

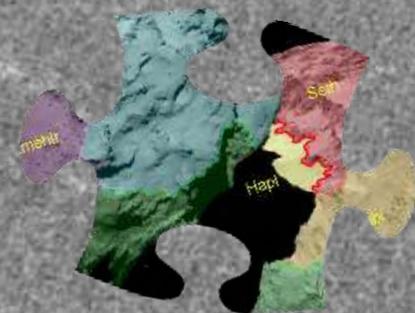
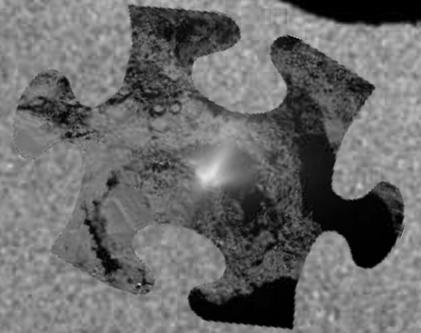
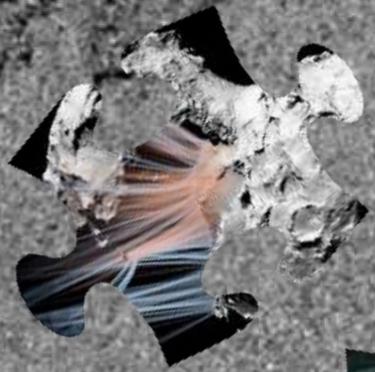


A Rosetta/OSIRIS View on the Jigsaw Puzzle of Modern Cometary Science

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1. What is really known?
2. What is relevant for the bigger picture?
(formation and evolution)





Fact sheet

Mass: 9.982×10^{12} kg

Volume: 18.7 km^3

Density: 0.53 g/cm^3

Ice: 14 – 33 wt%

Porosity: $\sim 71 \%$

Rotation period*: 12 hours

Albedo: 5.9 %

* in 2016

The shape poses challenges in modelling
(but it shouldn't be much of a surprise).



