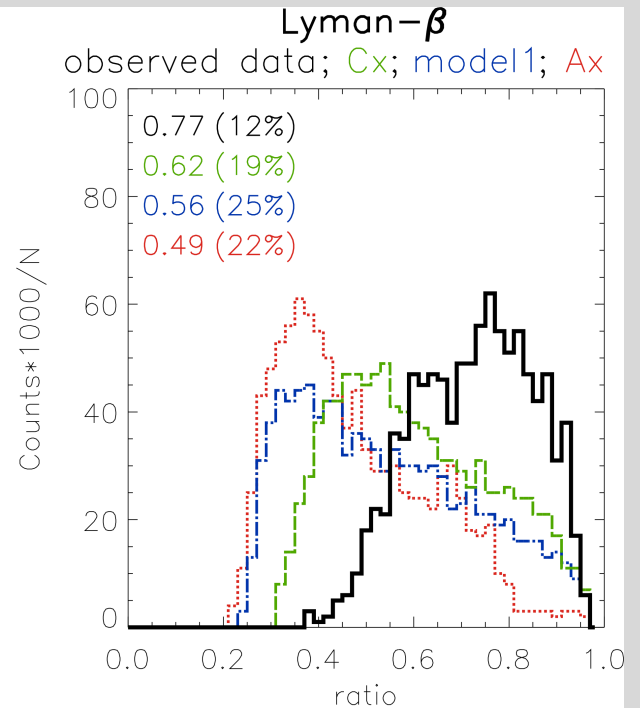
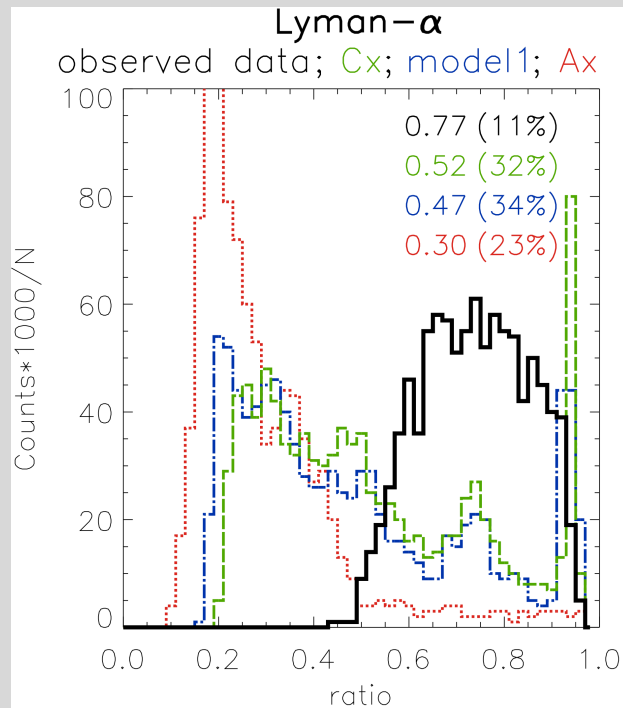
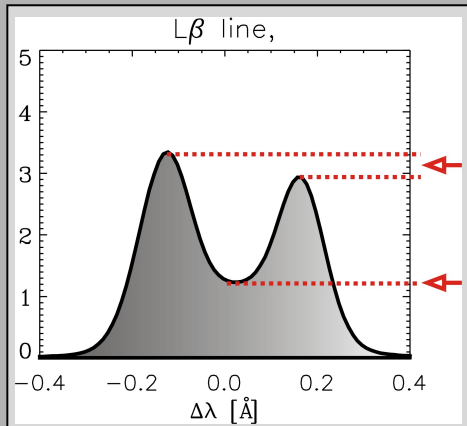
red-to-BLUE peak ratio
observed data; v5; model1; v20



blue-to-RED peak ratio
observed data; v5; model1; v20

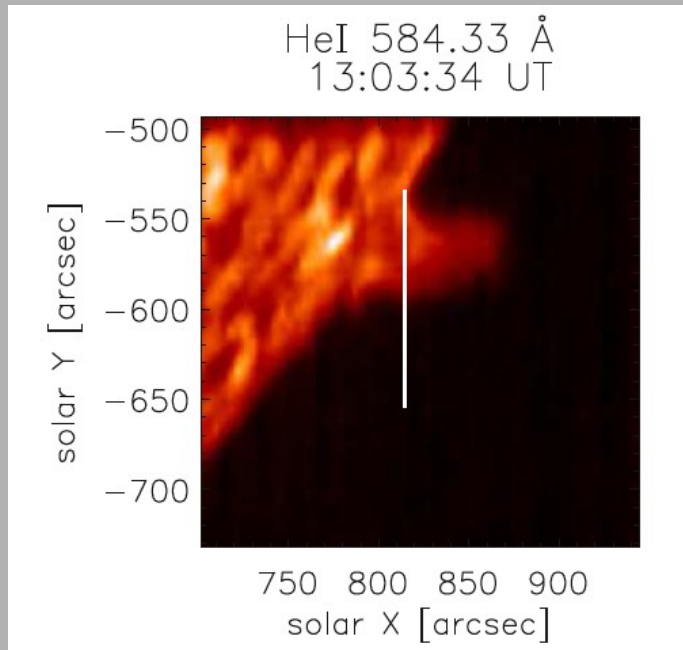


Reversal 'deepness'

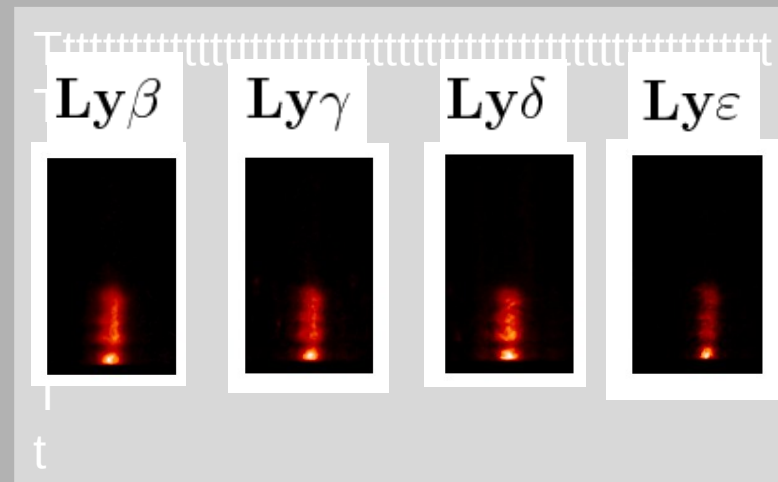


Prominence observations of April the 26, 2007

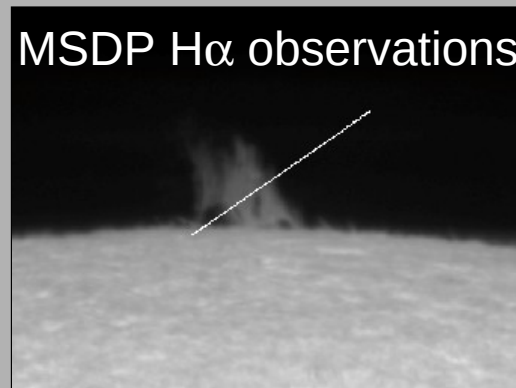
CDS observations

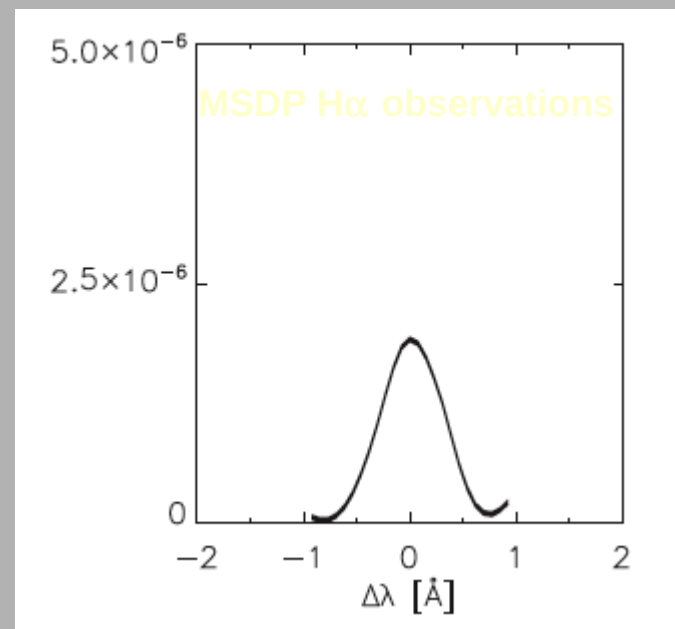
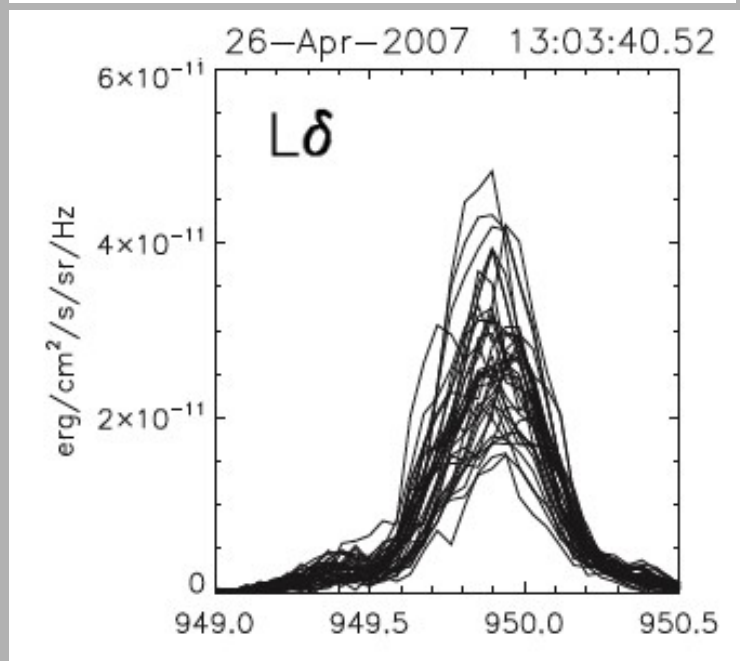
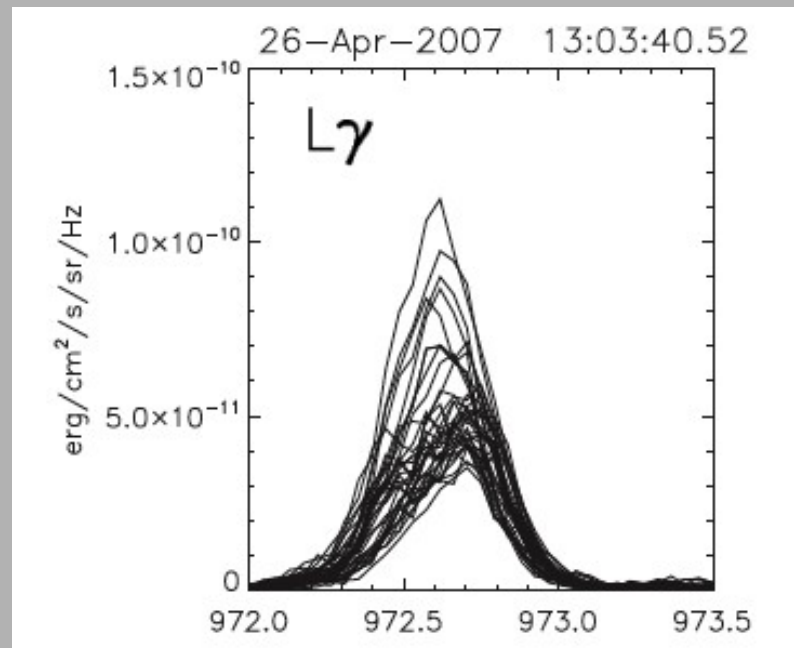
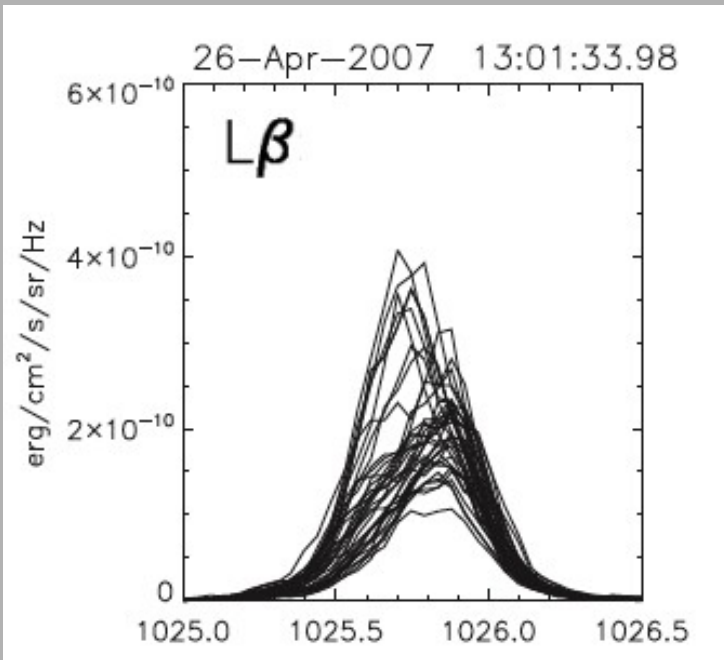


