

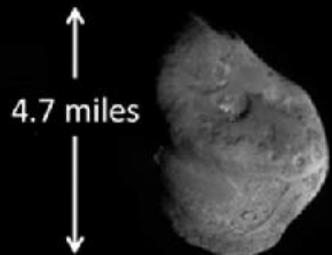
Observational evidence for a surface evolution trend of JFCs

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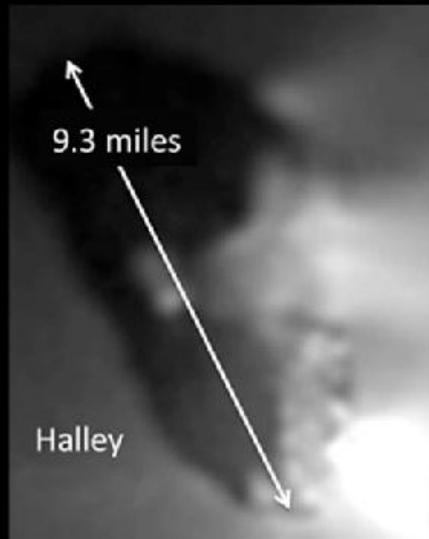


*In collaboration with: Pedro Lacerda, Colin Snodgrass, Simon Green, S. C. Lowry,
Y. R. Fernandez, C. Tubiana, A. Fitzsimmons, H. H. Hsieh, P. Nikolov, T. Bonev.*



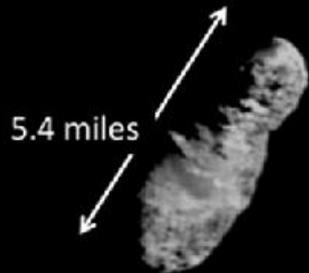
4.7 miles

9P/Tempel 1



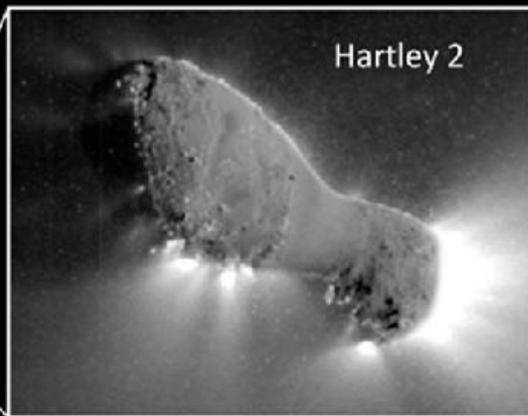
9.3 miles

Halley



5.4 miles

Borrelly



Hartley 2

1.25 miles



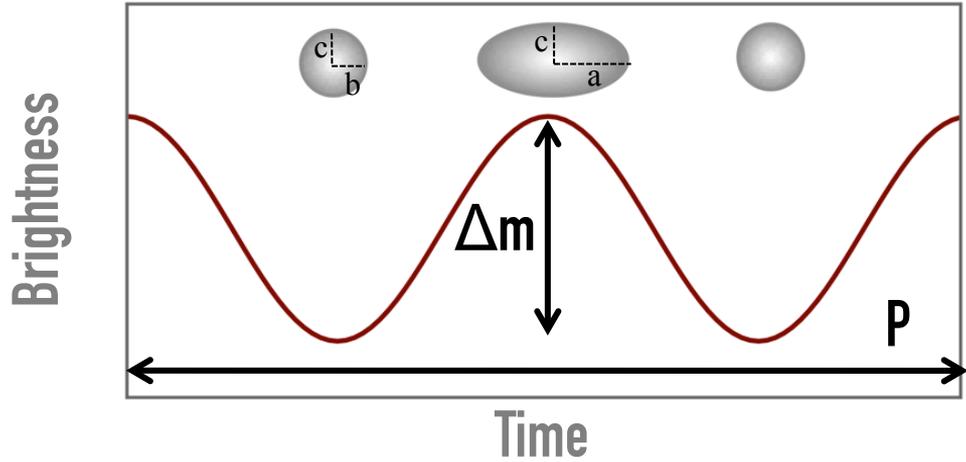
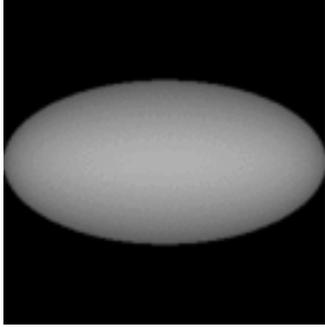
3.4 miles

Wild 2

67P/C-G

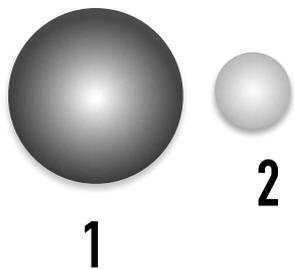


Rotational Lightcurves



Surface

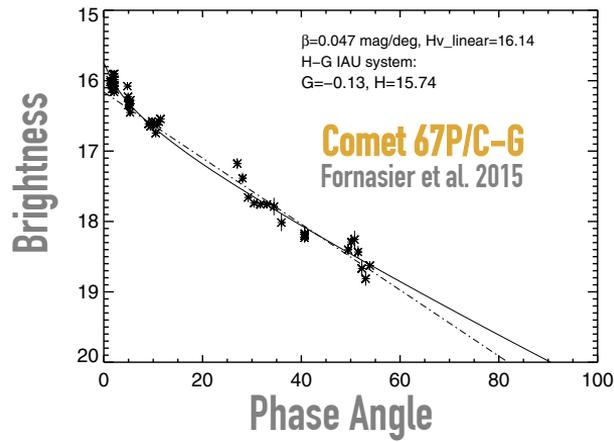
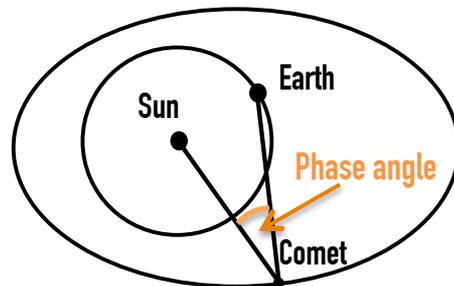
Albedo