1P/Halley - 1986



19P/Borelly - 2001

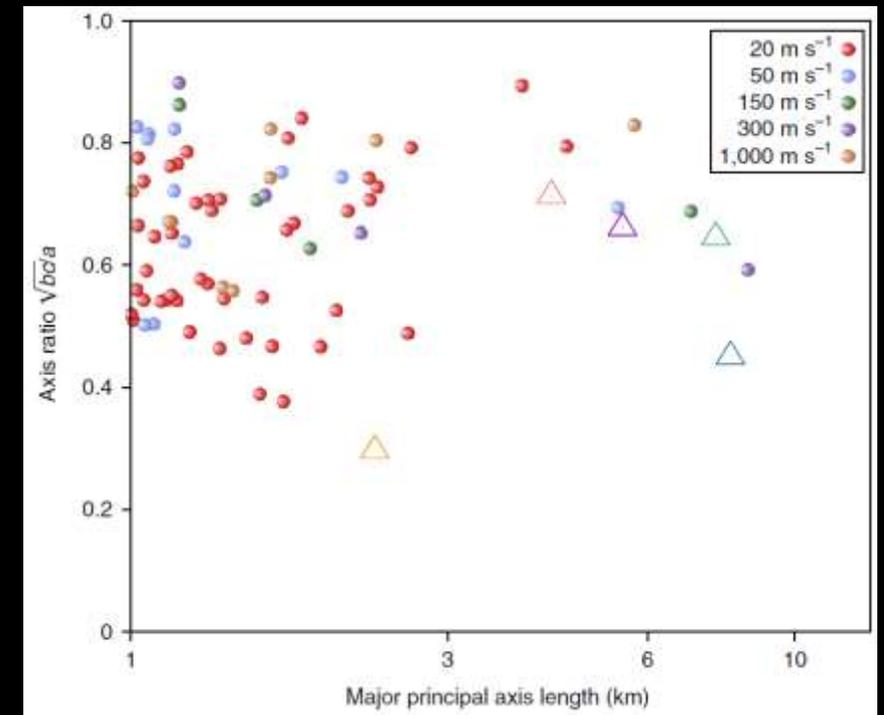
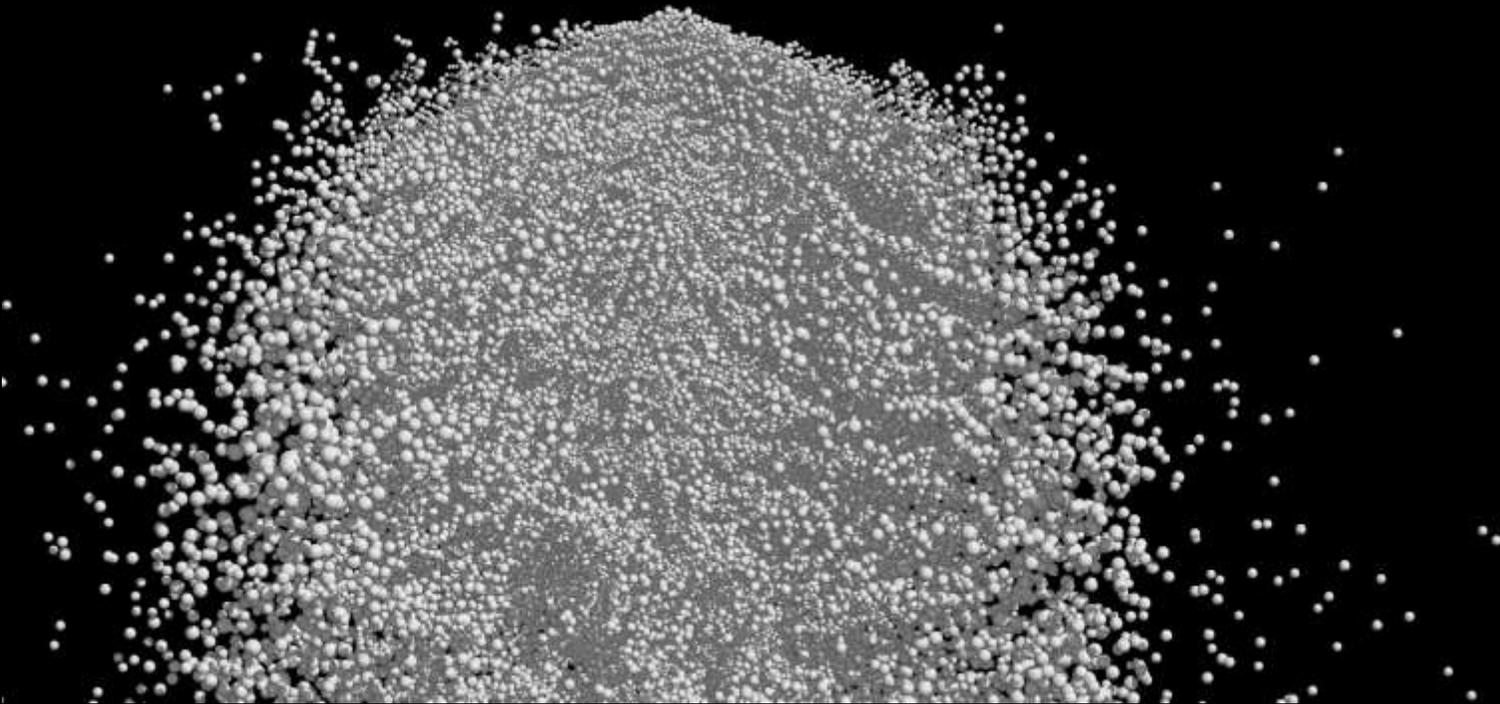