SUMER observations

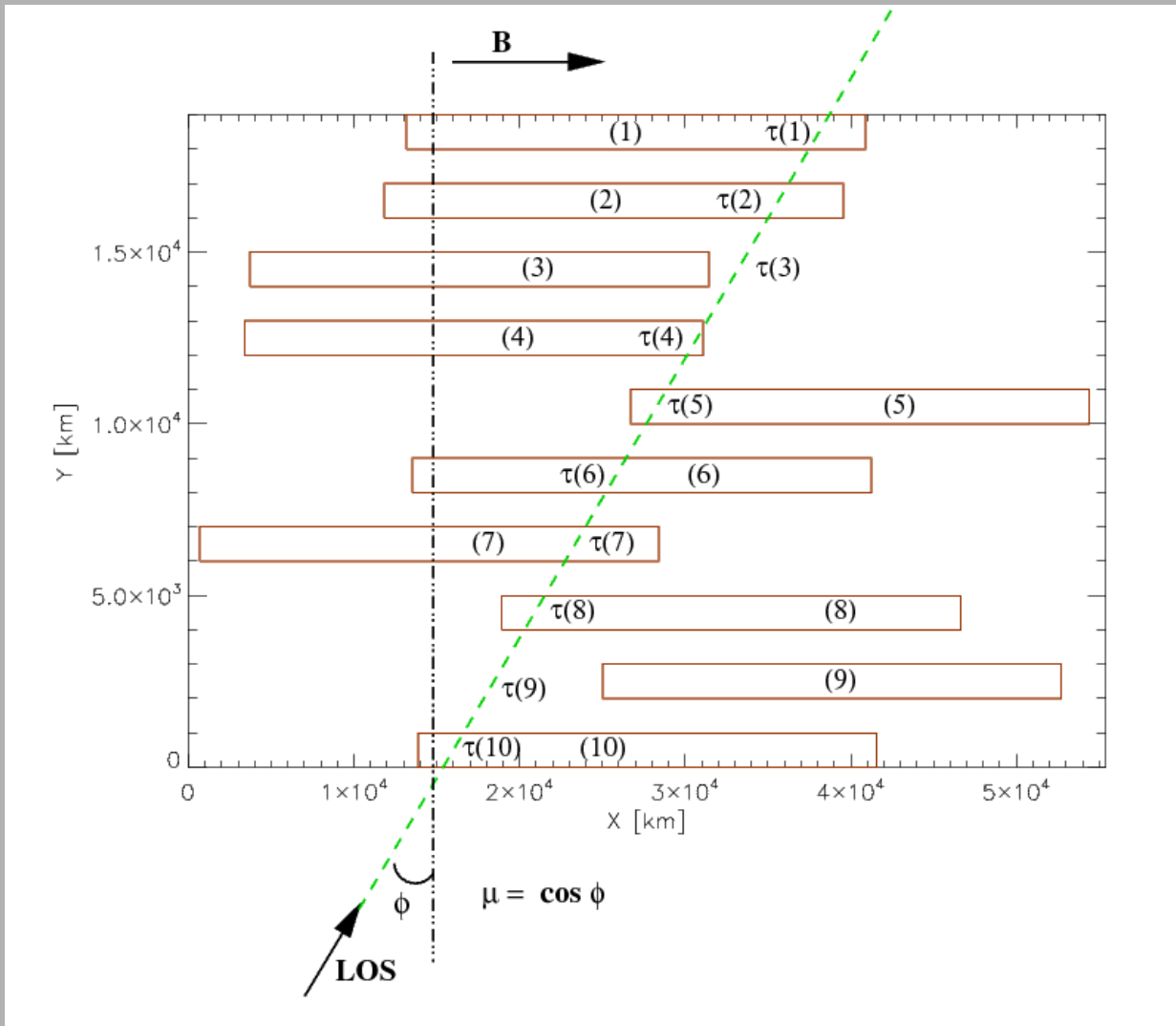


MSDP Hα observations





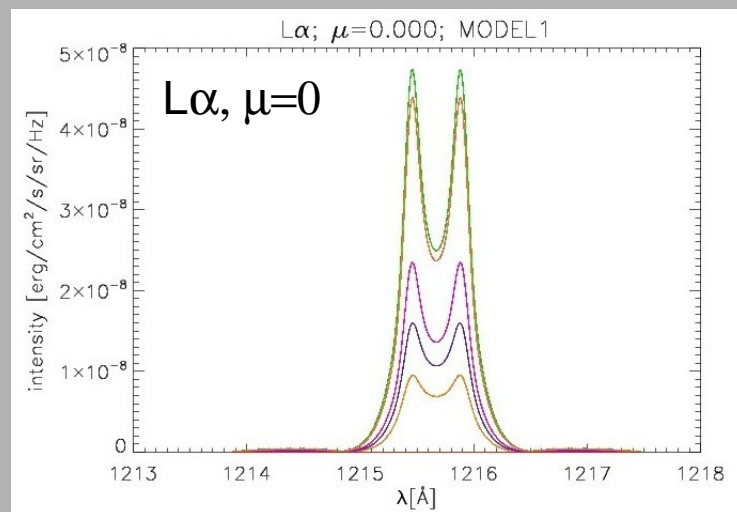
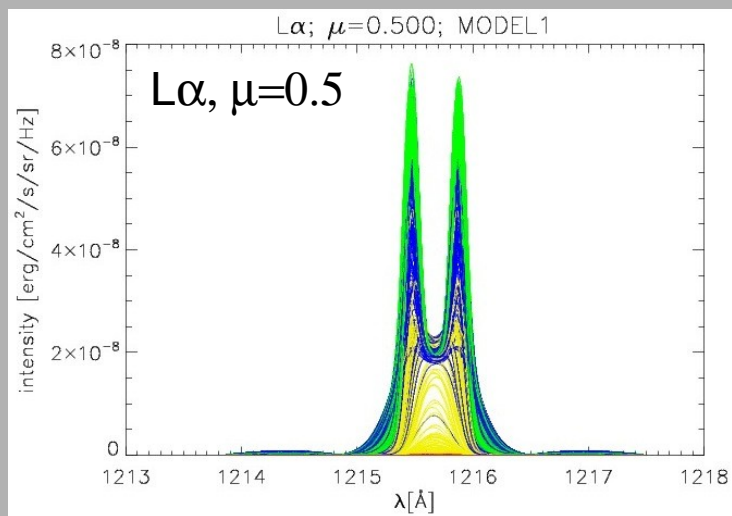
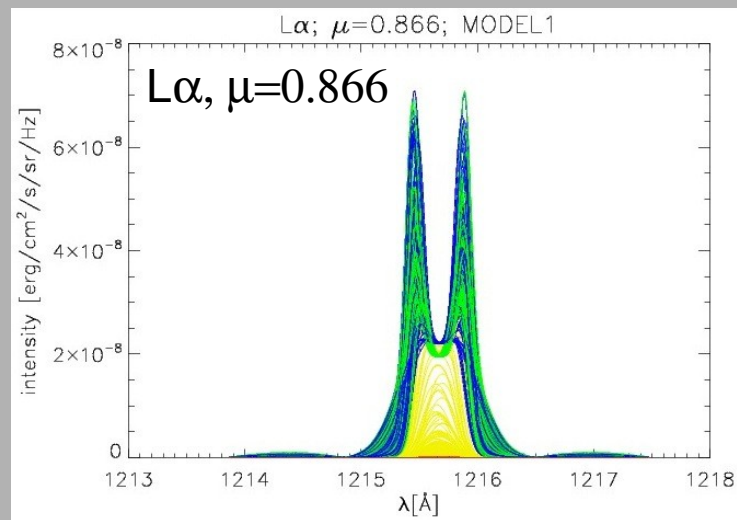
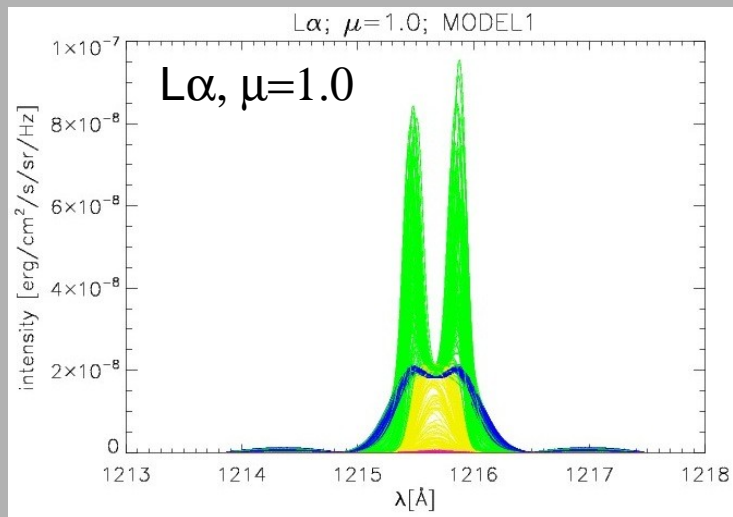
The multi-thread model at different view angles



MODEL1

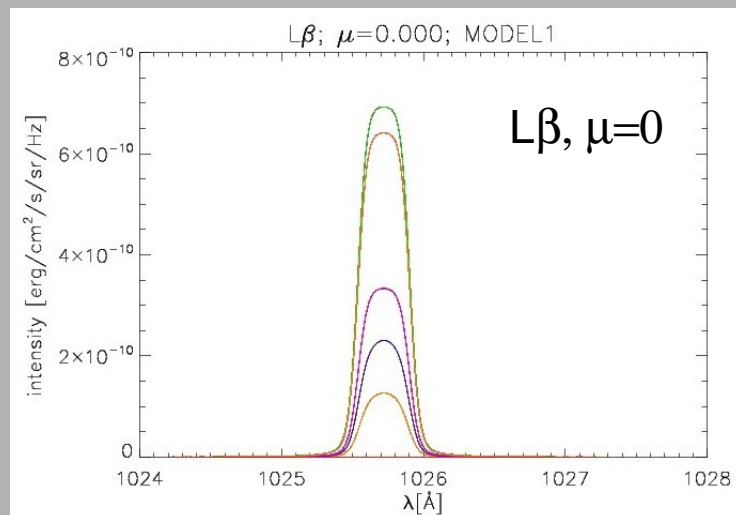
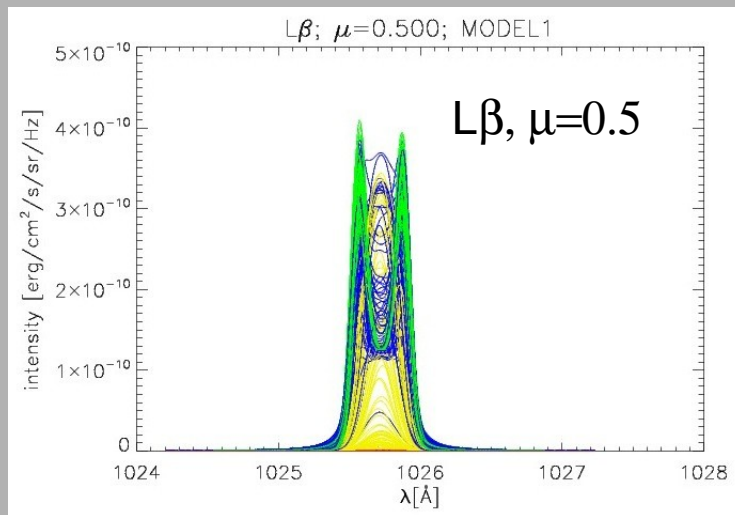
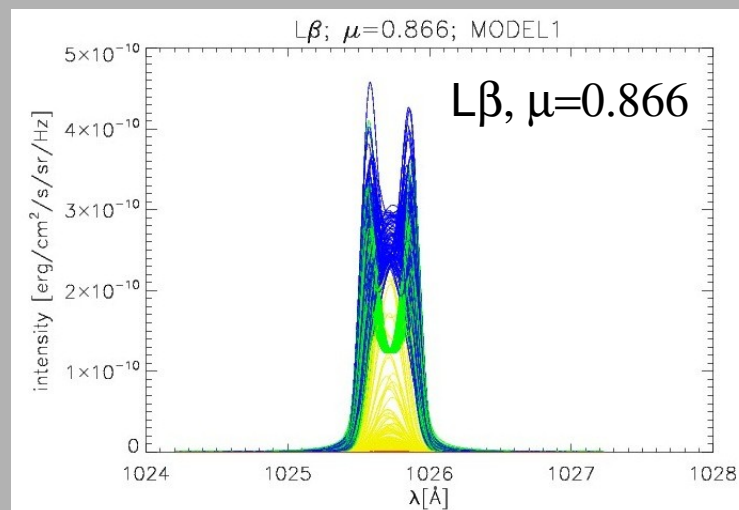
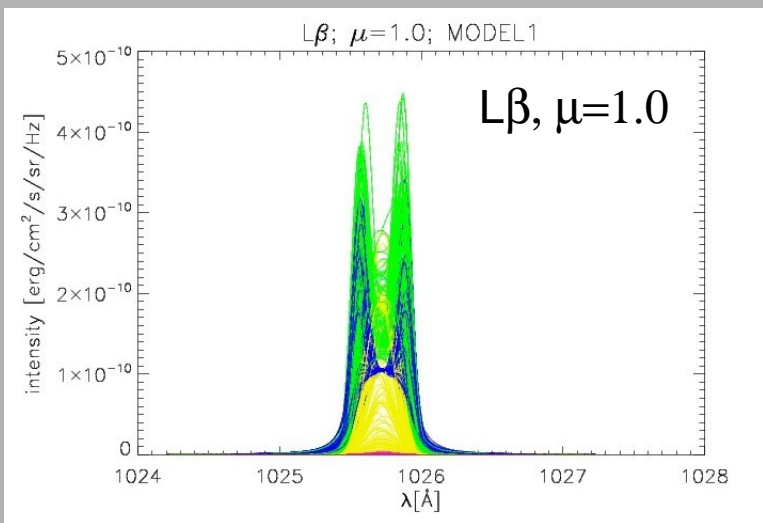
$T_0=7000$ K; $B_x(0)=6$ Gauss;
 $M_0=1.1 \times 10^{-4}$ g cm $^{-2}$; $p_0=0.015$ dyn cm $^{-2}$