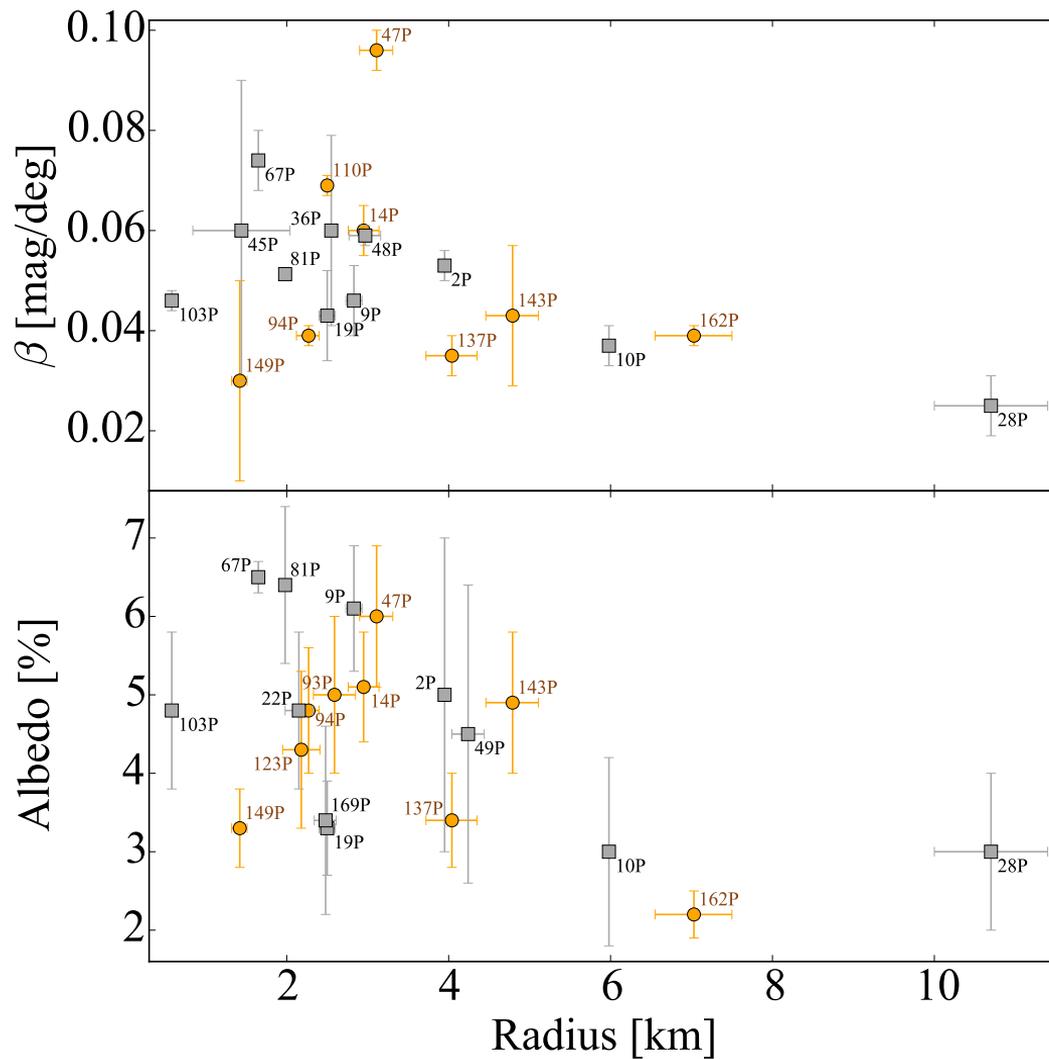
$$m_1 = m_2, R_1 > R_2$$
$$A_1 < A_2$$

If the size is known
(e.g. thermal IR measurements)