103P/Hartley 2 - 2010



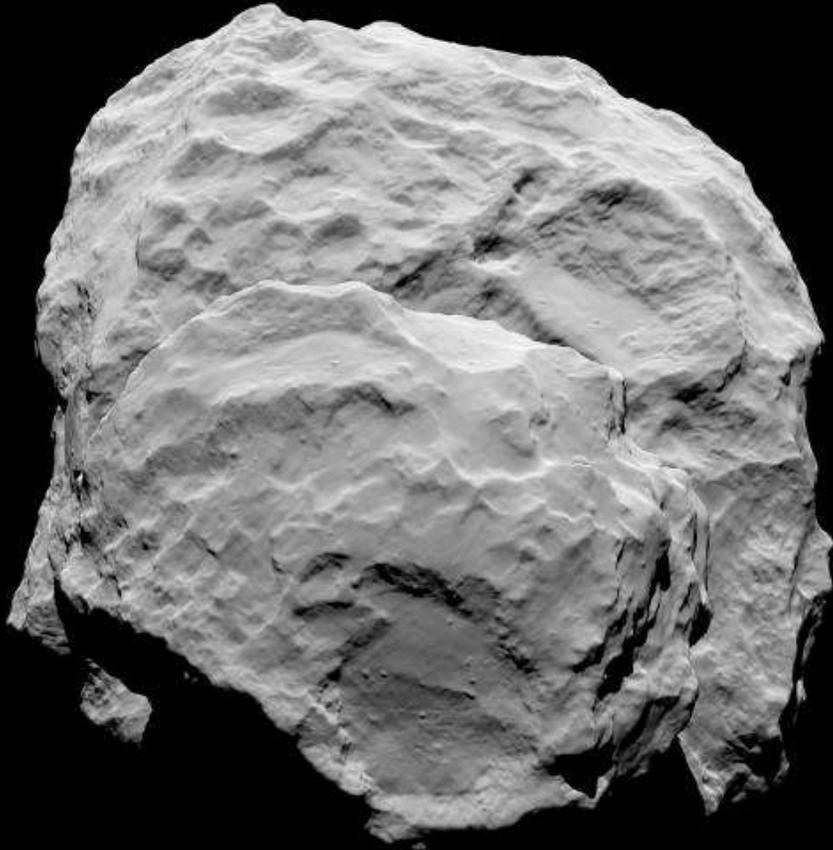
25143 Itokawa

Formation of bi-lobal shapes is omnipresent.



left: 150 m/s impact
general: no significant temperature or
porosity increase

Bulk density as low and porosity as high as expected.



SPC (v1.7) shape model

1. Volume Measurement:

- OSIRIS shape models
- $V = 18.56 \pm 0.02 \text{ km}^3$ (Preusker et al., 2017)

2. Mass Determination:

- Rosetta Radio Science Instrument (RSI)
- Measure mass (grav. pull) from Doppler shift
- $M = (9,982 \pm 3) \times 10^9 \text{ kg}$ (Pätzold et al., 2016)

3. Density & Porosity

- $\rho = M/V = 537.8 \pm 0.7 \text{ kg/m}^3$
- $14\% \leq f_{\text{ice}} \leq 33\%$ (Rotundi et al., 2015)
- 25% metal/sulfides, 42% rock/organics, and 32% ice (mass), thus $\rho = 1820 \text{ kg/m}^3$ (Davidsson et al., 2016)
- $\psi = (71 \pm 2)\%$

Strength in the expected range, no proof for “consolidated terrain”.