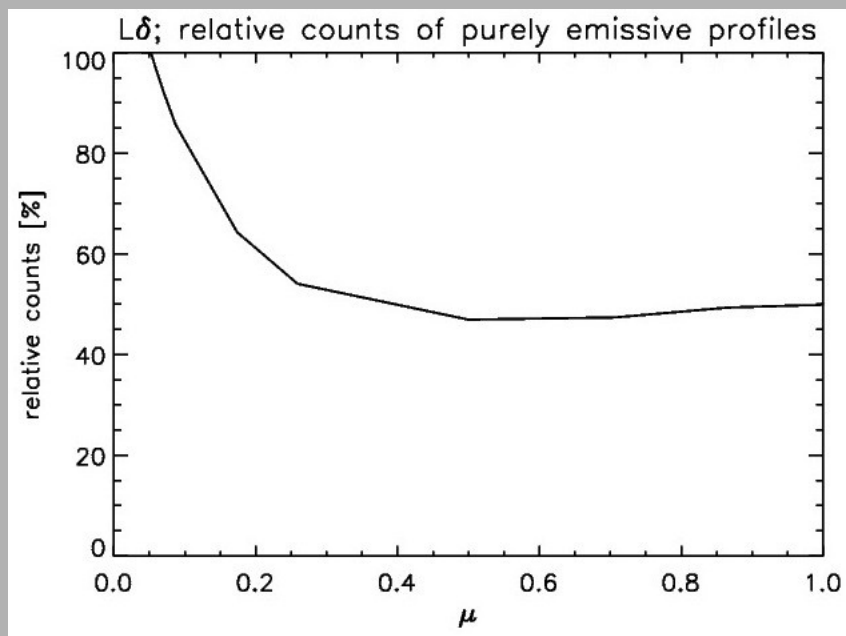
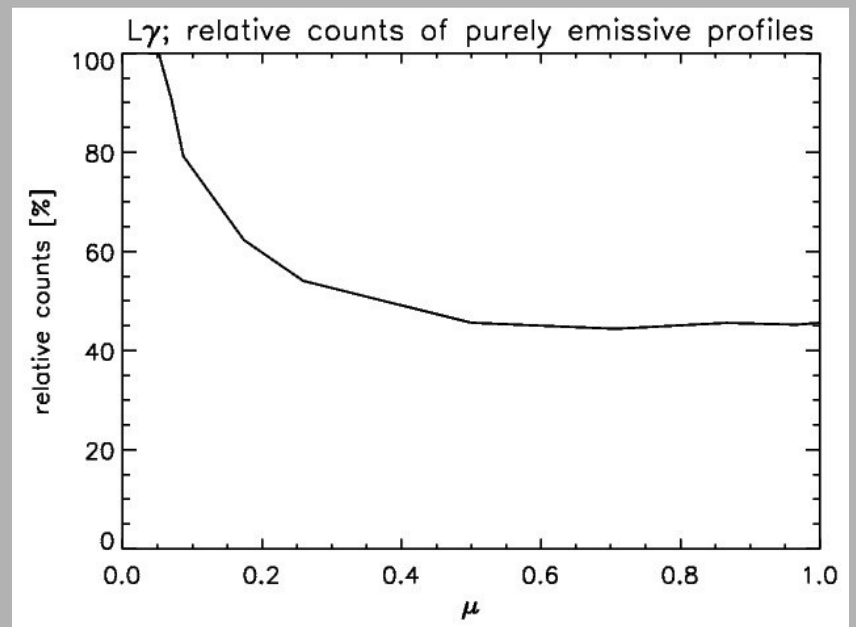
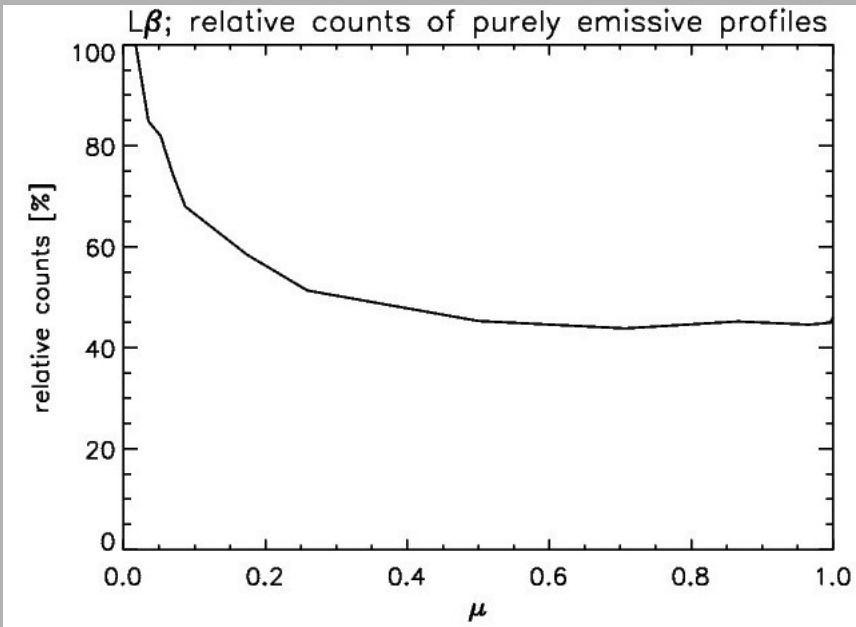
$\gamma_1=10$
 $\gamma_2=60$



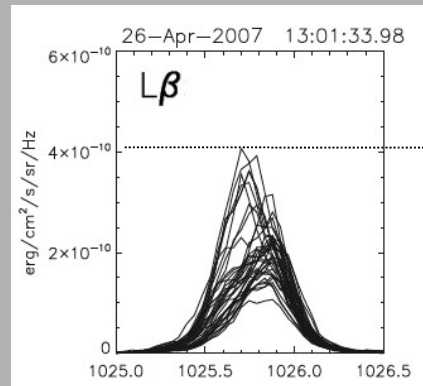
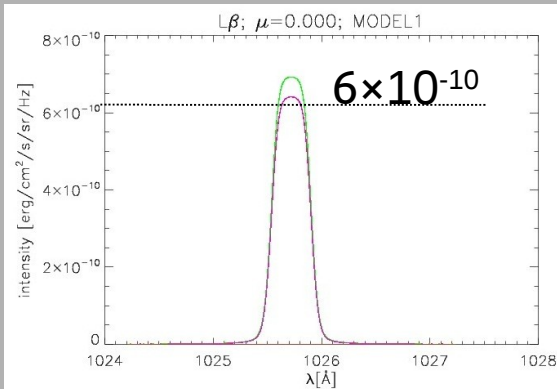
MODEL1

$T_0=7000$ K; $B_x(0)=6$ Gauss; $\gamma_1=10$
 $M_0=1.1 \times 10^{-4}$ g cm $^{-2}$; $p_0=0.015$ dyn cm $^{-2}$ $\gamma_2=60$

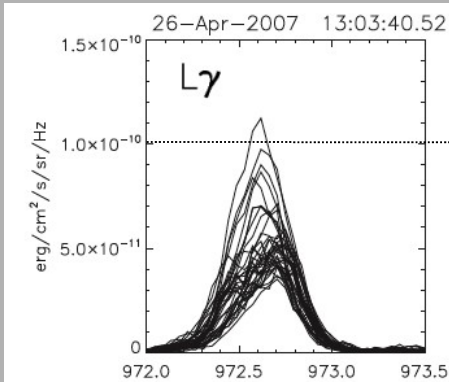
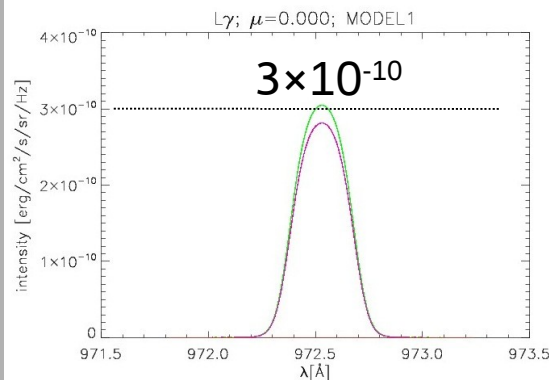




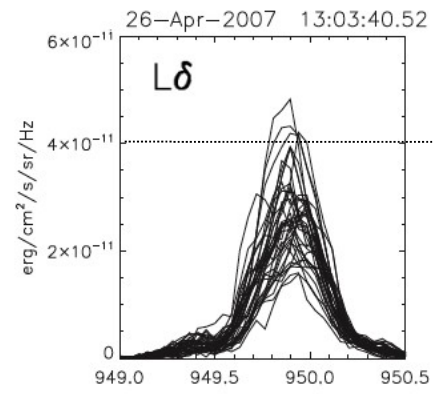
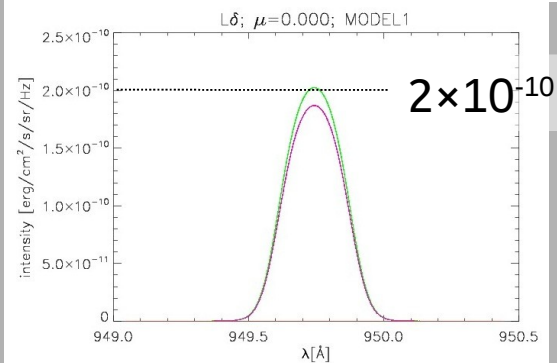
Comparison of the synthetic Lyman line profiles of MODEL1 for $\mu=0$ with the observations



1.5-times higher synthetic intensities than observed



3-times higher synthetic intensities than observed

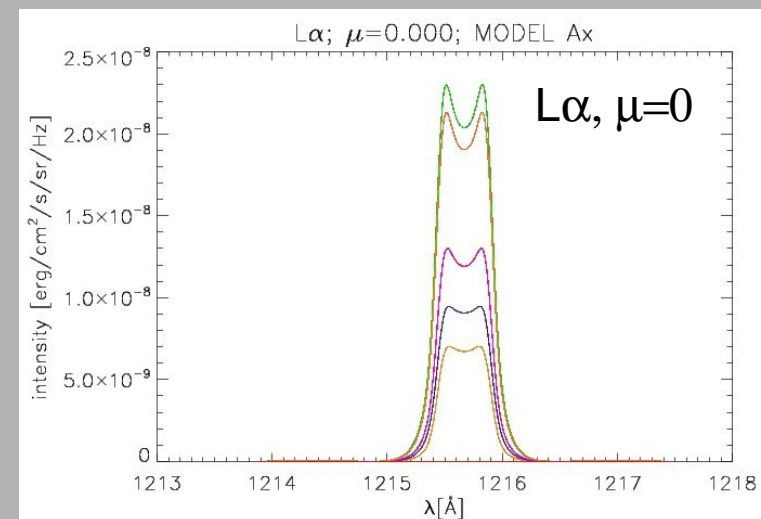
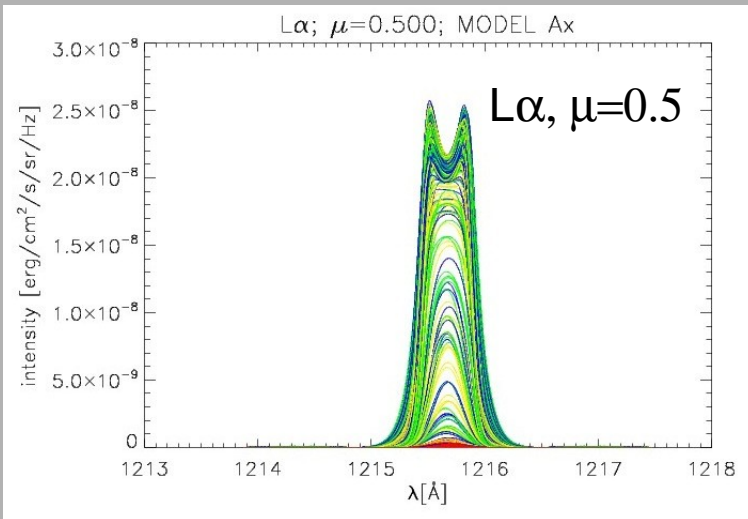
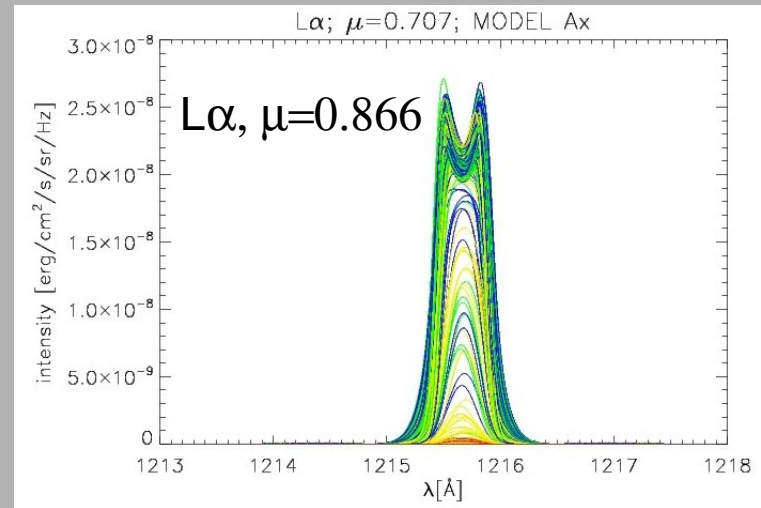
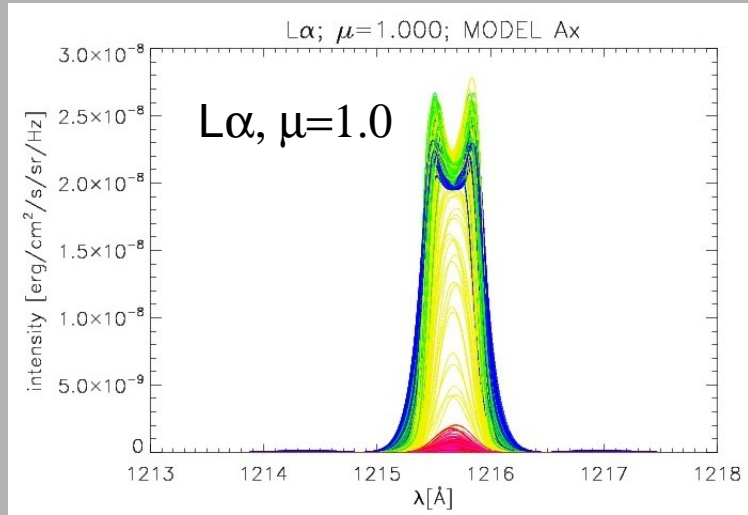


5-times higher synthetic intensities than observed

Ax1

$T_0=8000$ K; $B_x(0)=5$ Gauss;
 $M_0=1 \times 10^{-5}$ g cm $^{-2}$; $p_0=0.03$ dyn cm $^{-2}$

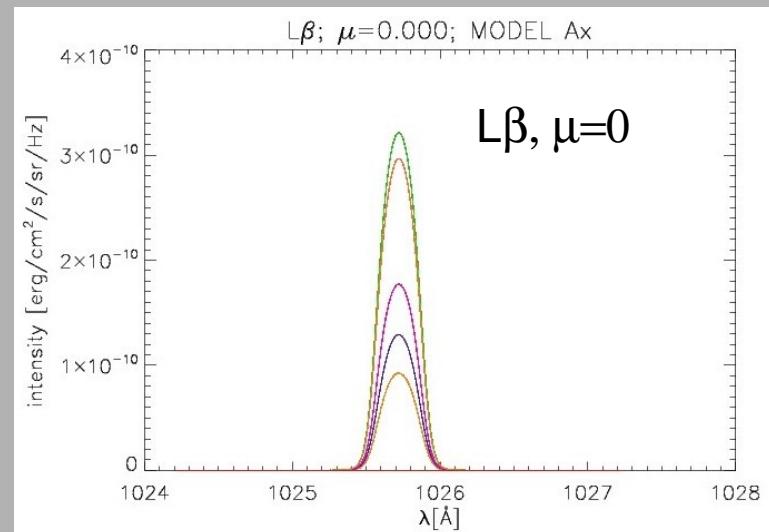
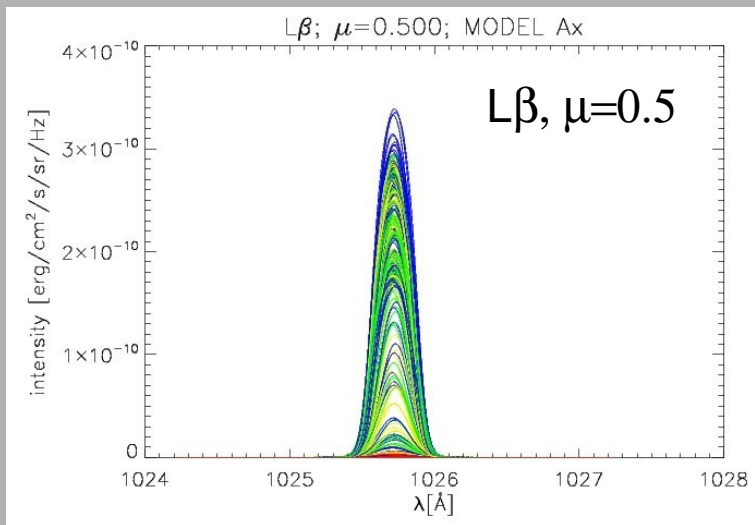
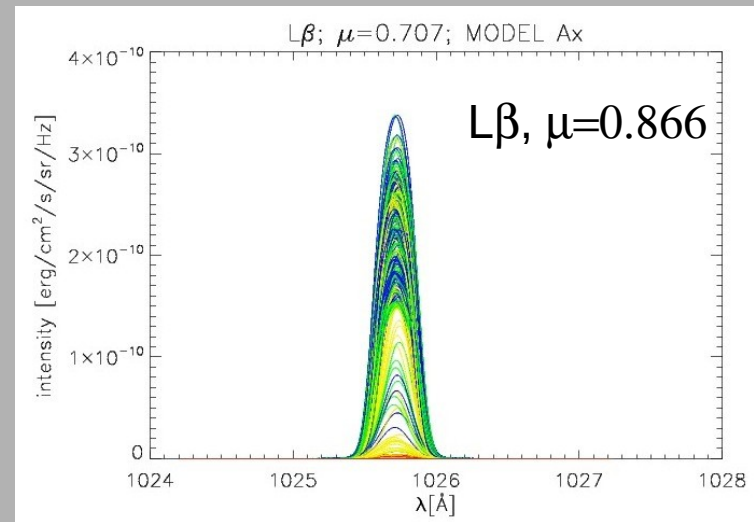
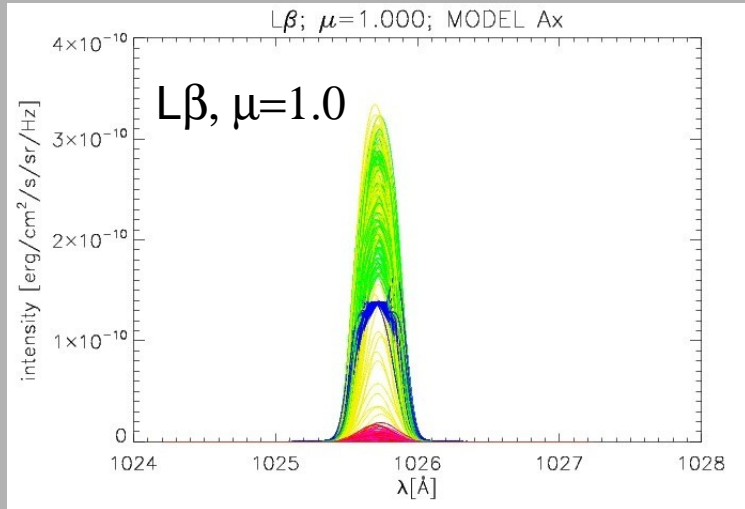
$\gamma_1=5$
 $\gamma_2=30$