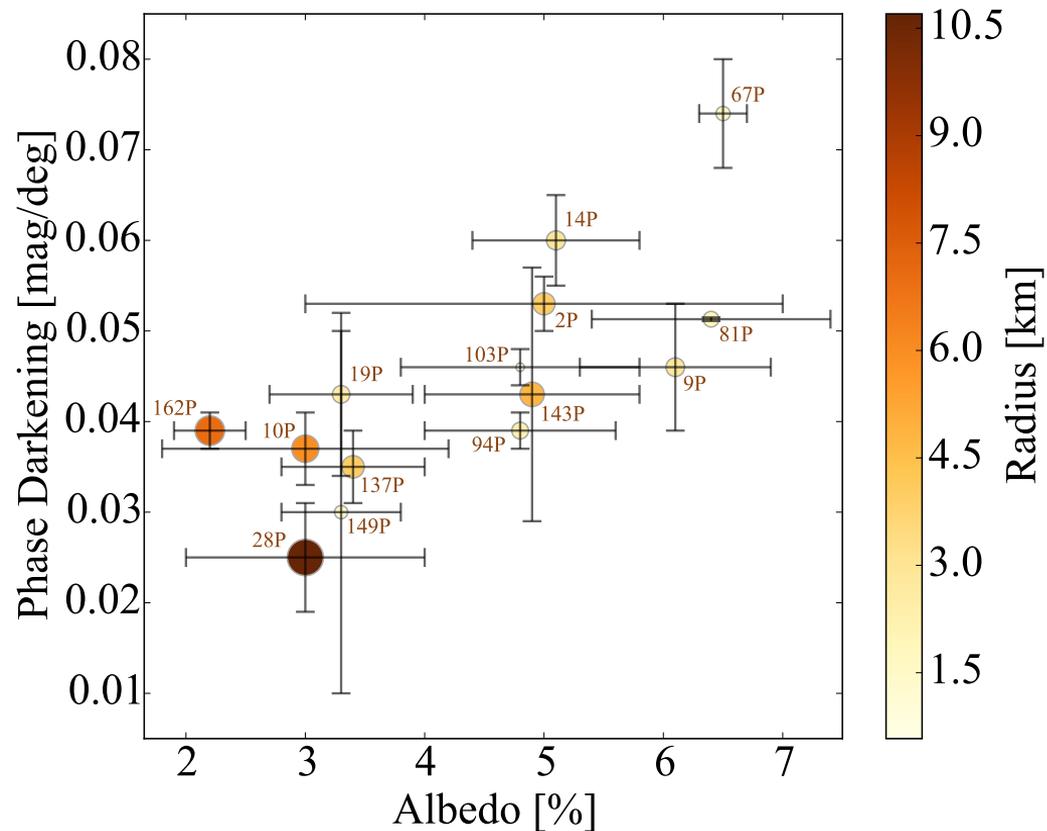
Phase Darkening



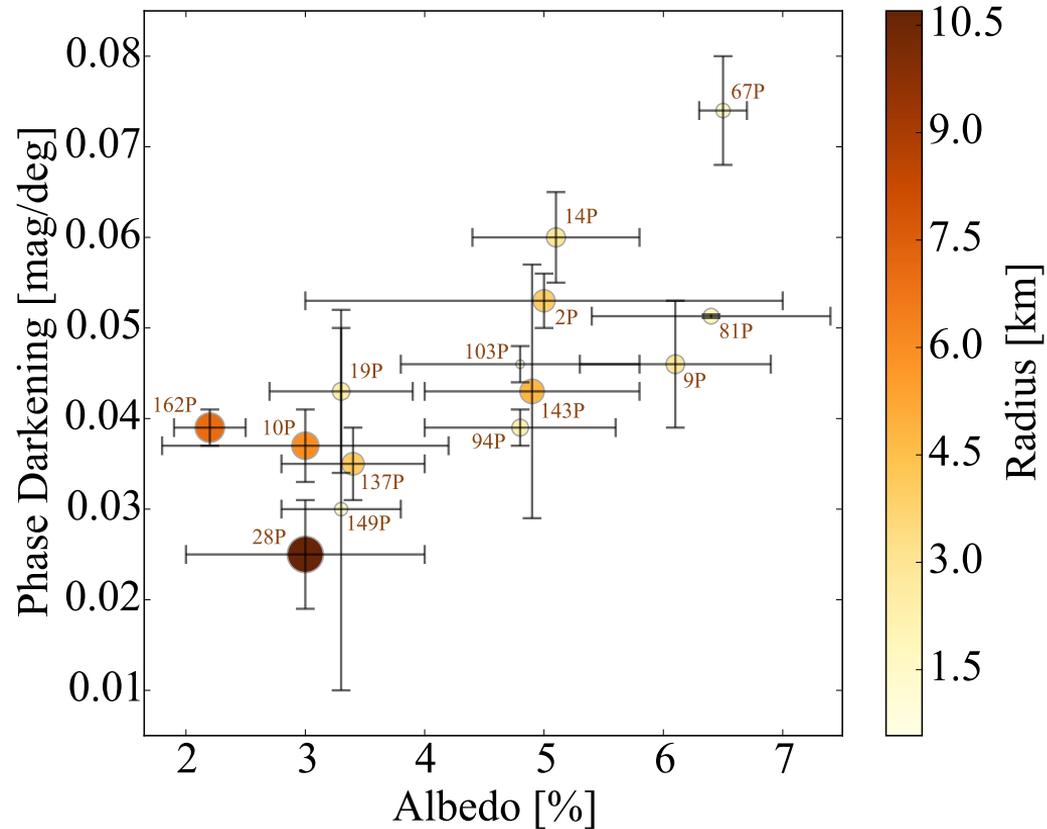
JFC Surfaces



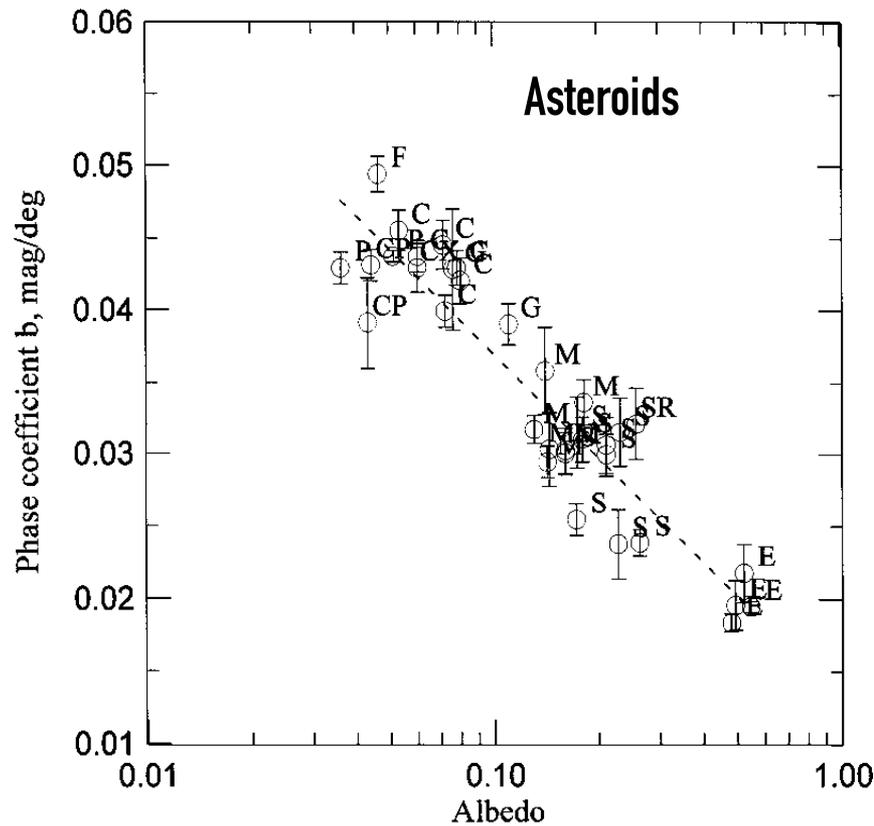
JFC Surfaces



JFC Surfaces

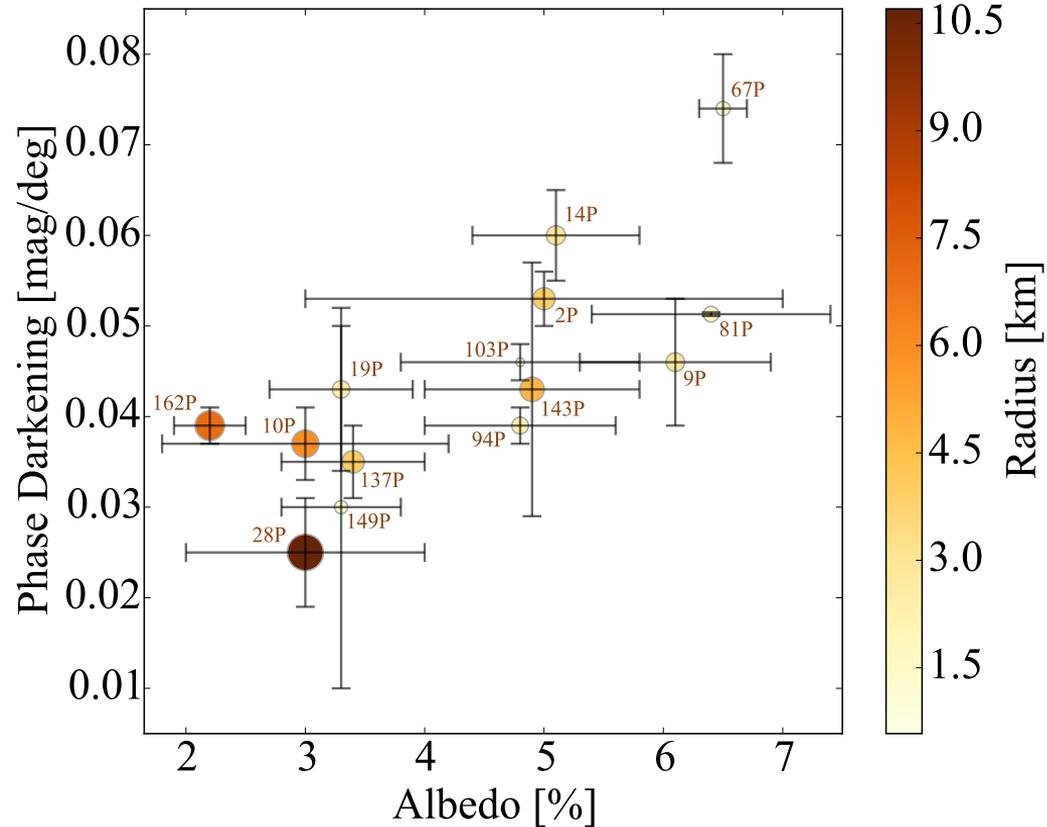


Kokotanekova et al. 2018

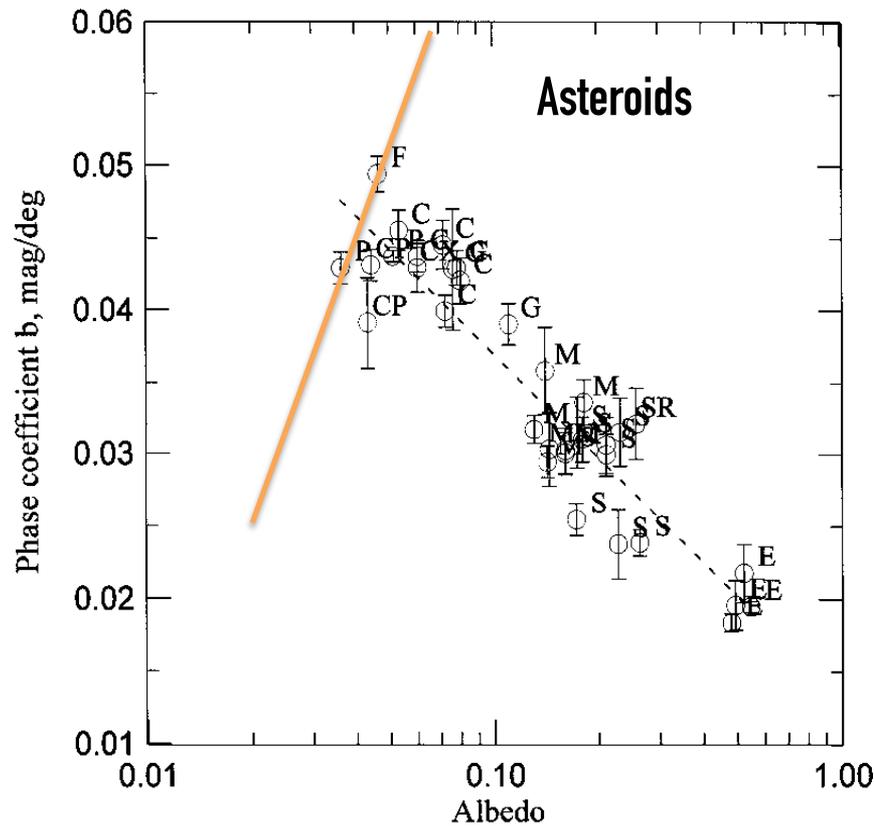


Belskaya & Shevchenko. 2000