Applied data:

- Overhangs, collapsed structures, boulders, cliffs, and Philae’s footprint
- Images and shape models

Tensile strength σ_T :

< 150 Pa (collapsed structures)

3-15 Pa (overhangs)

Shear strength σ_S :

> 30 Pa (Hathor cliffs)

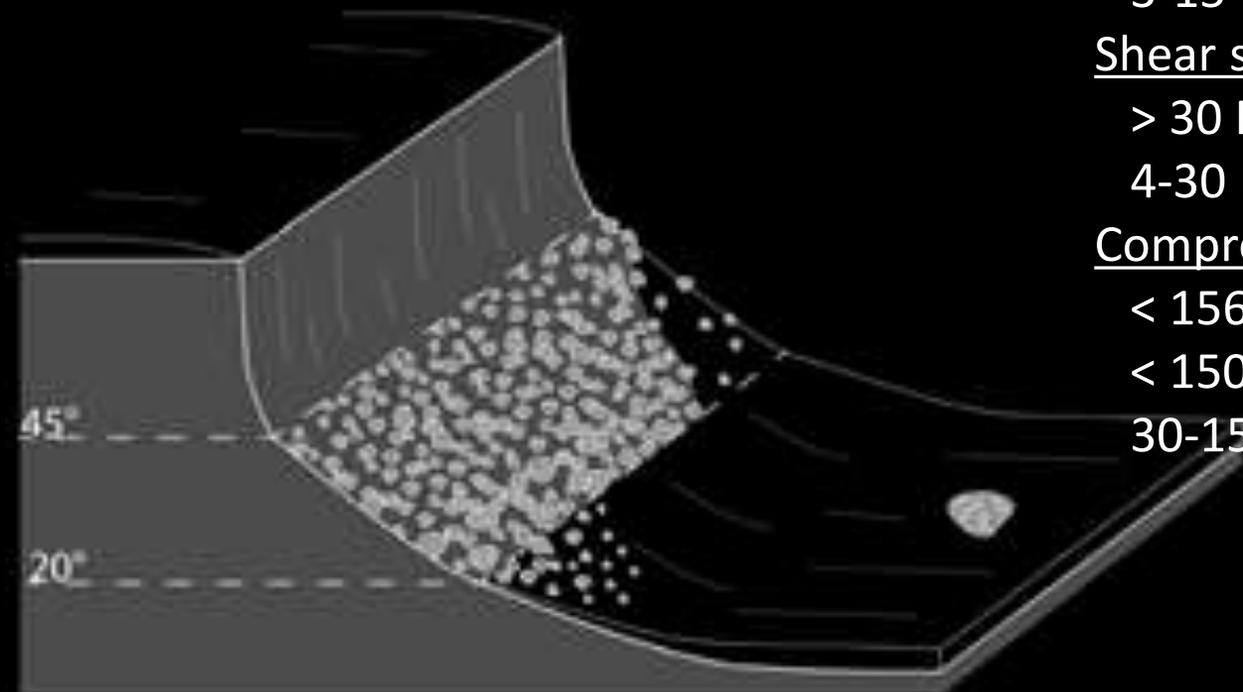
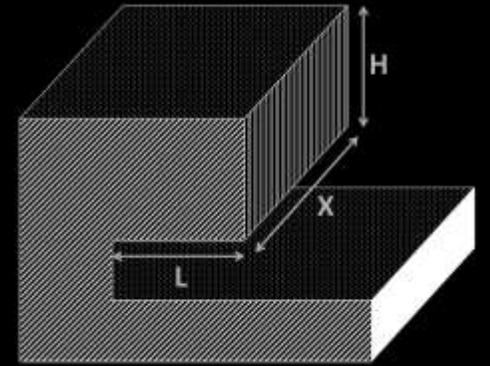
4-30 Pa (fine surface materials and boulders)

Compressive strength σ_C :