Ax1

$$T_0=8000 \text{ K}; \quad B_x(0)=5 \text{ Gauss};$$
$$M_0=1 \times 10^{-5} \text{ g cm}^{-2}; \quad p_0=0.03 \text{ dyn cm}^{-2}$$

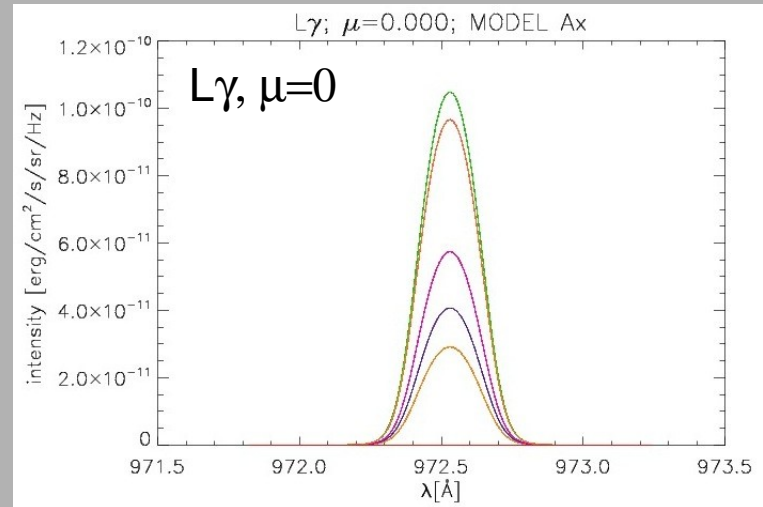
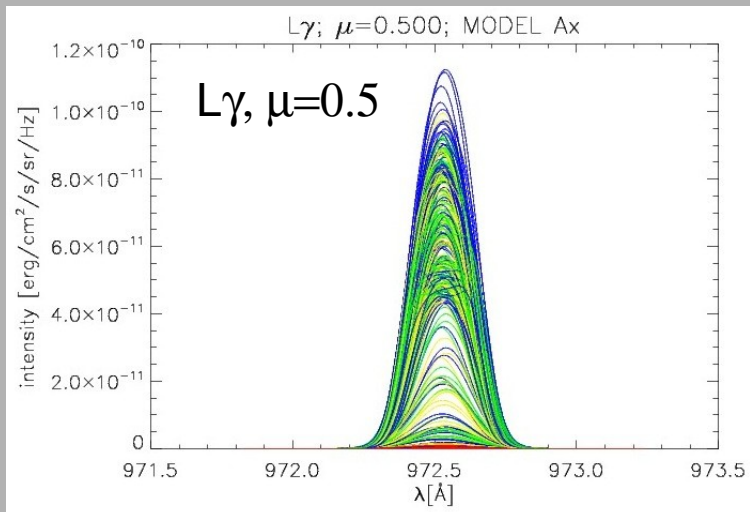
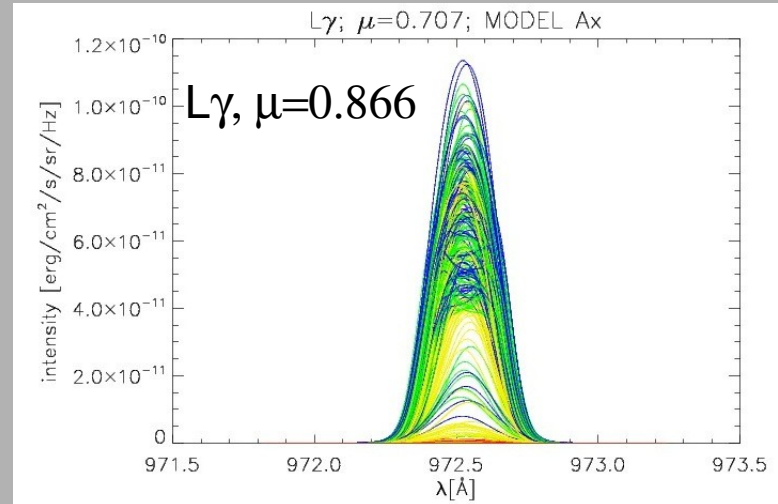
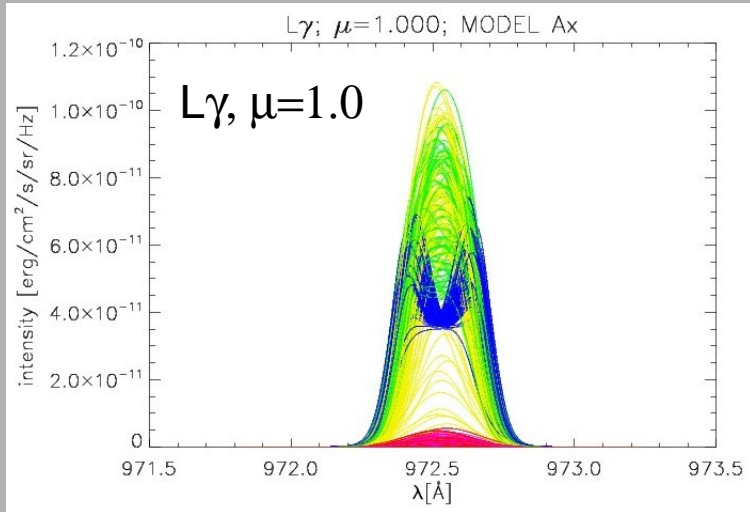
$$\gamma_1=5$$
$$\gamma_2=30$$



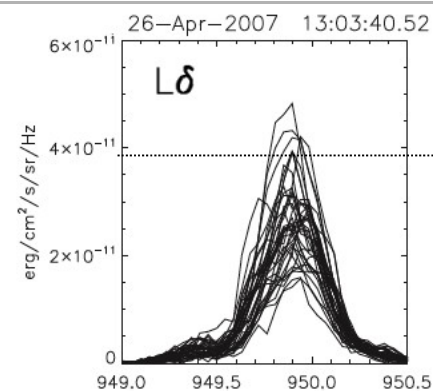
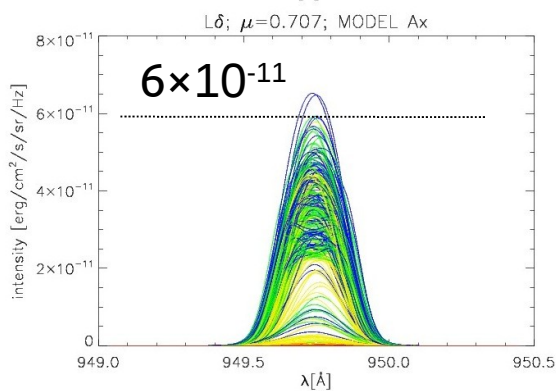
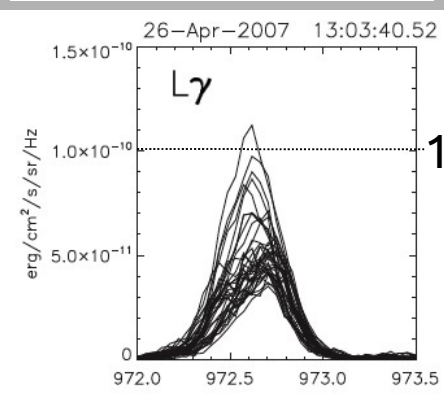
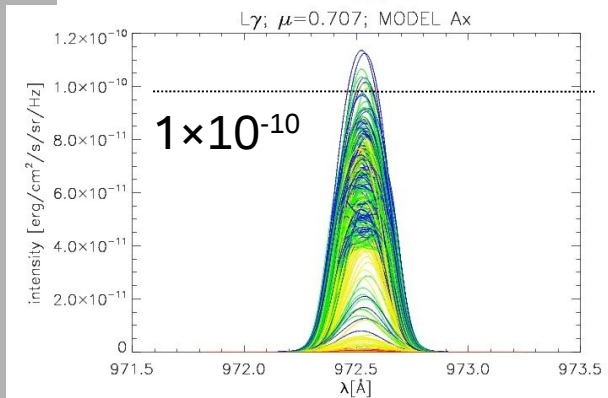
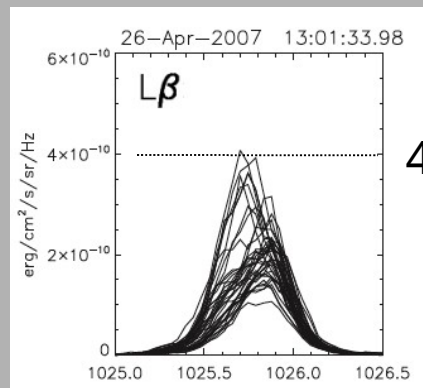
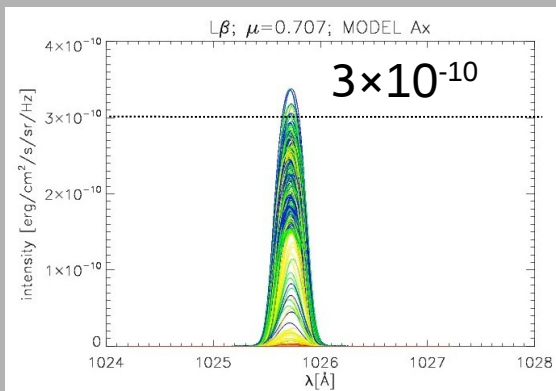
Ax1

$T_0=8000$ K; $B_x(0)=5$ Gauss;
 $M_0=1 \times 10^{-5}$ g cm $^{-2}$; $p_0=0.03$ dyn cm $^{-2}$

$\gamma_1=5$
 $\gamma_2=30$

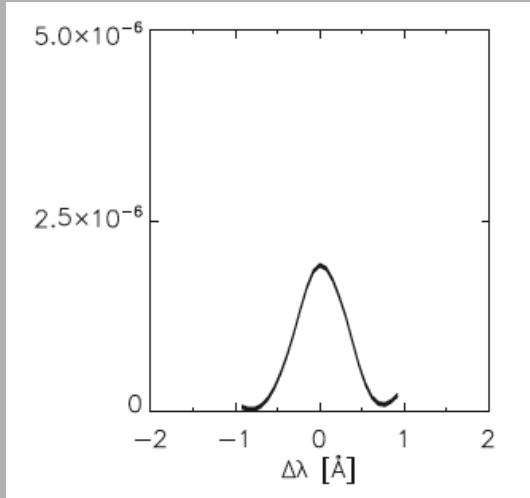


Comparison of the synthetic Lyman line profiles of MODEL Ax1 for $\mu=0.5$ with the observations

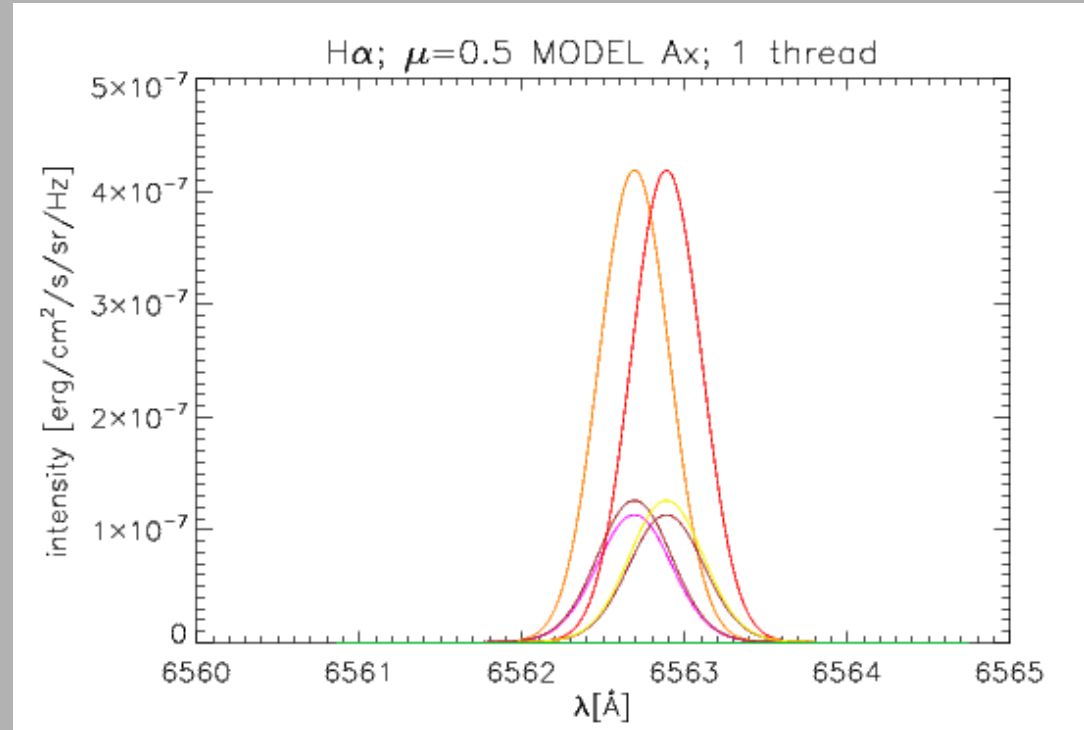


Comparison of the synthetic H α profiles of MODEL Ax1 for $\mu=0.5$ with the observations