JFC Surfaces

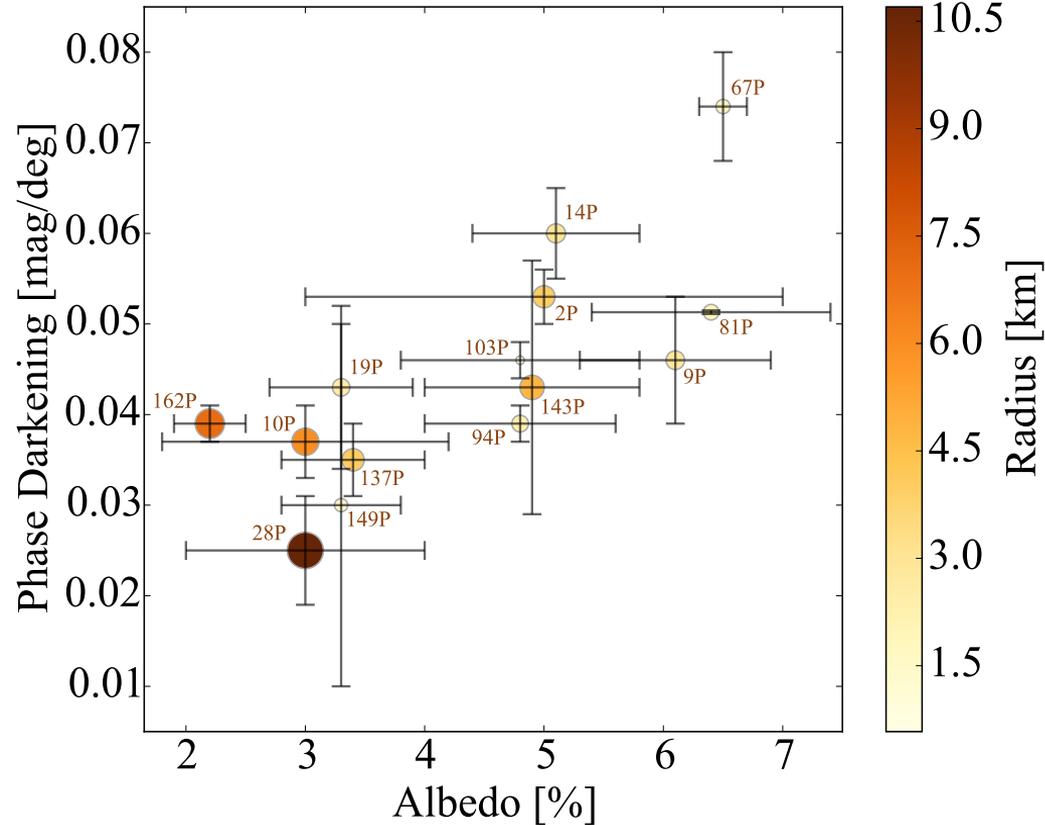


Kokotanekova et al. 2018

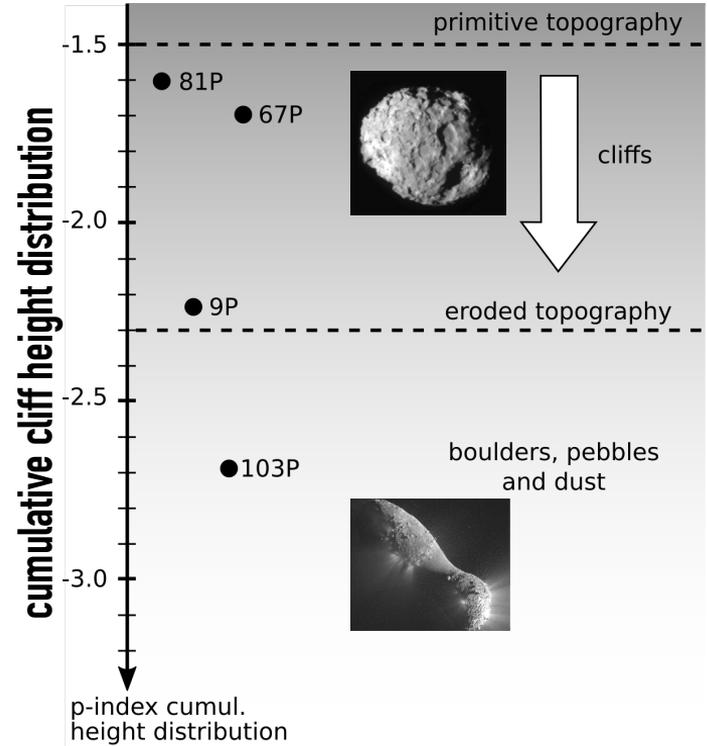


Belskaya & Shevchenko. 2000

Interpretation: 1. Cliff Height Distribution



Kokotanekova et al. 2018



Vincent et al. 2017

Ip et al. 2016

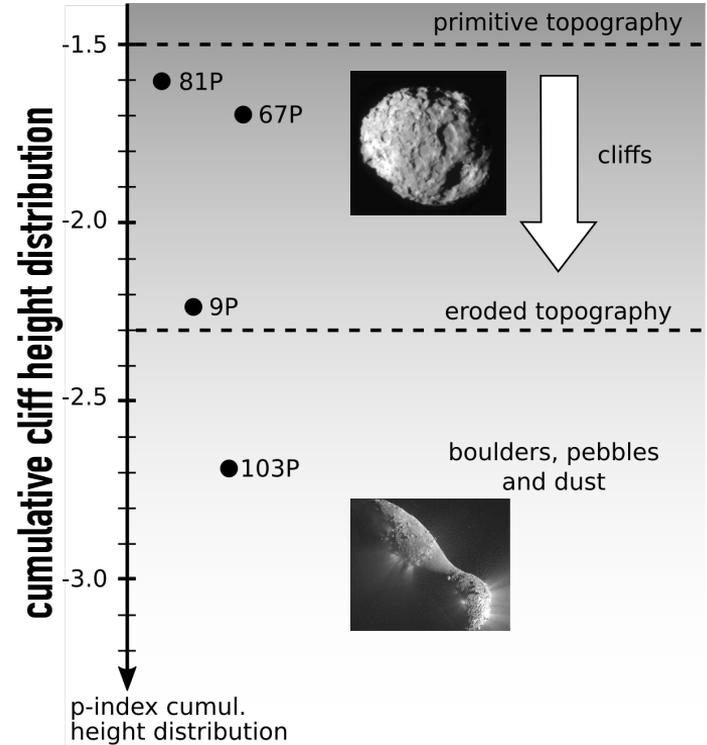