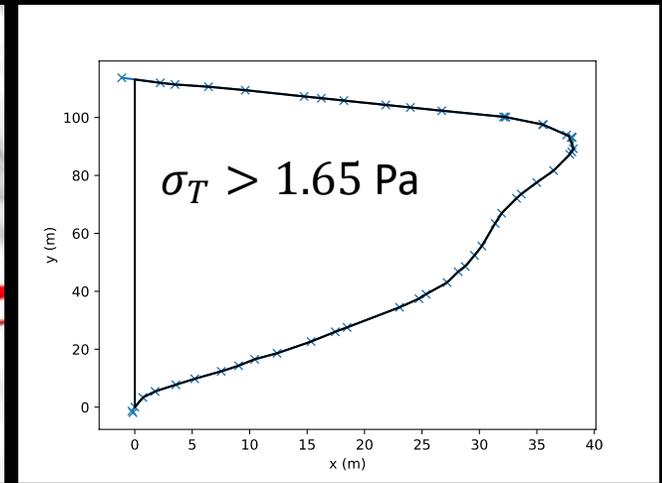
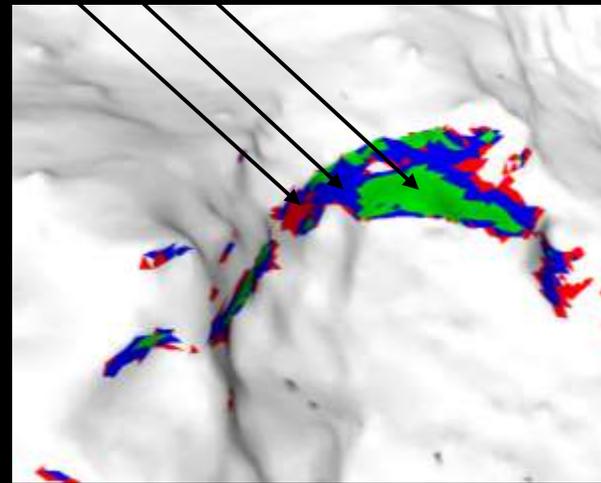
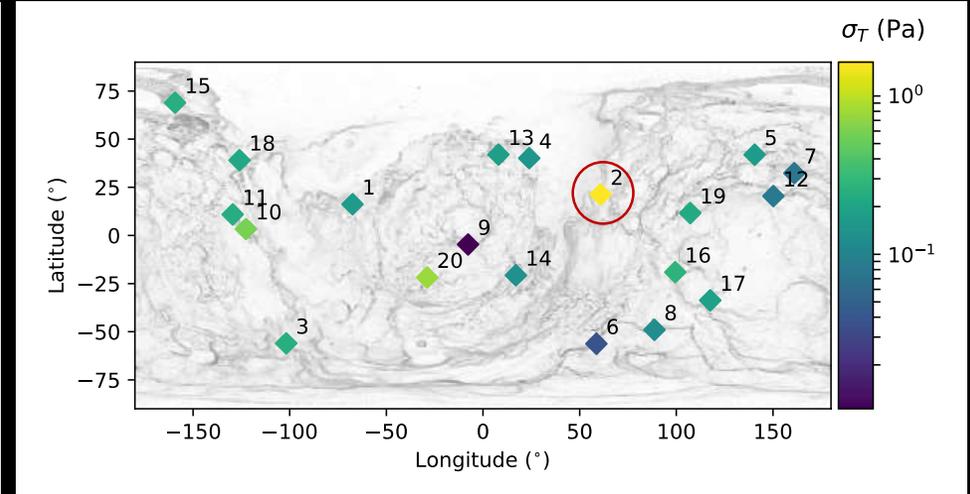
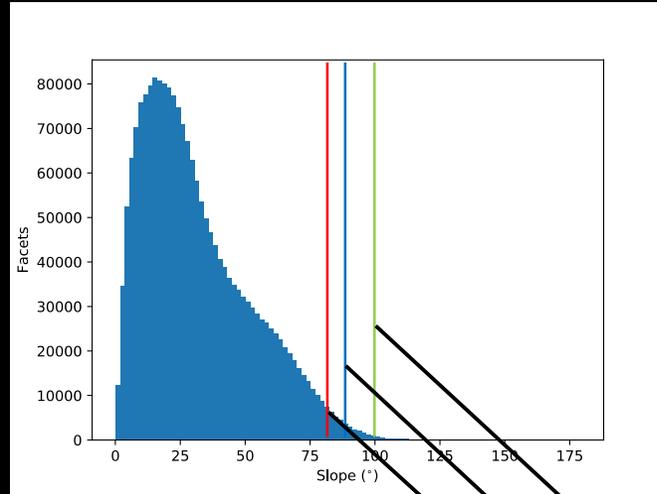
< 15600 Pa (fine surface materials)