MSDP H α observations

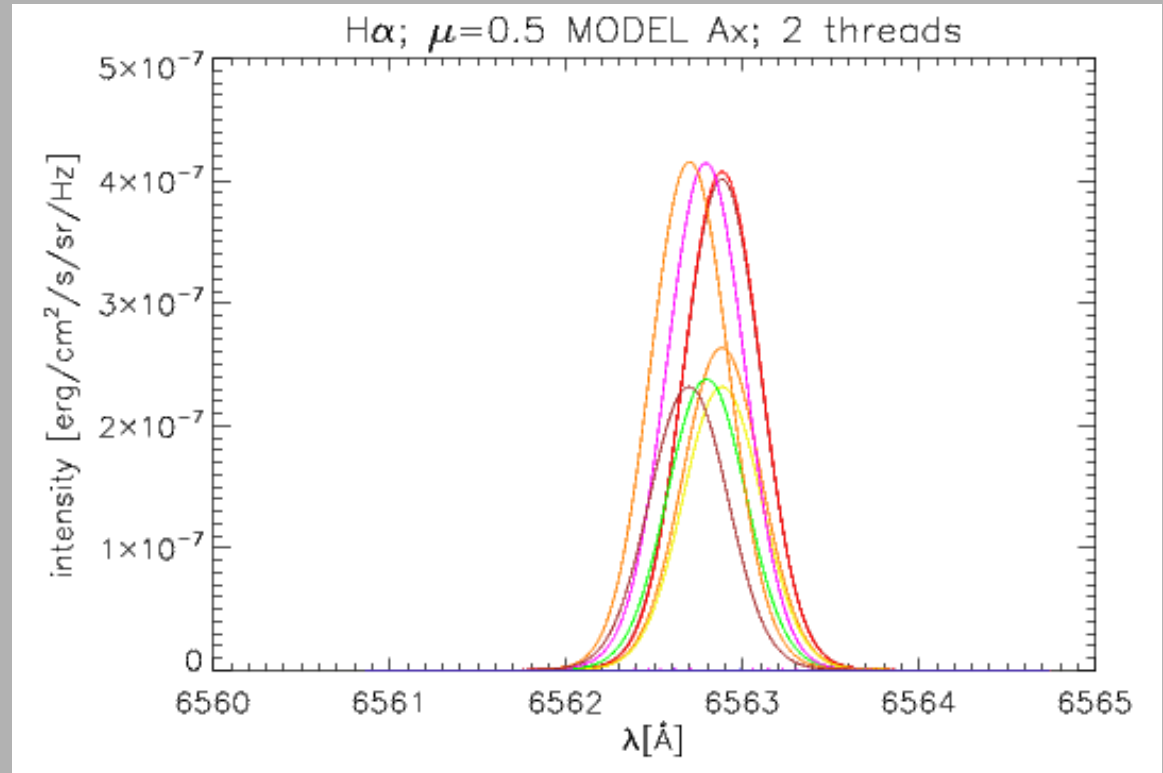
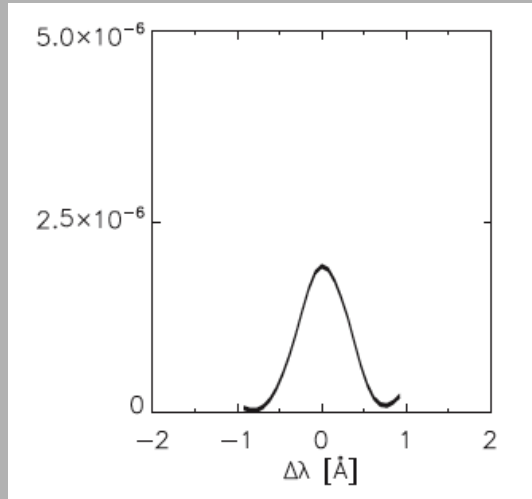


calculated profiles



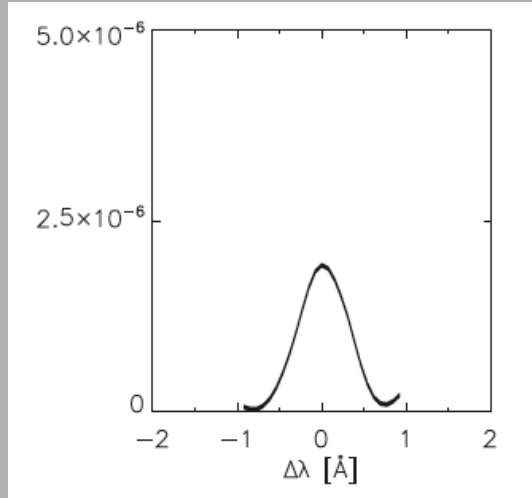
Comparison of the synthetic H α profiles of MODEL Ax1 for $\mu=0.5$ with the observations

MSDP H α observations

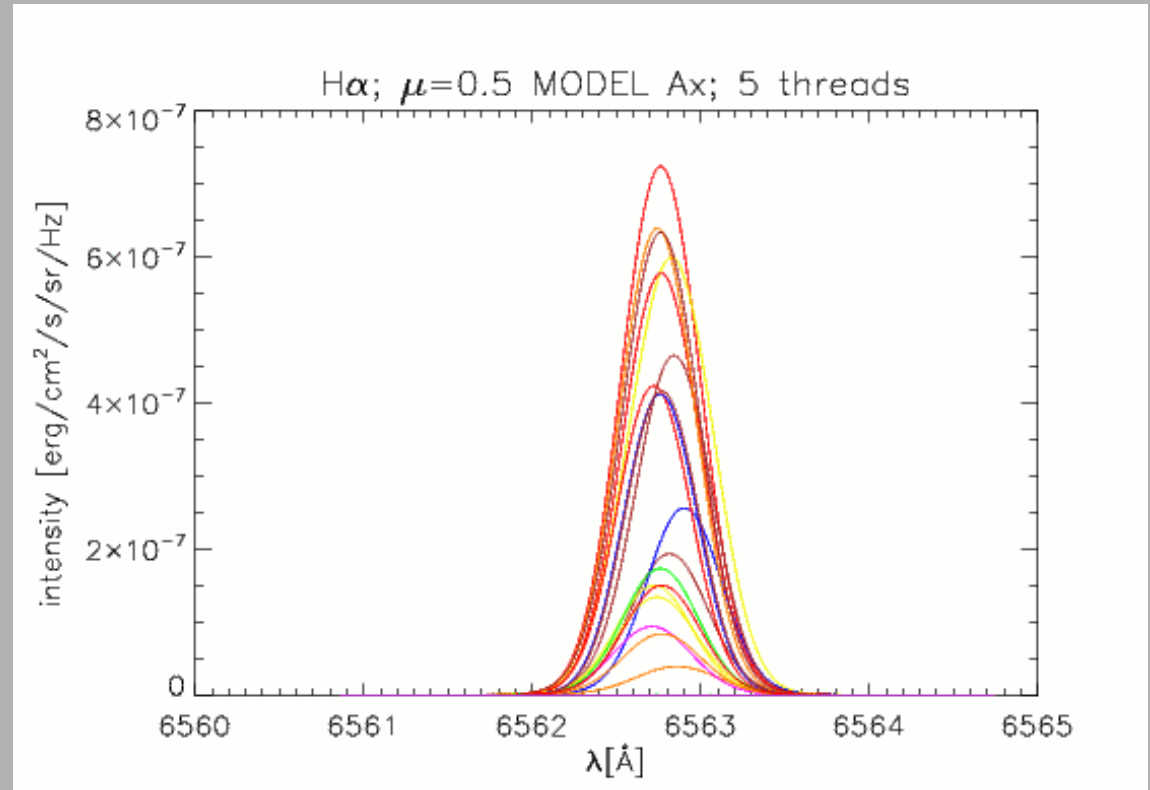


Comparison of the synthetic H α profiles of MODEL Ax1 for $\mu=0.5$ with the observations

MSDP H α observations

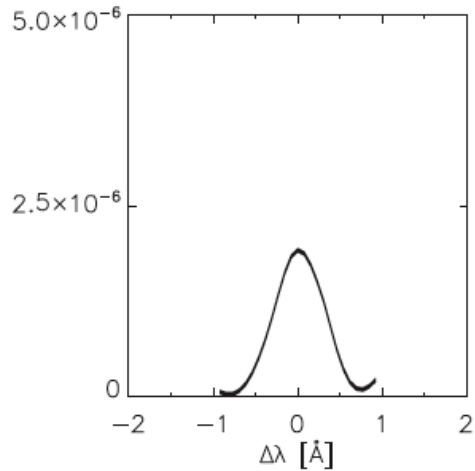


calculated profiles

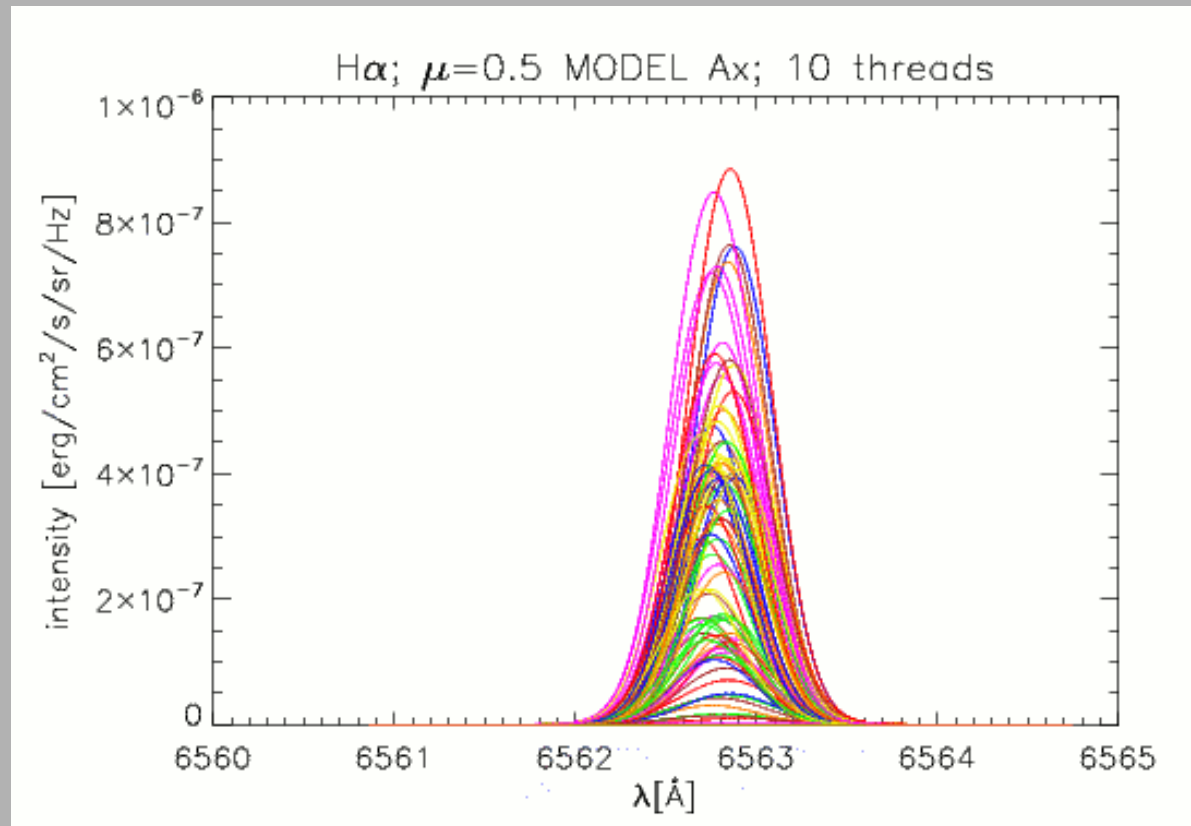


Comparison of the synthetic H α profiles of MODEL Ax1 for $\mu=0.5$ with the observations

MSDP H α observations

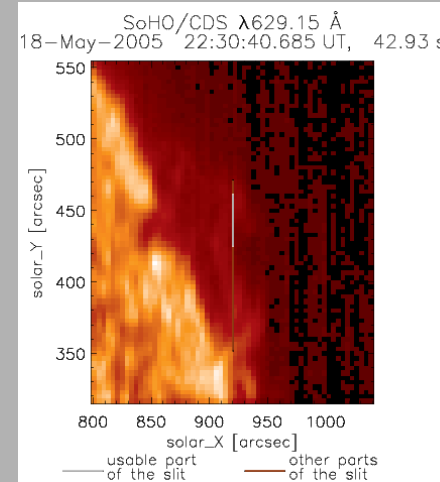
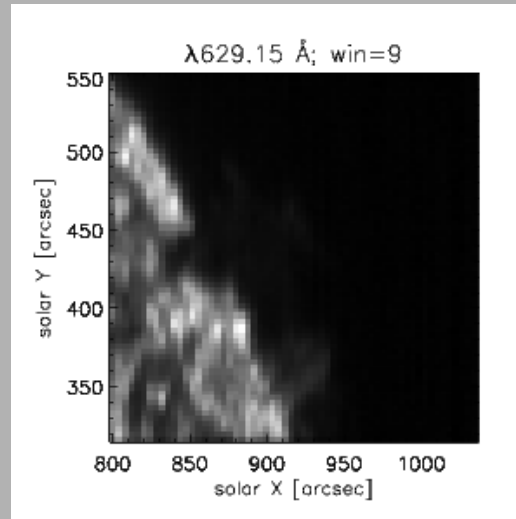


calculated profiles

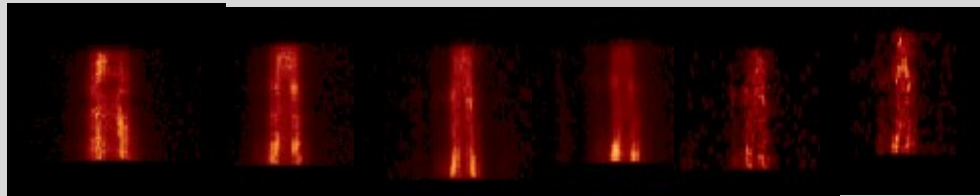


Prominence observations of May the 18, 2005

CDS
observations



$\text{Ly}\alpha$ $\text{Ly}\beta$ $\text{Ly}\gamma$ $\text{Ly}\delta$ $\text{Ly}5$ $\text{Ly}6$



