



Slovenská
Astronomická
Spoločnosť
pri Slovenskej akadémii vied

LONG-TERM EVOLUTION OF COLOUR INDICES OF THE SYMBIOTIC SYSTEM AX PERSEI

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Bezovec 2023 – Conference of Young Astronomers

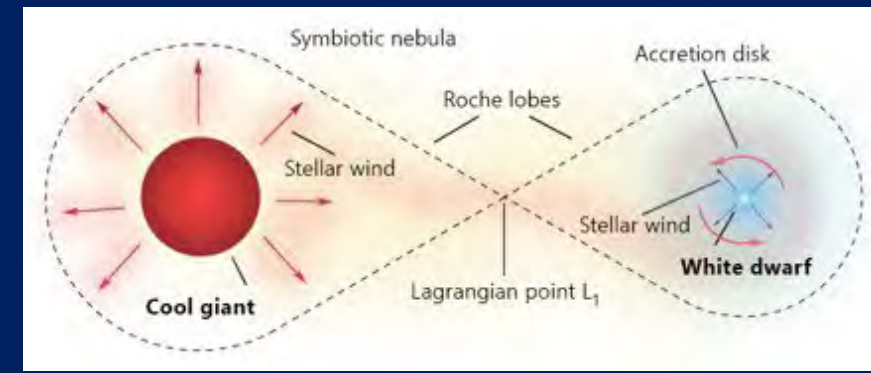
The contribution was supported by the grants APVV-20-0148
and VVGS-2023-2517

THE GOALS OF OUR RESEARCH

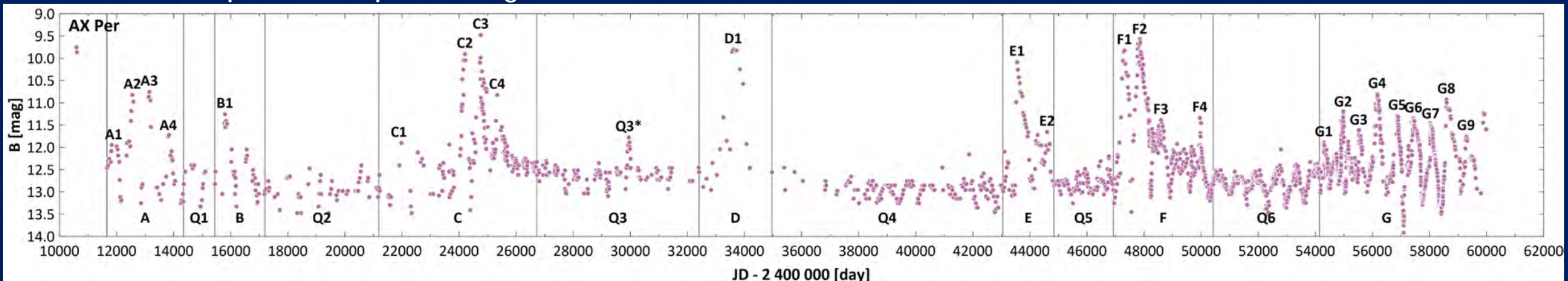
- Detailed analysis of the long-term evolution of $U-B$, $B-V$, $V-R_c$, $V-I_c$ and R_c-I_c colour indices of the symbiotic system AX Persei.
- Determination and interpretation of the timing and the magnitudes of the extrema of the curves of indices.
- Correlation analysis of the curves of indices.

AX PERSEI

- Symbiotic character discovered by P. W. Merrill and L. Humason in 1932.
- Consisting of a M4.5 III giant and a white dwarf surrounded by symbiotic nebula.
- Classical symbiotic binary – active and quiescent phases.
- Outbursts – the sign of activity, significant photometric and spectroscopic changes.



The model of symbiotic system.
Source: Merc et al., 2019



Light curve of AX Persei in the B filter.