Basilevsky & Keller 2006

3. In situ surface properties observations

Surface roughness and photometric properties are related:

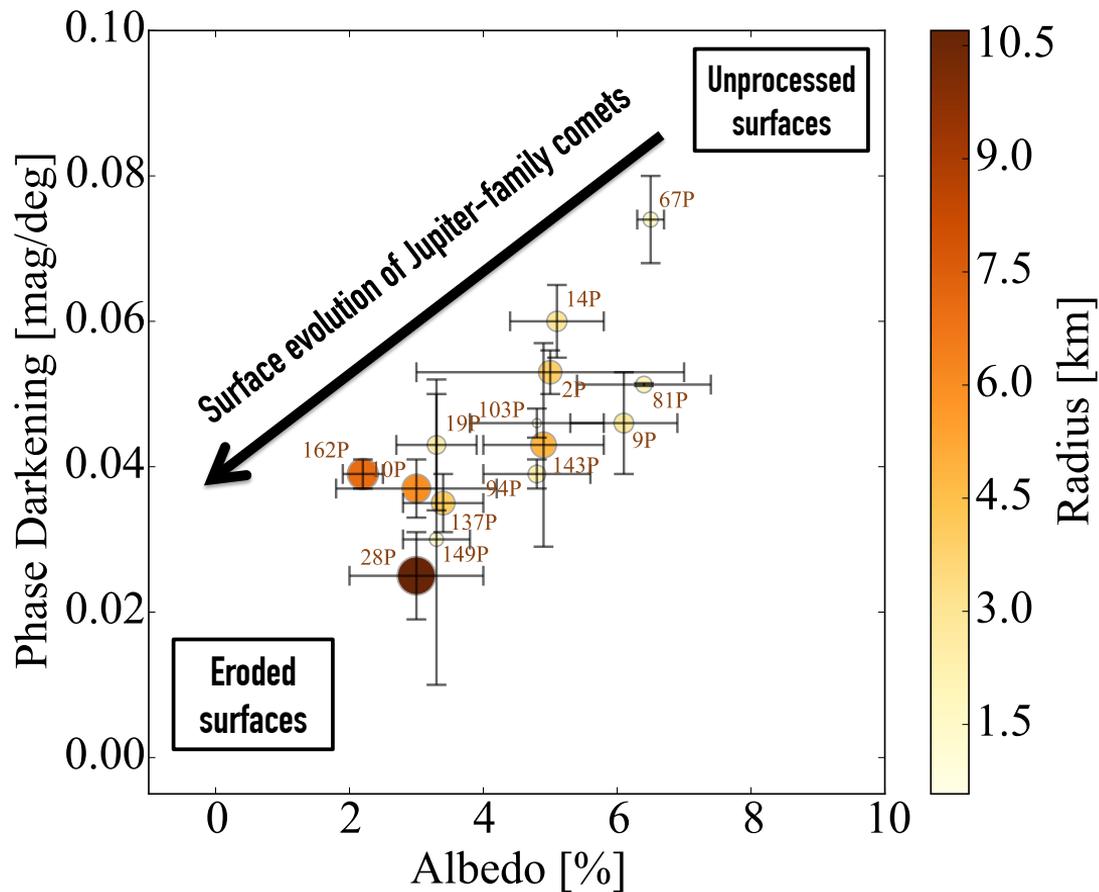
- Comparison between the comets visited by spacecraft
- Comparison between 'rough' regions – Anuket and Ash and 'smooth' regions: Hapi and Imhotep