SUMER
observations

model name: Model1_8000_Mo8.5E-5_Bxo7_po2E-2, one thread

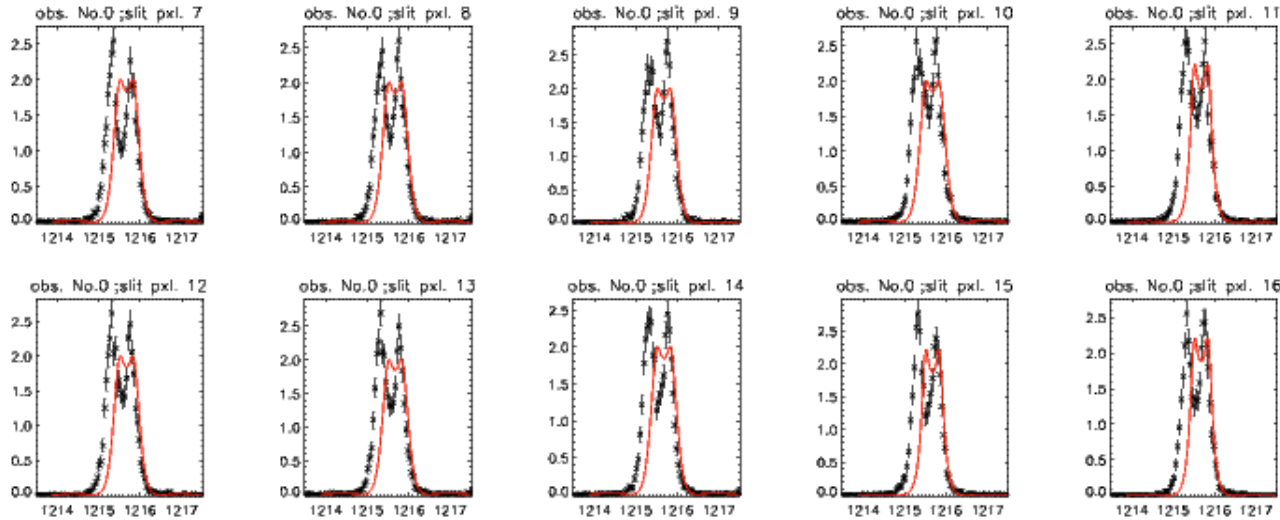
$\mu=1.0000000$; line: Ly α

λ in Å on abscissas

intensities in 10^{-8} erg/cm²/s/sr/Hz on ordinates

observations: 18-May-2005 22:06:58.147 + 19-May-2005 01:32:46.811 UT

total $\chi^2=7.6286E+05$; 25.58% of all synth. profiles assigned



searching for model according 3 criterion:

- visual comparison
- the smallest χ^2 between observed and assigned synth. profiles
- percentage of assigned synthetic profiles from all calculated

Mo9.5E-5_9000_Bxo6_po2.5E-2

$T_0=9000$ K; $B_x(0)=6$ Gauss; $\gamma_1=10$

$M_0=9.5 \times 10^{-5}$ g cm⁻²; $p_0=0.025$ dyn cm⁻²; $\gamma_2=60$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2

$T_0=8000$ K; $B_x(0)=7$ Gauss; $\gamma_1=10$

$M_0=8.5 \times 10^{-5}$ g cm⁻²; $p_0=0.02$ dyn cm⁻²; $\gamma_2=60$

Model_A4_Mo4E-6_po0.03

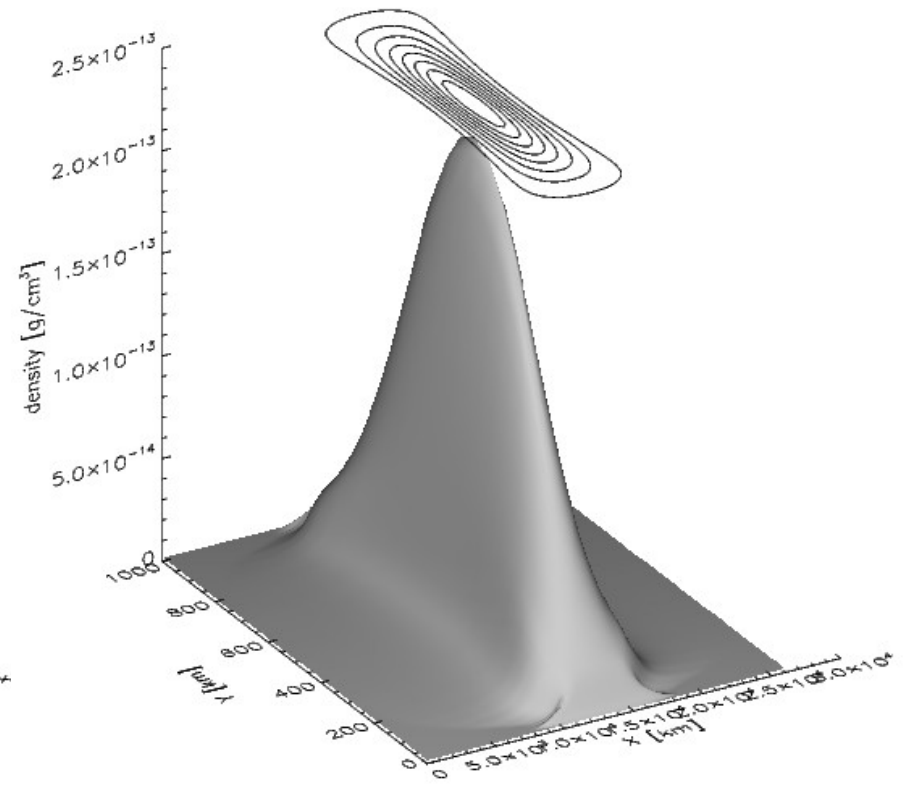
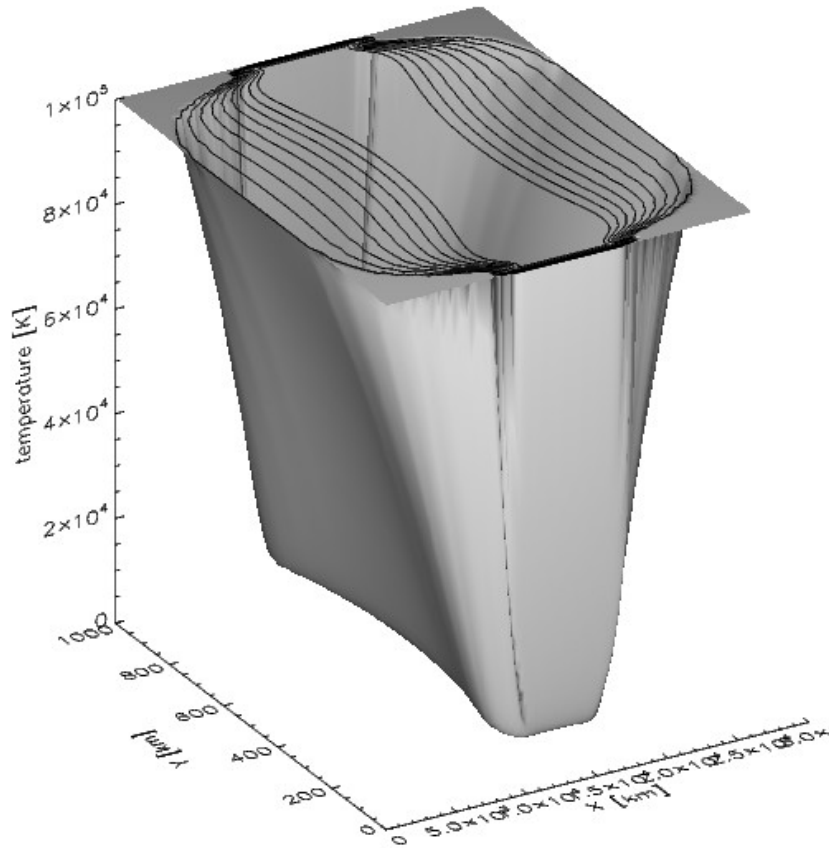
$T_0=8000$ K; $B_x(0)=1$ Gauss; $\gamma_1=5$

$M_0=4.0 \times 10^{-6}$ g cm⁻²; $p_0=0.03$ dyn cm⁻²; $\gamma_2=30$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2

temperature

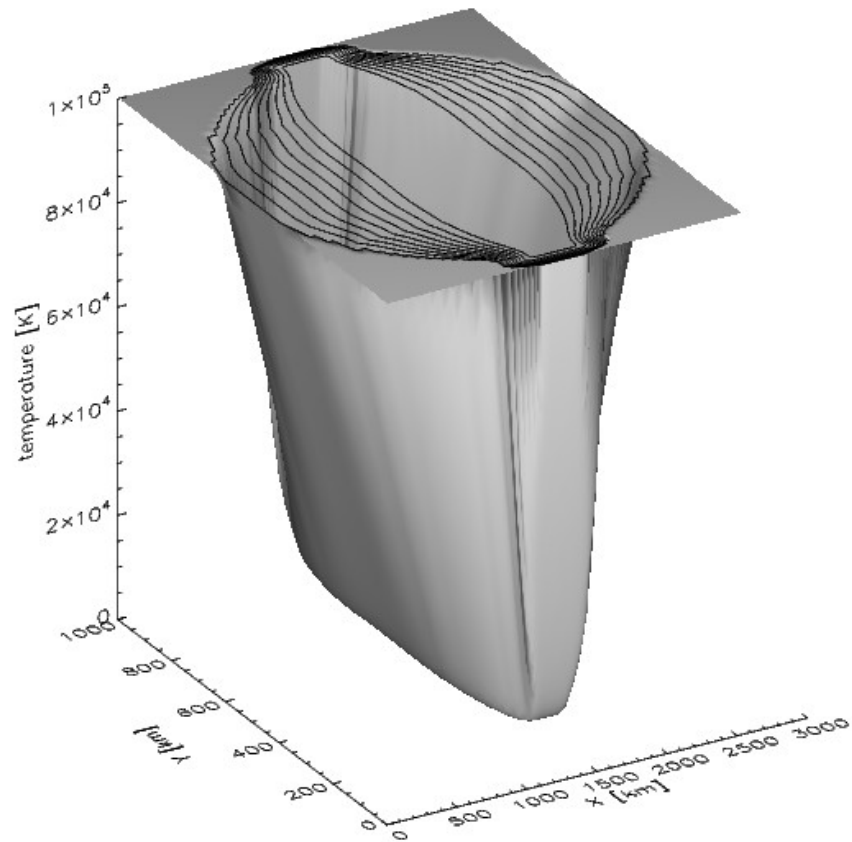
plasma density



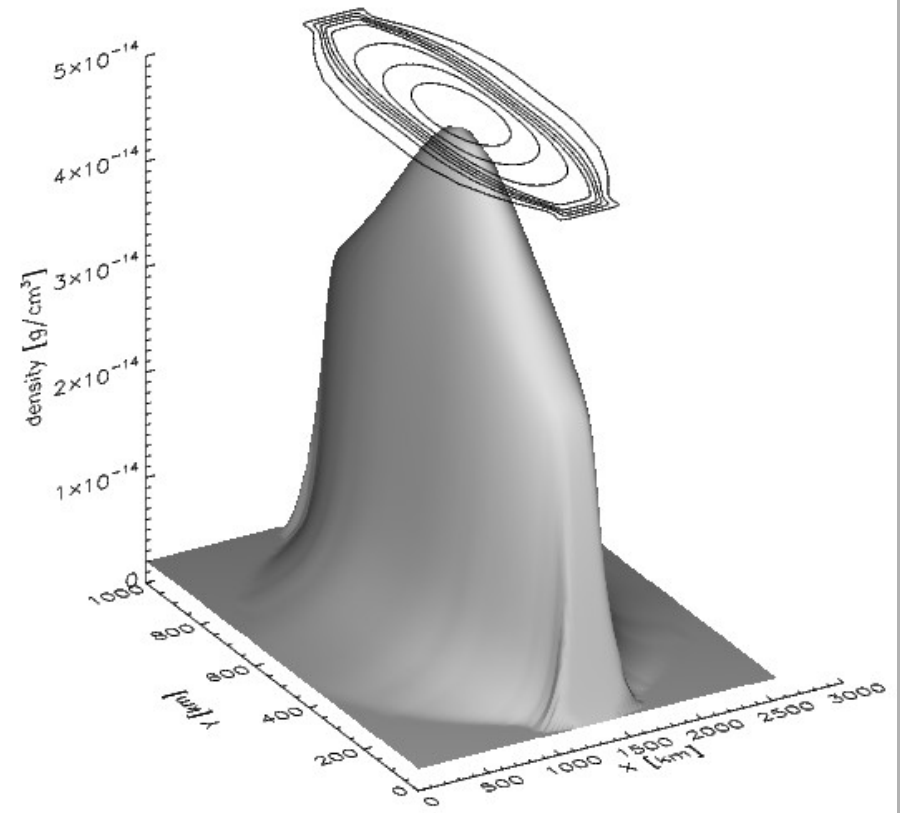
similar plasma properties also for Mo9.5E-5_9000_Bxo6_po2.5E-2

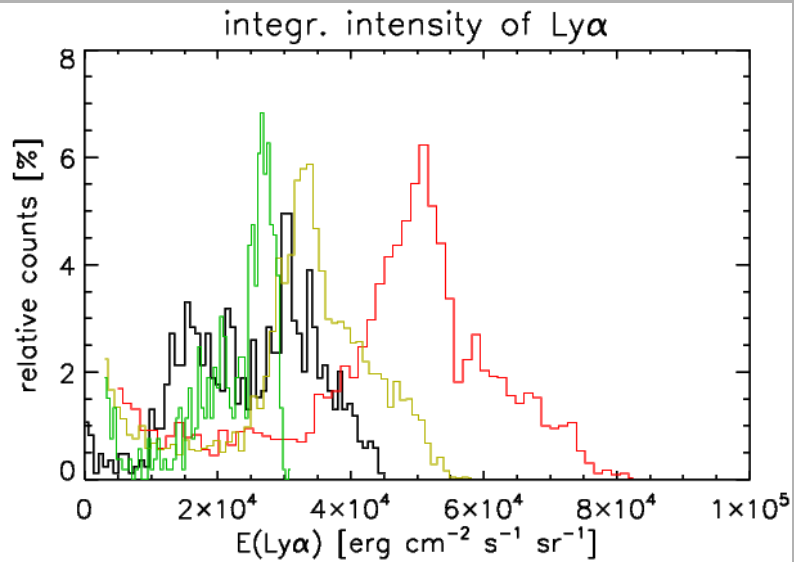
Model_A4_Mo4E-6_po0.03

temperature



plasma density



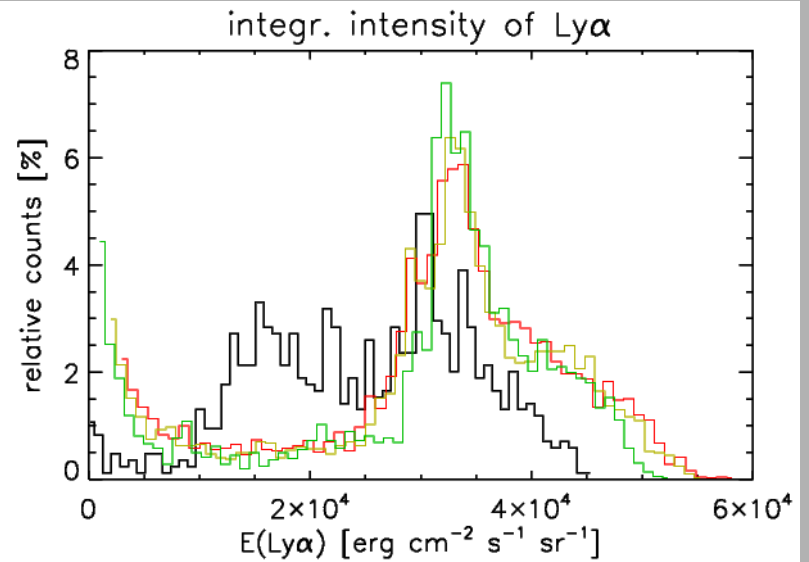


observations

Mo9.5E-5_9000_Bxo6_po2.5E-2, 10 threads
 $\mu=1.0000000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=1.0000000$