ORBITAL PERIOD OF AX PERSEI

- Basic parameter of binary systems.
- Light curves in U , B , V , R_c and I_c filters: orbital period of (680.4 ± 4.3) days (Mártonfi et al. 2021)
- Radial velocity curve (radial velocity values adopted from literature and obtained by the cross-correlation analysis): orbital period of (681.2 ± 4.2) days (Mártonfi & Gális 2022).
- Eclipses, minima and maxima of the light curve of the system are regularly repeated with this period.
- Employed linear ephemeris from Skopal et al. (2011):

$$JD_{min} = 2\,447\,551.26 + E \times 680.83$$

DATA PROCESSING AND ANALYSIS

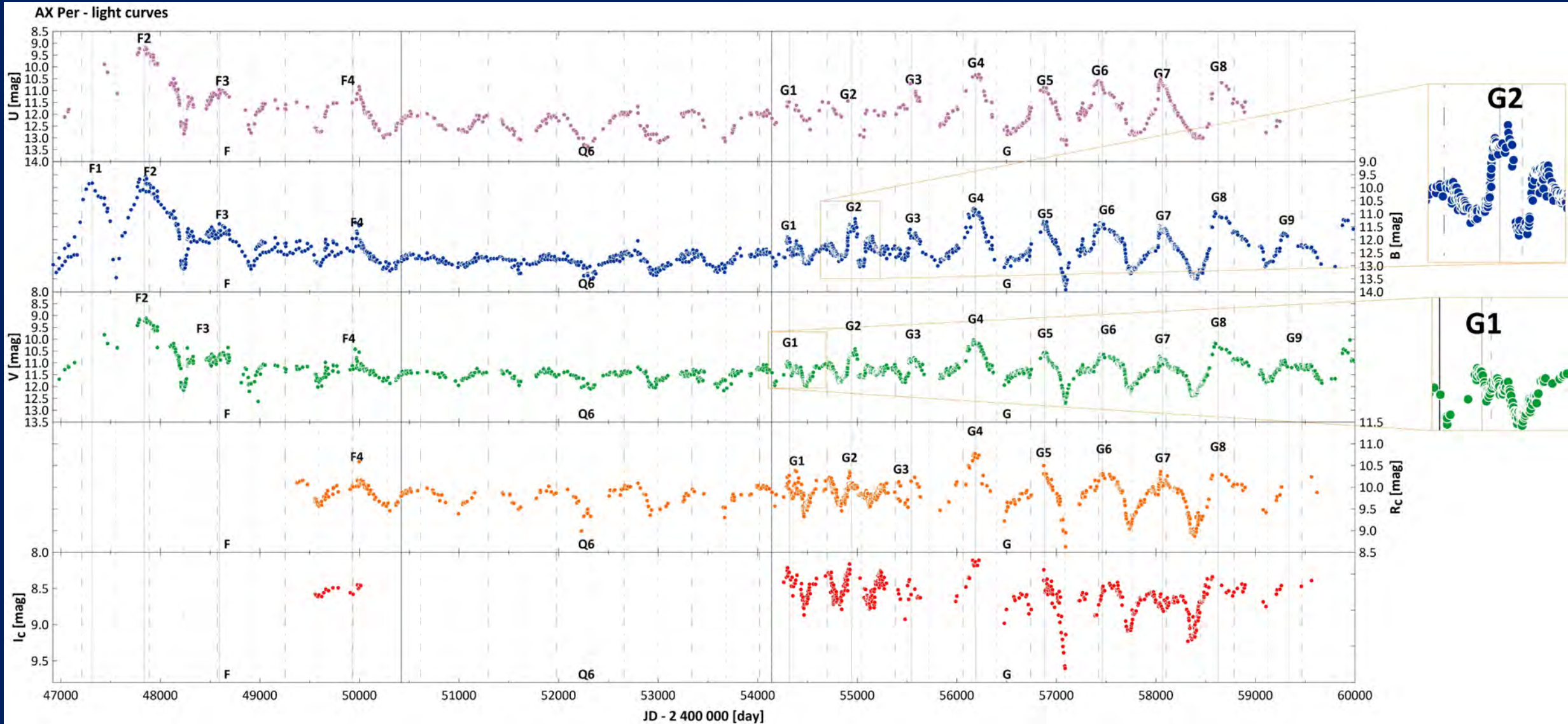
- Sources of the photometric data: available scientific publications – papers, telegrams (Atel, CBET), circulars (IAUC) and the database of **American Association of Variable Star Observers*** (AAVSO).
- Magnitudes in U , B , V , R_c and I_c filters; light curves constructed and analysed in the previous research (Mártonfi et al. 2021).