Longobardo et al. (2017)

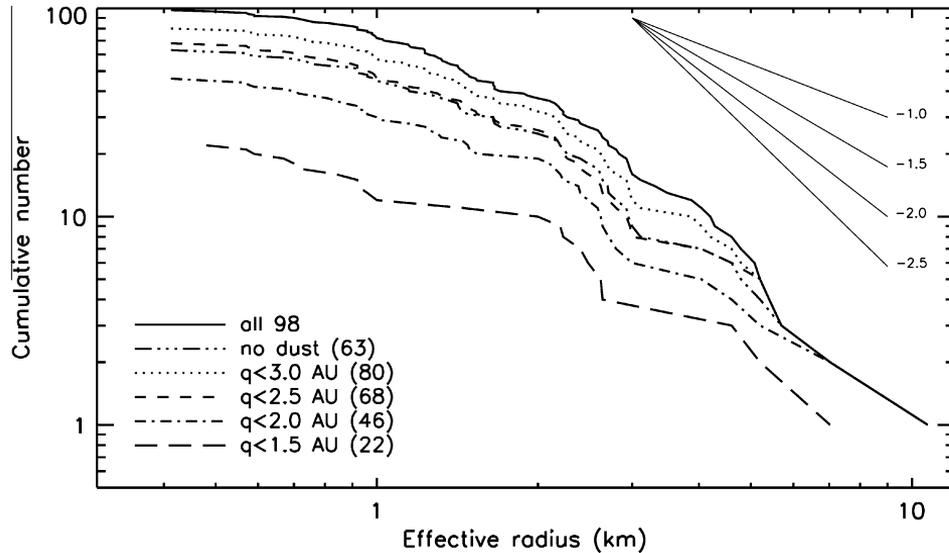


Vincent et al. 2017

JFC Surface Evolution

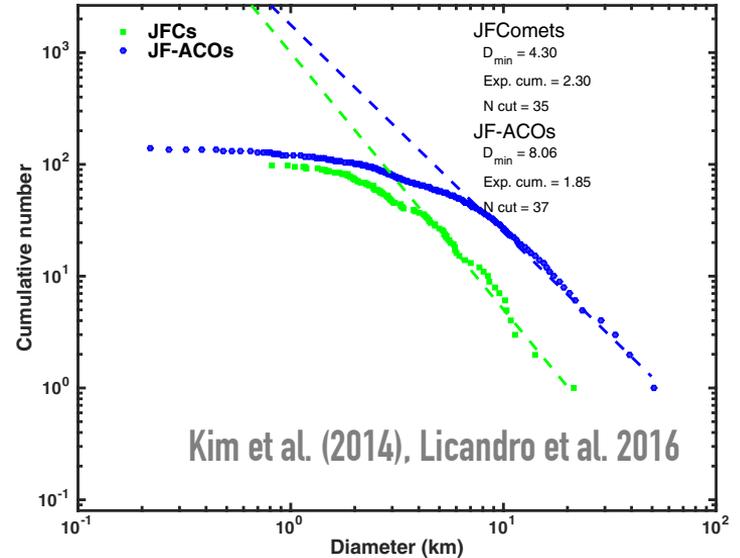


Cumulative size distribution of JFCs



“Excess of comets with
radii 3–6 km”

Fernandez et al. (2013)



Flatter distribution for ACOs than for JFCs –
> The larger nuclei preferentially survive
the active phase

2. Dynamical simulations

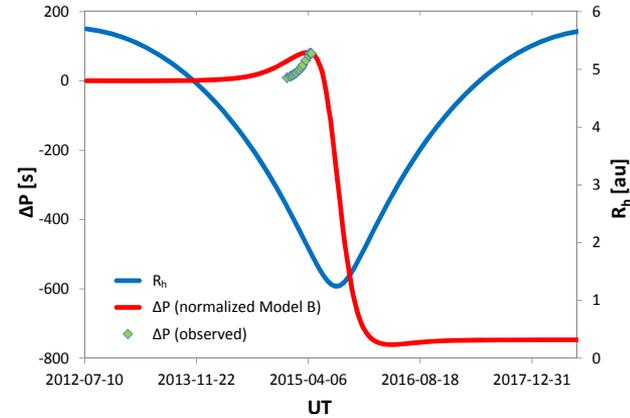
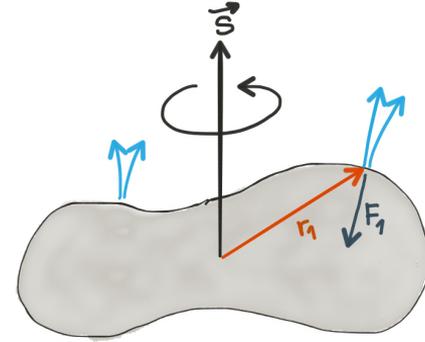
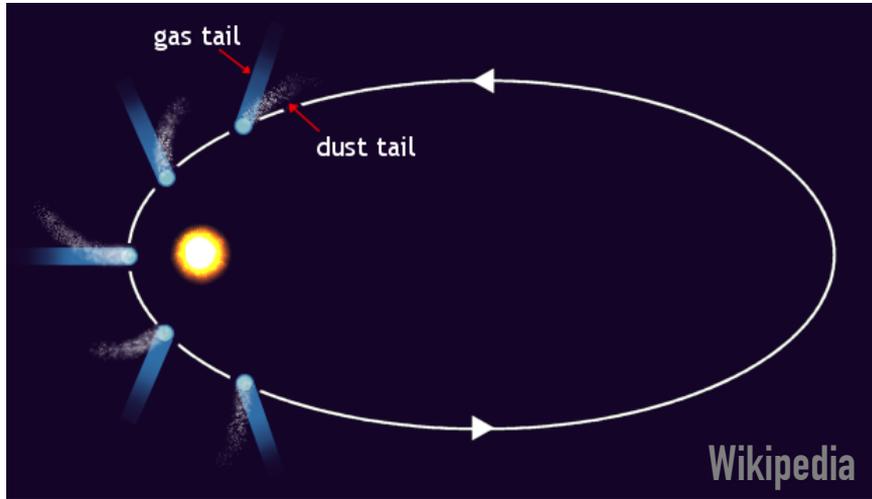
- **Dynamical simulations require longer physical lifetimes for larger nuclei in order to reproduce the number of observed comets**