Model_A4_Mo4E-6_po0.03, 10 threads
 $\mu=1.0000000$

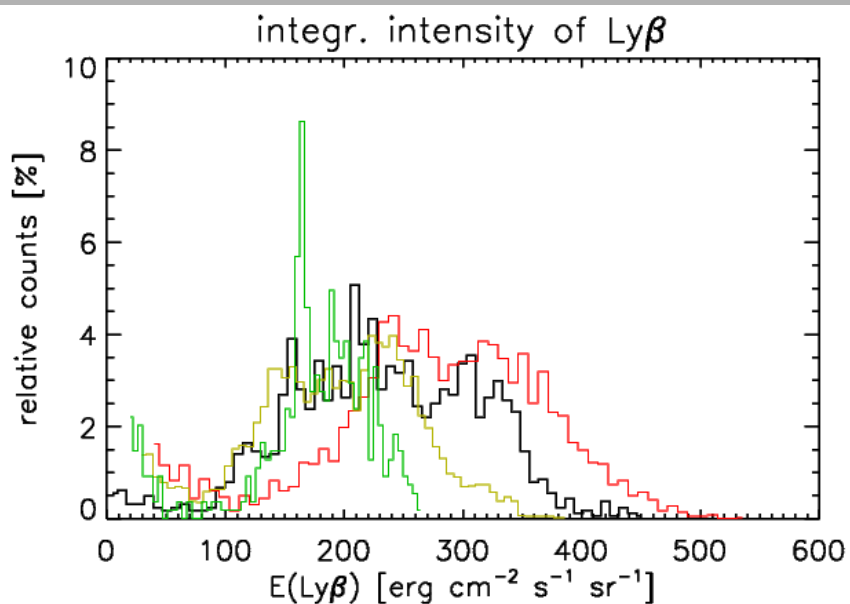


observations

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=1.0000000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=0.86600000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=0.50000000$

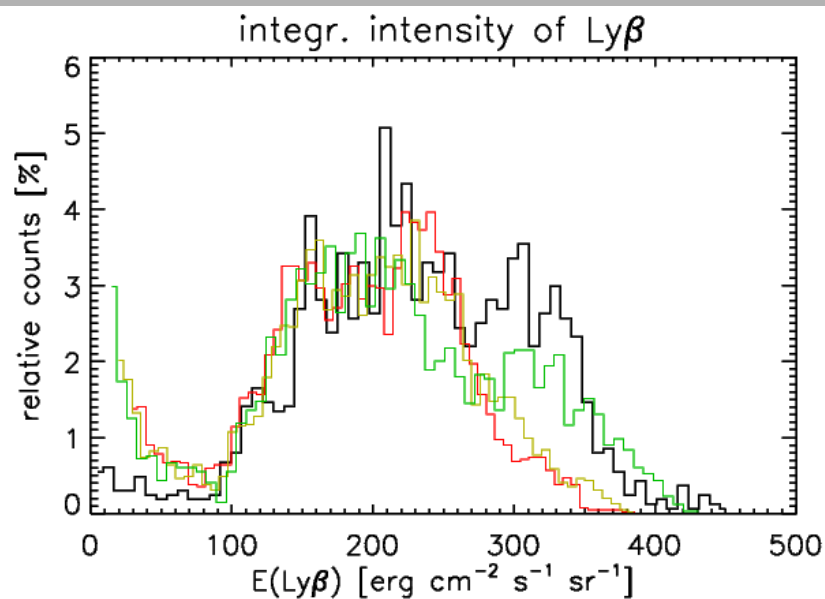


observations

Mo9.5E-5_9000_Bxo6_po2.5E-2, 10 threads
 $\mu=1.0000000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=1.0000000$

Model_A4_Mo4E-6_po0.03, 10 threads
 $\mu=1.0000000$



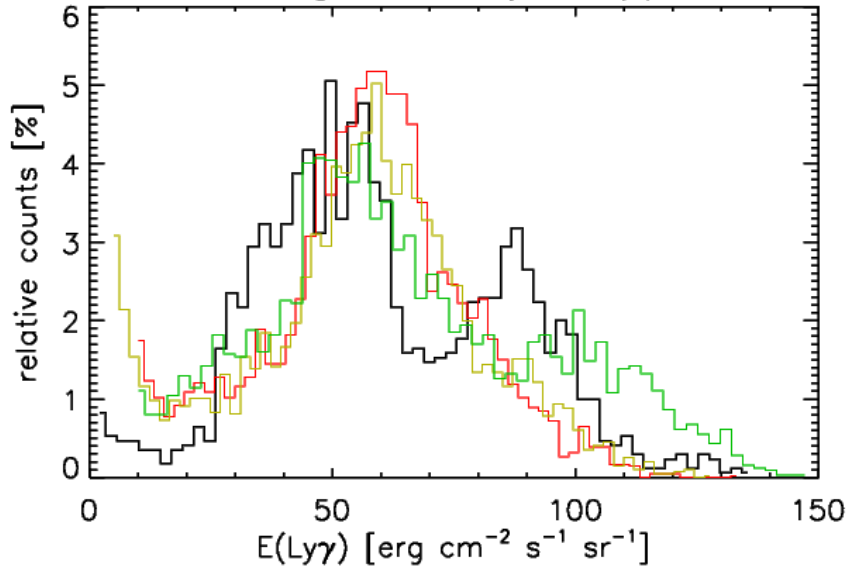
observations

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=1.0000000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=0.86600000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=0.50000000$

integr. intensity of Ly γ



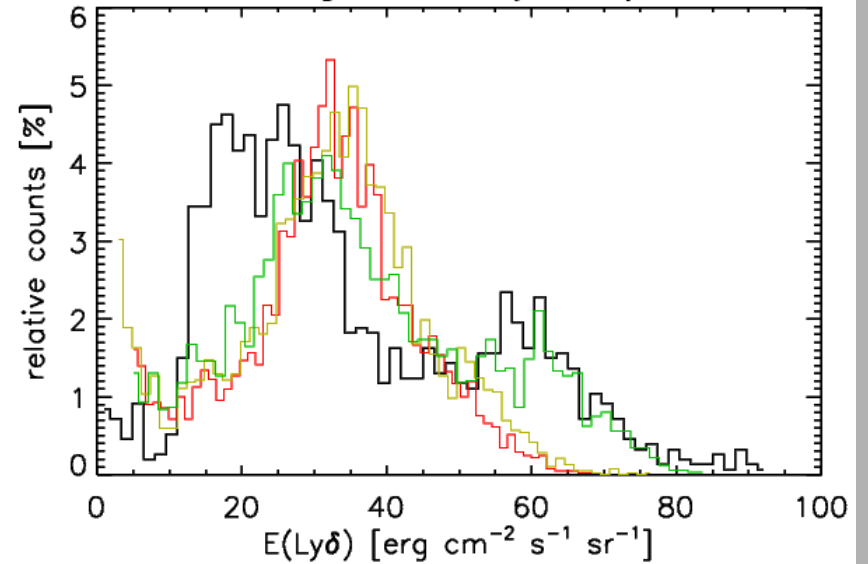
observations

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=1.0000000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=0.86600000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=0.50000000$

integr. intensity of Ly δ

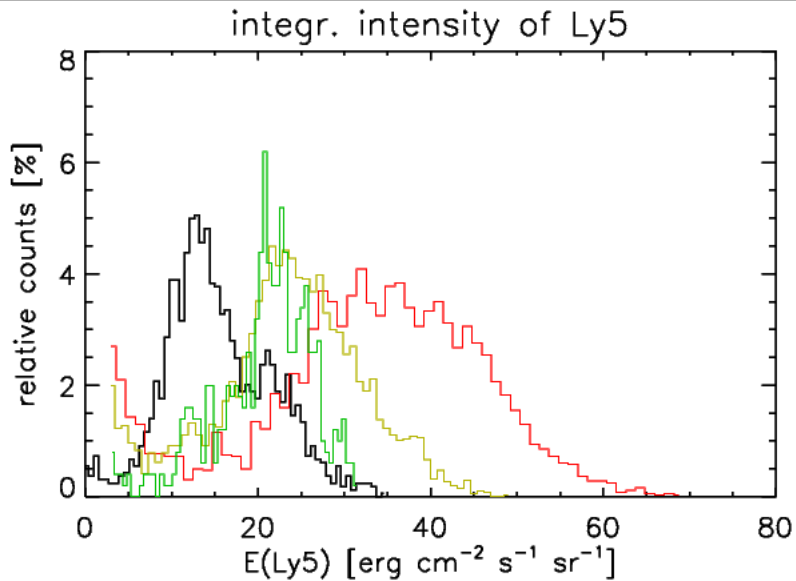


observations

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=1.0000000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=0.86600000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=0.50000000$

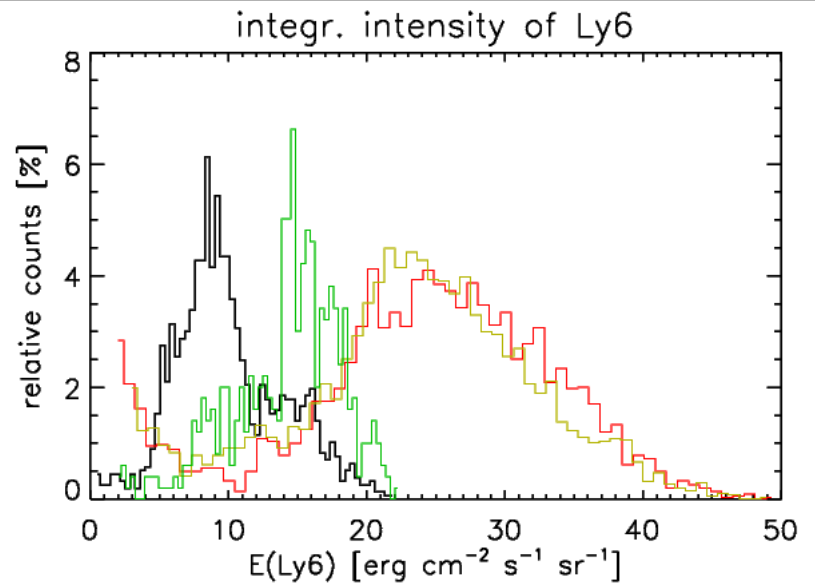


observations

Mo9.5E-5_9000_Bxo6_po2.5E-2, 10 threads
 $\mu=1.0000000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=1.0000000$

Model_A4_Mo4E-6_po0.03, 10 threads
 $\mu=1.0000000$

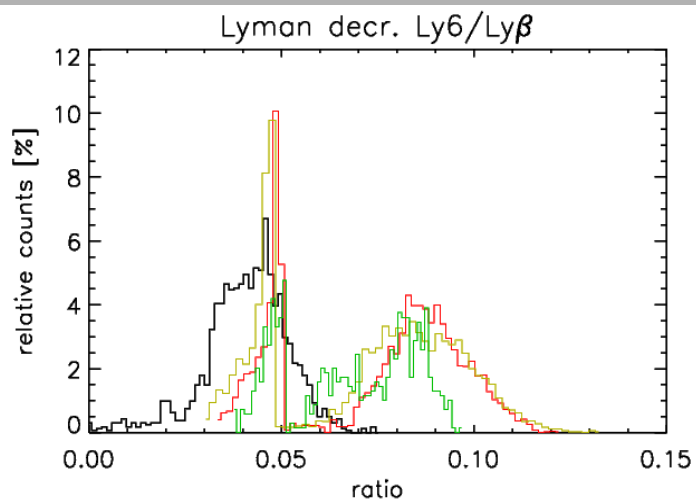
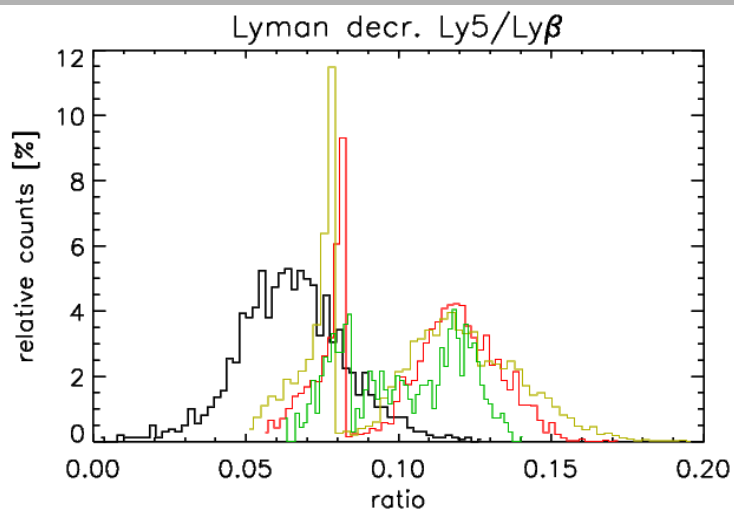
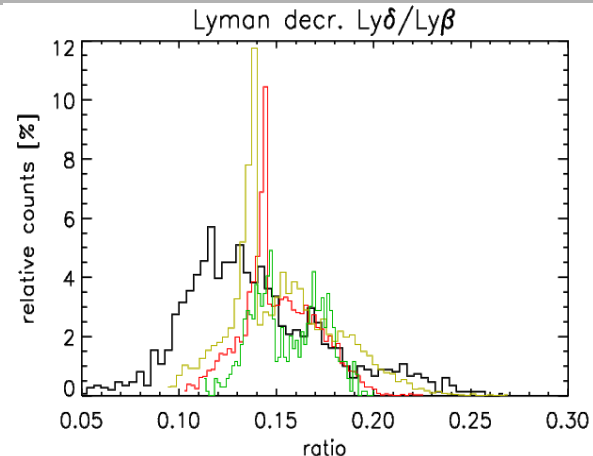
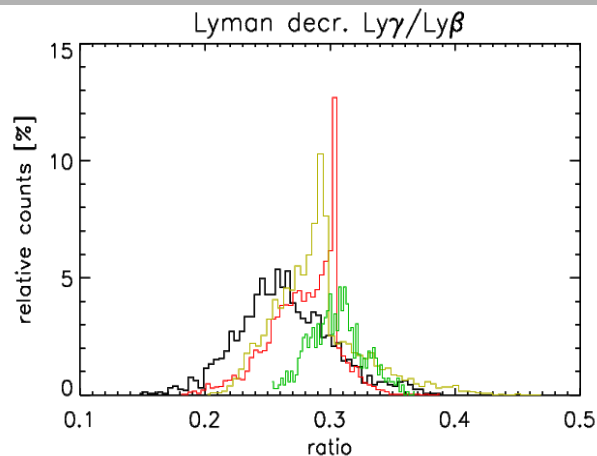
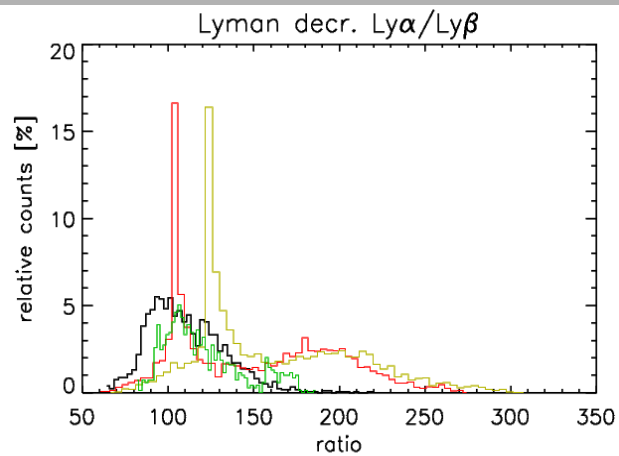


observations

Mo9.5E-5_9000_Bxo6_po2.5E-2, 10 threads
 $\mu=1.0000000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=1.0000000$