*<https://www.aavso.org/>

DATA PROCESSING AND ANALYSIS

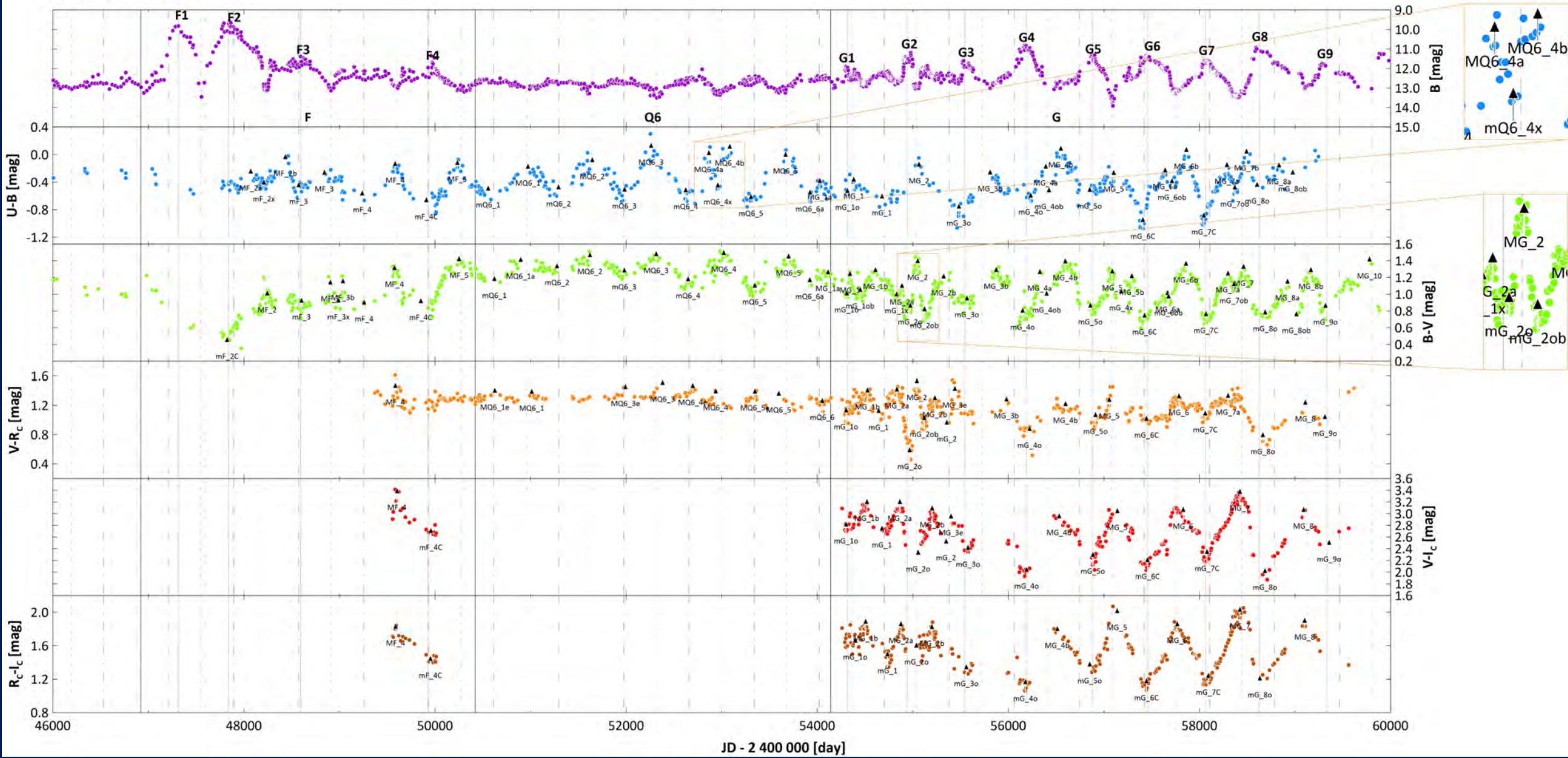
- **Derivation of colour indices:** searching for pair of magnitudes (in the filters of the given colour index) with the Julian date (JD) difference < 1 day.
- Omitting the points with the deviation $> 2\sigma$ of the given part of the curve of indices.
- The curves of the 5-day averages of the indices were constructed.
- **Extremes determination:** MAVKA software, 3 methods of approximation (polynomial, parabolic spline, asymptotic parabola).
- **Correlation analysis:** *CorrLAB* software, classical and discrete correlation analysis.