Di Sisto et al. 2009

Nesvorny et al. 2017

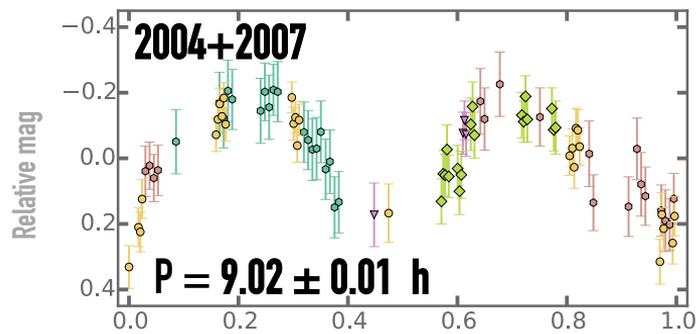
Rickman et al. 2017

Spin Changes Due to Outgassing

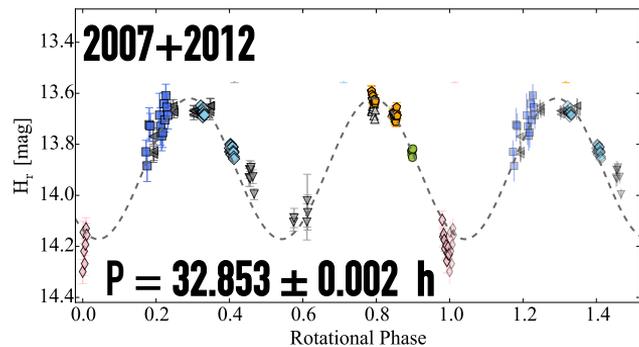
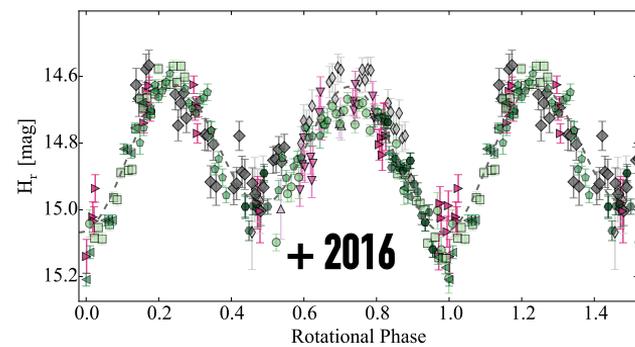


67P

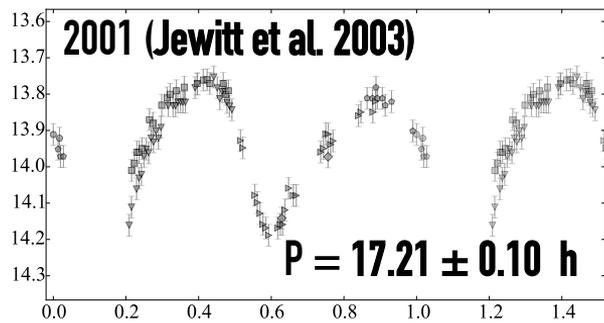
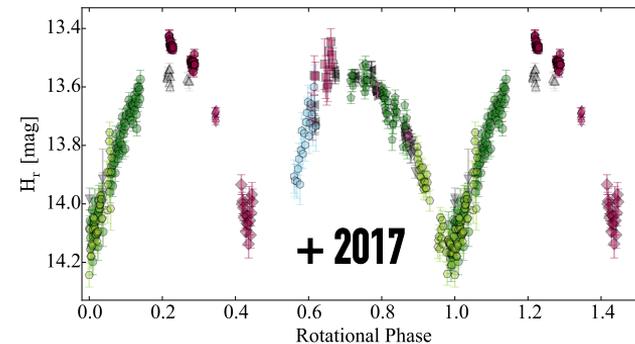
Keller et al 2015;