Model_A4_Mo4E-6_po0.03, 10 threads
 $\mu=1.0000000$



observations

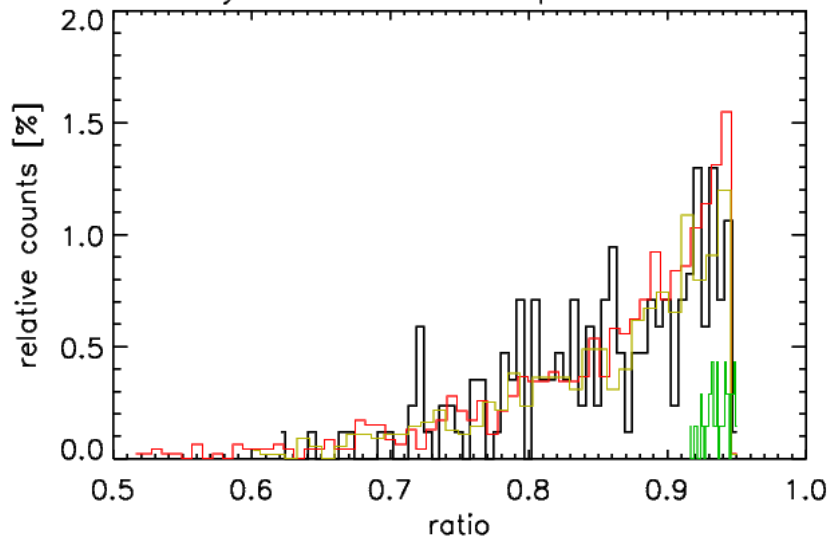
Mo9.5E-5_9000_Bxo6_po2.5E-2, 10 threads
 $\mu=1.0000000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=1.0000000$

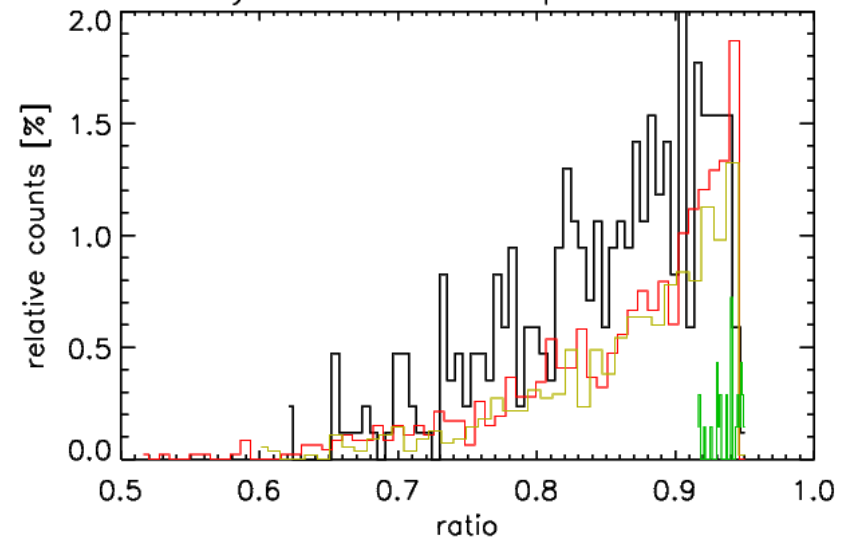
ModelA4_Mo4E-6_po0.03, 10 threads
 $\mu=1.0000000$

**very low sensitivity to μ
 from interval 0.5 – 1 !!!!**

Ly α red-to-BLUE peak ratio



Ly α blue-to-RED peak ratio



observations

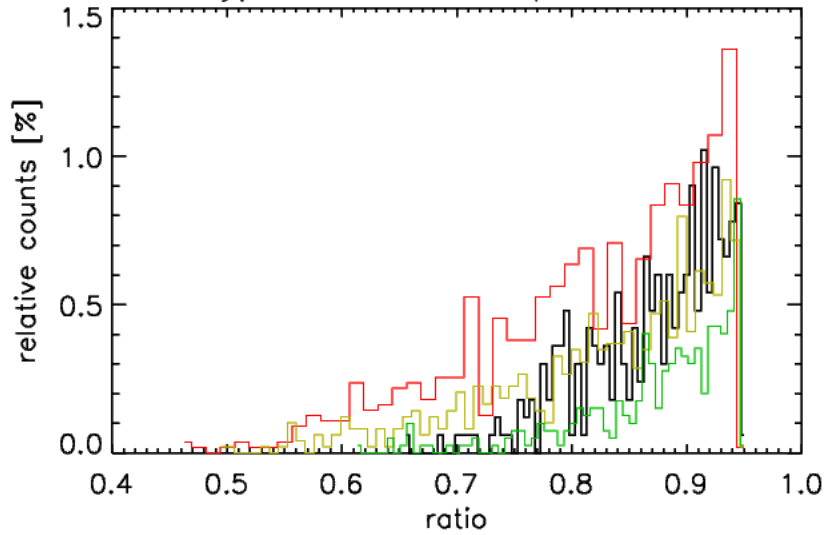
Mo9.5E-5_9000_Bxo6_po2.5E-2, 10 threads
 $\mu=1.0000000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=1.0000000$

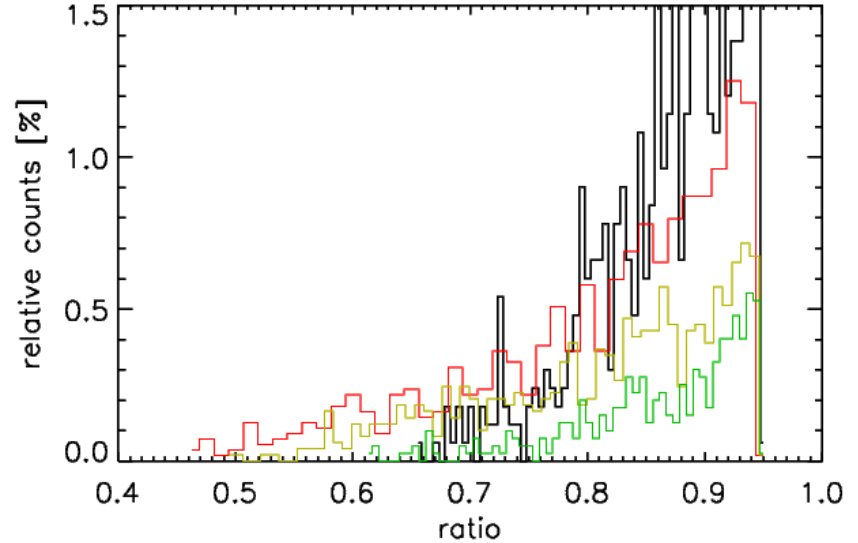
Model_A4_Mo4E-6_po0.03, 10 threads
 $\mu=1.0000000$

**very low sensitivity to μ
from interval 0.5 – 1 !!!!**

Ly β red-to-BLUE peak ratio



Ly β blue-to-RED peak ratio



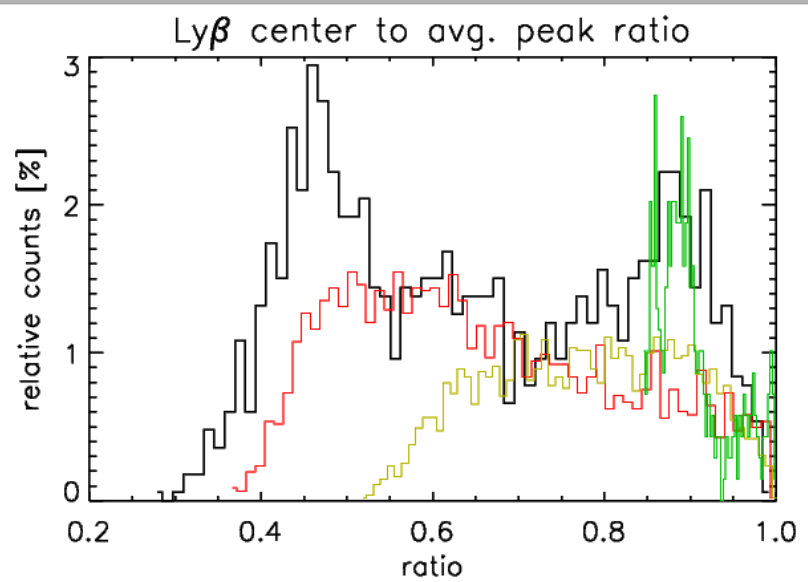
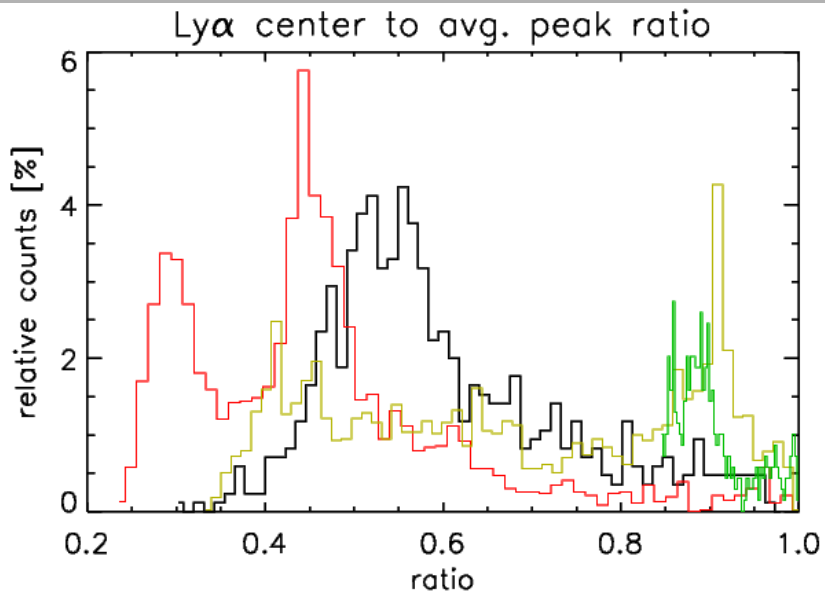
observations

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=1.0000000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=0.86600000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=0.50000000$

similar also for higher Lyman lines



observations

Mo9.5E-5_9000_Bxo6_po2.5E-2, 10 threads
 $\mu=1.0000000$

Model1_8000_Mo8.5E-5_Bxo7_po2E-2, 10 threads
 $\mu=1.0000000$

Model_A4_Mo4E-6_po0.03, 10 threads
 $\mu=1.0000000$

**very low sensitivity to μ
 from interval 0.5 – 1 !!!!**

similar as Ly β also for higher Lyman lines

Conclusions

Models of one thread close to MODEL1 used in model of prominence fine structure are suitable for modelling of prominences with majority of reversed profiles of Ly β – Ly δ (prominences of 25 May and 18 May 2005):

Mo9.5E-5_9000_Bxo6_po2.5E-2	$T_0=9000\text{ K};$ $M_0=9.5 \times 10^{-5}\text{ g cm}^{-2};$	$B_x(0)=6\text{ Gauss};$ $p_0=0.025\text{ dyn cm}^{-2};$	$\gamma_1=10$ $\gamma_2=60$
Model1_8000_Mo8.5E-5_Bxo7_po2E-2	$T_0=8000\text{ K};$ $M_0=8.5 \times 10^{-5}\text{ g cm}^{-2};$	$B_x(0)=7\text{ Gauss};$ $p_0=0.02\text{ dyn cm}^{-2};$	$\gamma_1=10$ $\gamma_2=60$
MODEL1	$T_0=7000\text{ K};$ $M_0=1.1 \times 10^{-4}\text{ g cm}^{-2};$	$B_x(0)=6\text{ Gauss};$ $p_0=0.015\text{ dyn cm}^{-2}$	$\gamma_1=10$ $\gamma_2=60$

Purely emissive profiles of the lines L β – L δ observed at the prominence on 26 April 2007, are due to lower column mass not due magn. field orientation: purely emissive profiles calculated using the MODEL1 with $M_0=1.1 \times 10^{-4}\text{ g cm}^{-2}$ for one thread, occurs at μ almost equal to 0 only (below 0.009 for L β , below 0.052 for L γ and below 0.017 for L δ) \Rightarrow very low probability to observe the prominence with magnetic field oriented almost or even strictly to the observer. Synthetic profiles of L β – L δ computed using MODEL1 at $\mu=0$ are 2 – 4 times higher than the observed ones Using the model Ax1 with $M_0=1 \times 10^{-5}\text{ g cm}^{-2}$, purely emissive profiles can be obtained even at $\mu=0.5$ or lower. These synthetic profiles resemble better observed profiles.

Conclusions – continuation

There is problem with values of reversal depth in modelling the Lyman line profiles – intensities at the reversals are OK, but peaks of synthetic profiles are higher. Including the effect of mutual irradiations of threads made the problem even worse. Even after improvements of the multi-thread model (numerics, using atlas spectrum of BG radiation for IR lines) the problem still persists mainly for the Ly α line

Assuming that optical thickness at the H α center is around unity, intensity in the linecenter can be used to estimate number of the threads in the observed arc of the prominence.

By comparison of histograms of some properties of the observed profiles with those of synthetic profiles by the 18 May prominence it can be seen that models close to MODEL1 cannot explain their behaviour completely and that presence of low density threads are also needed.

Question of whether there is one general prominence-corona transition region (PCTR) or each thread has its own or both exist still open – it is to be solved with applying threads with different plasma conditions in the multithread model

Thank for your attention

This work was supported by grant P209/12/0906 of the Grant Agency of The Czech republic and by the project VEGA 2/0108/12 of the Science Grant Agency. P.S. acknowledges support from the Slovak Research and Development Agency project under contract No. APVV-0816-11.