THE RESULTS OF THE RESEARCH SPECIFIC CHARACTERISTICS OF THE LIGHT CURVES OF AX PERSEI



THE CURVES OF INDICES OF AX PERSEI

AX Per - colour indices



CORRELATION ANALYSIS OF COLOUR INDICES

- Strong correlation between light curves mutually (except I_c light curve) and light curves and colour indices.
- Curves of indices: mutual correlation coefficients are lower than with light curves.
- $V-R_c$ in Q6: no correlation with other light and indices curves: different character of variations during the Q6 phase.
- $U-B$: lower correlation coefficients with light curves and other curves of indices due to the double-peak structures during reddenings.

	U	B	V	R_c	U-B	B-V	V- R_c
U	1,00	0,87	0,80	0,84	0,92	0,83	-0,09
B	0,87	1,00	0,84	0,81	0,69	0,68	0,10
V	0,80	0,84	1,00	0,86	0,61	0,45	0,15
R_c	0,84	0,81	0,86	1,00	0,68	0,61	-0,02
U-B	0,92	0,69	0,61	0,68	1,00	0,77	-0,25
B-V	0,83	0,68	0,45	0,61	0,77	1,00	-0,10
V- R_c	-0,09	0,10	0,15	-0,02	-0,25	-0,10	1,00

Correlation coefficients – phase Q6

	U	B	V	R_c	I_c	U-B	B-V	V- R_c	V- I_c	R_c - I_c
U	1,00	0,87	0,83	0,83	0,50	0,60	0,72	0,56	0,83	0,80
B	0,87	1,00	0,96	0,84	0,61	0,31	0,78	0,72	0,82	0,69
V	0,83	0,96	1,00	0,88	0,71	0,22	0,62	0,73	0,82	0,70
R_c	0,83	0,84	0,88	1,00	0,76	0,32	0,49	0,46	0,76	0,75
I_c	0,50	0,61	0,71	0,76	1,00	0,04	0,19	0,44	0,40	0,37
U-B	0,60	0,31	0,22	0,32	0,04	1,00	0,46	-0,02	0,33	0,50
B-V	0,72	0,78	0,62	0,49	0,19	0,46	1,00	0,51	0,60	0,49
V- R_c	0,56	0,72	0,73	0,46	0,44	-0,02	0,51	1,00	0,67	0,43
V- I_c	0,83	0,82	0,82	0,76	0,40	0,33	0,60	0,67	1,00	0,88
R_c - I_c	0,80	0,69	0,70	0,75	0,37	0,50	0,49	0,43	0,88	1,00

Correlation coefficients – phase G

CONCLUSION

- **Specific characteristics of the light curves:** Q6: variations due to reflection effect and eclipses, active phases: strong effect of outbursts, different orbital phase of outburst in the phase G.
- **The curves of indices:** Q6: variations due to the reflection effect (except $V-R_c$ colour index), reddenings in primary minima, blueings in the phase 0.5 or outbursts, double-peak structures of reddenings.
- **Correlation analysis of colour indices:** strong correlation between light curves and curves of indices, lower mutual coefficients between curves of indices, different behaviour of $V-R_c$ (Q6) and $U-B$ (G) curves.
- **Future research:** period analysis of curves of indices, colour index – magnitude diagram, new ephemeris of AX Persei.

REFERENCES

- Mártonfi, P., Gális, R. 2022, JASB, 3(1), 15
- Mártonfi, P., Gális, R., Merc, J. 2021, OEJV, 220, 26
- Merc, J., Gális, R., Wolf, M. 2019, Astronomische Nachrichten, 340, 598
- Skopal, A., Tarasova ,T. N., Cariková Z. et al. 2011, A&A, 536, A27

ACKNOWLEDGEMENTS

- I would like to thank to Assoc. Prof. RNDr. Rudolf Gális, PhD. and RNDr. Jaroslav Merc, PhD. for help, valuable advice and comments.
- This contribution was supported by *Slovak Research and Development Agency* under the contract No. APVV-20-0148 and *VVGS PF UPJŠ* under the contract No. VVGS-2023-2517.

THANK YOU FOR YOUR ATTENTION