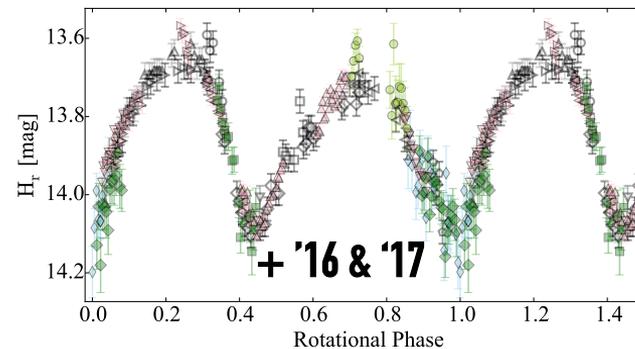
14P
 $\Delta P < 4.2 \text{ min}$



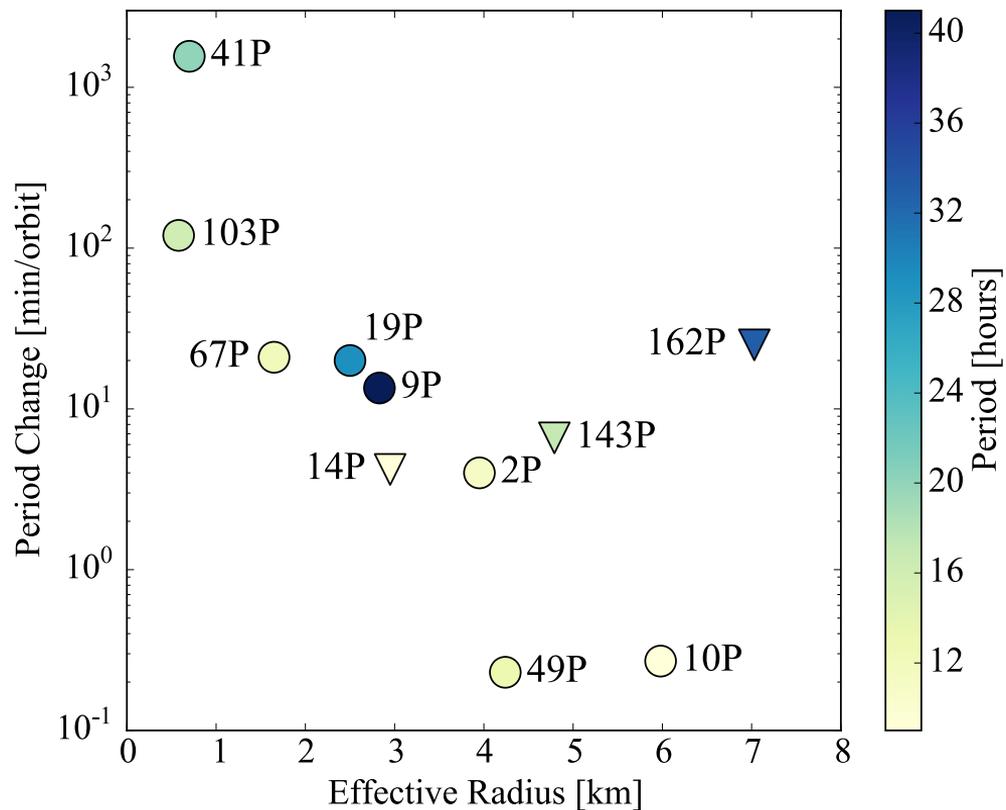
162P
 $\Delta P < 25 \text{ min}$



143P
 $\Delta P < 6.6 \text{ min}$



Spin Changes Due to Outgassing

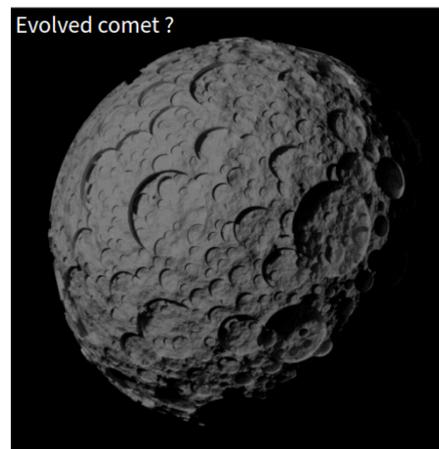
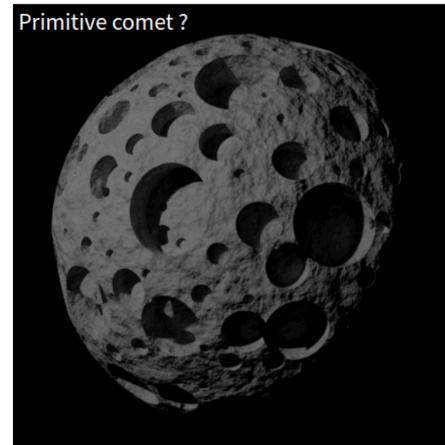
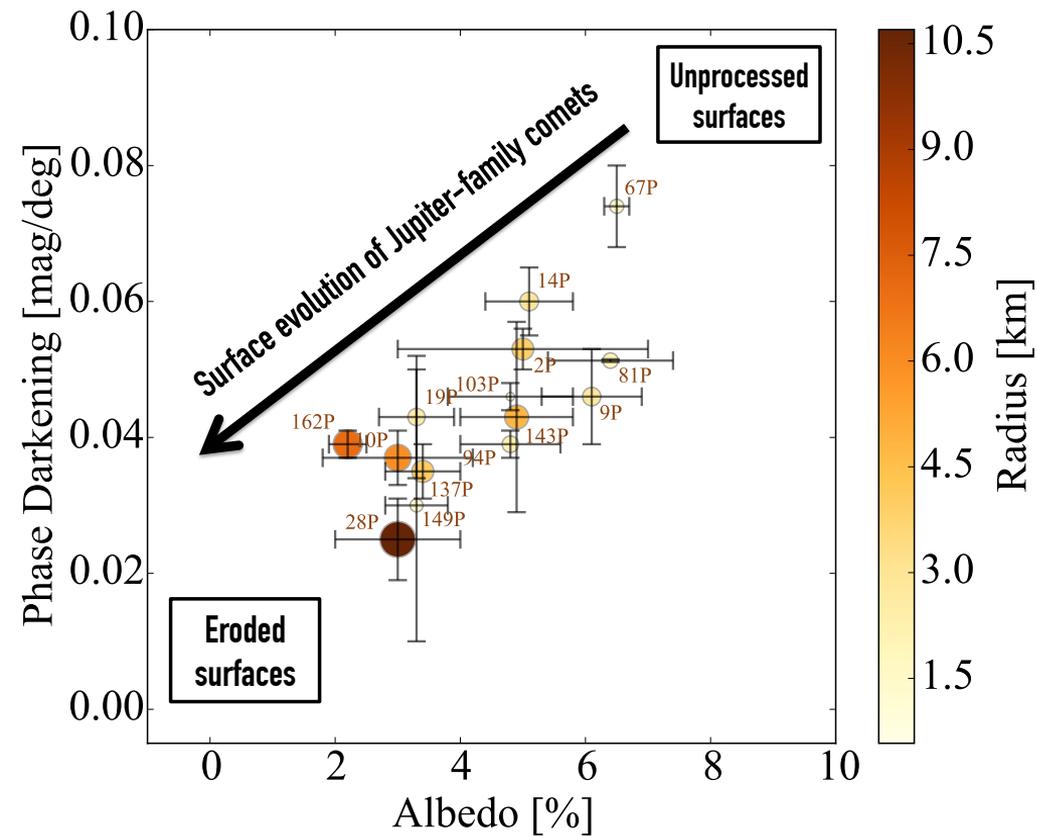


Mechanisms to prolong the physical lifetime of large JFCs

1. **More difficult to spin up (larger and less active)**
2. Larger nuclei have experienced larger hydrostatic compression, and therefore have larger tensile strengths which obstruct the activity (Gundlach & Blum 2016)
3. Larger comets have larger gravitational potential, they will retain more infalling dust and will quickly build an insulating layer (e.g. Thomas 2015)

Conclusion: Larger comet nuclei have larger physical lifetimes

JFC Surfaces



Synthetic models from
Vincent, J.-B., LPSC 2018

Summary

- 1. Photometric time series are a powerful tool to study the physical properties and surface characteristics of JFCs.**
- 2. JFCs with higher albedo have steeper phase functions – a possible evolutionary trend.**
- 3. Large nuclei do not experience significant period changes, and are less likely to undergo a break-up.**
- 4. Large nuclei are more likely to reach the later stages of surface evolution.**