< 1500 Pa (collapsed structures)

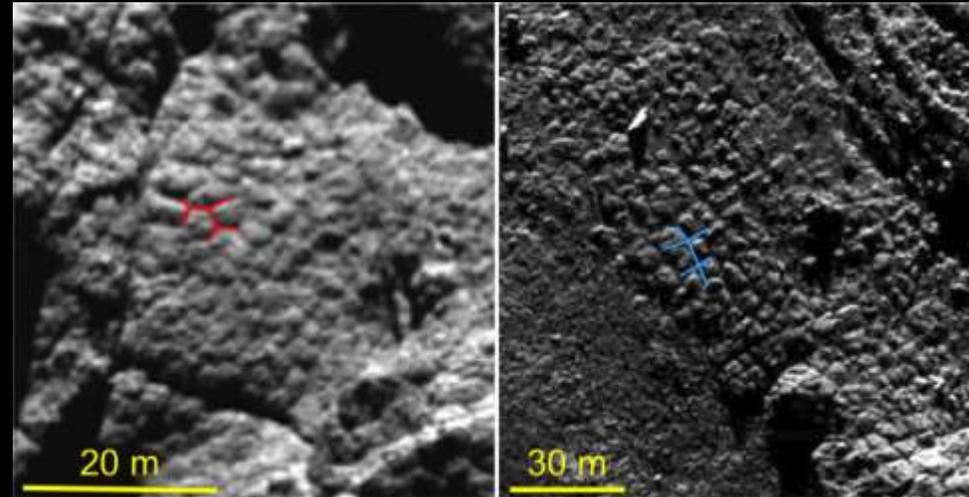
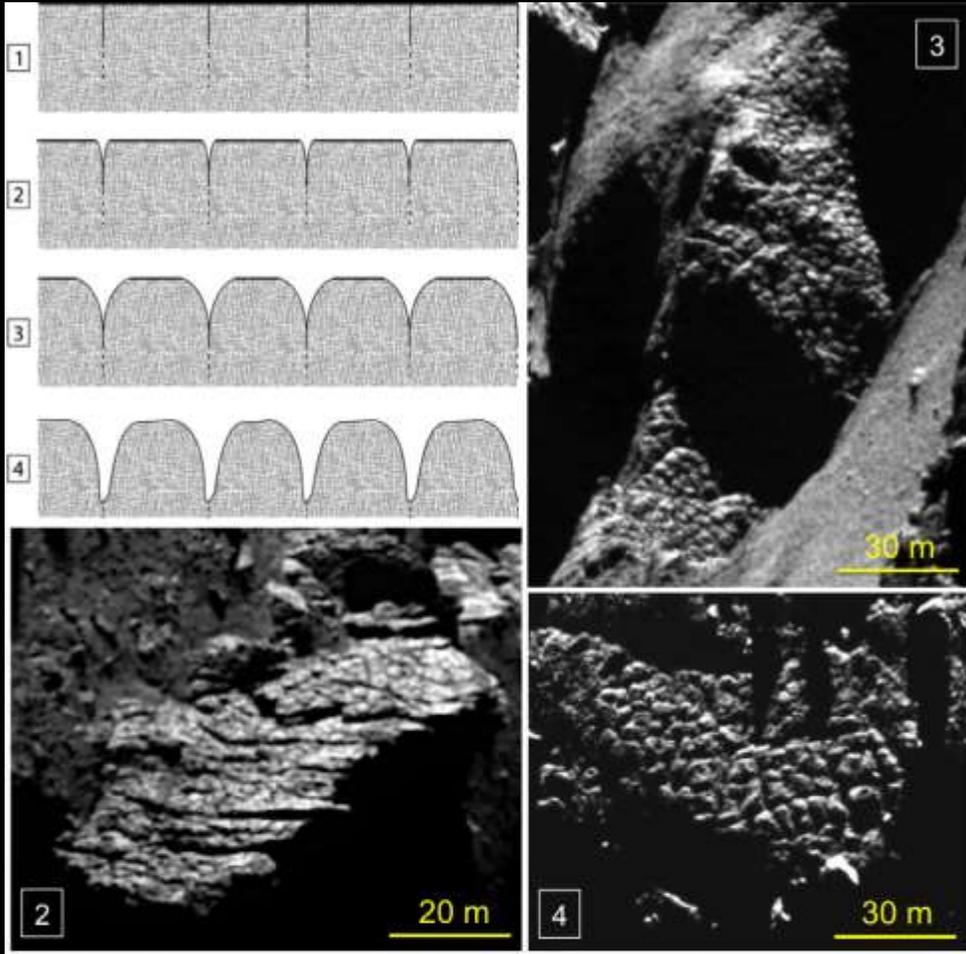
30-150 Pa (overhangs)



Strength in the expected range, no proof for “consolidated terrain”.



Formation and evolution of cracks debated.



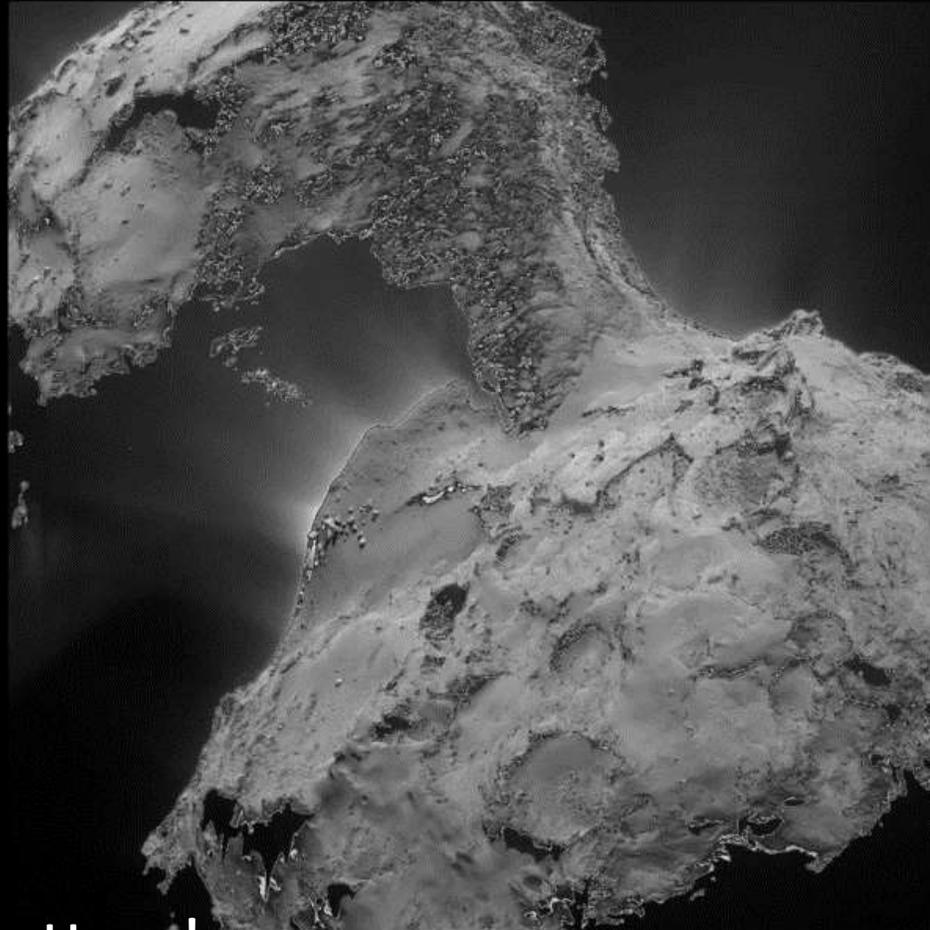
“Their formation is consistent with the diurnal or seasonal temperature variations in a hard (MPa) and consolidated sintered layer of water ice, located a few centimeters below the surface.”

What is the required material strength to propagate cracks on the meter scale?

Activity has many faces:
diffuse (?), jet-like, transient, outbursts, sunset.



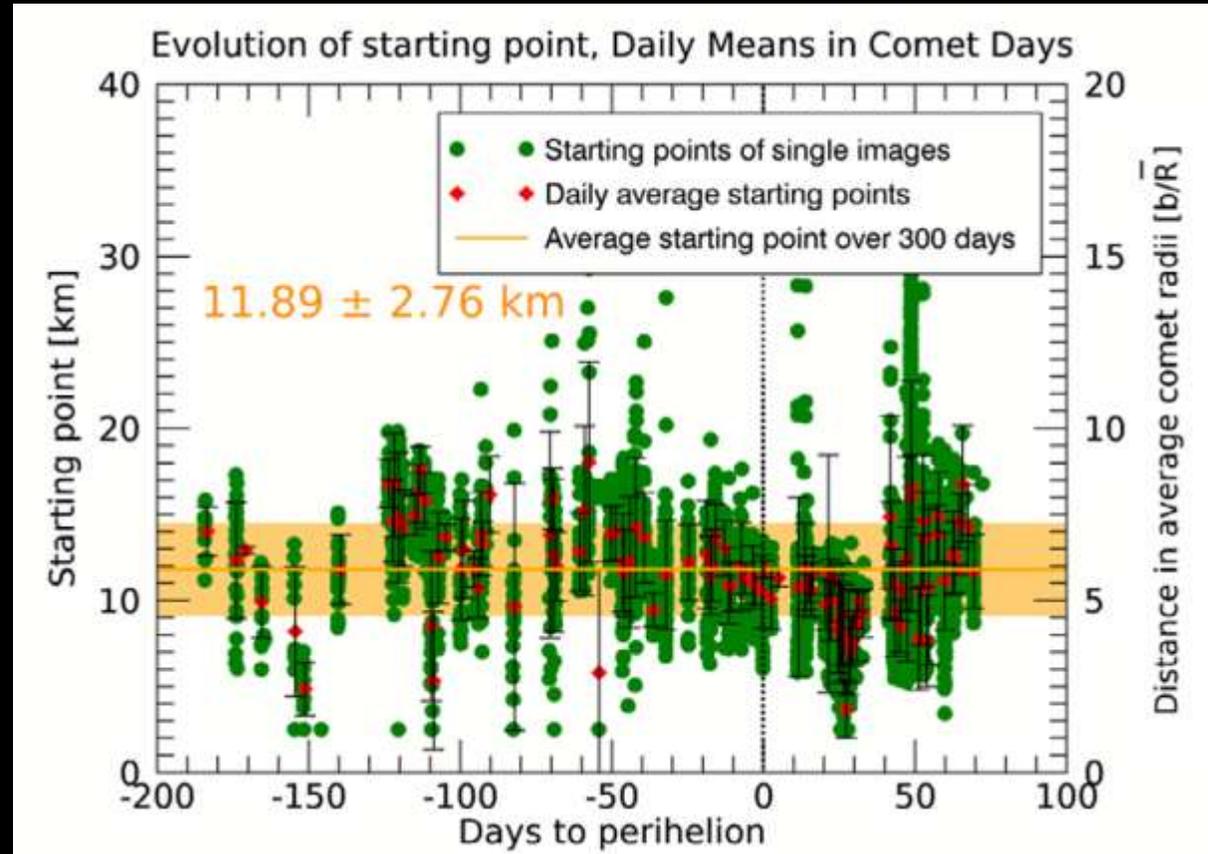
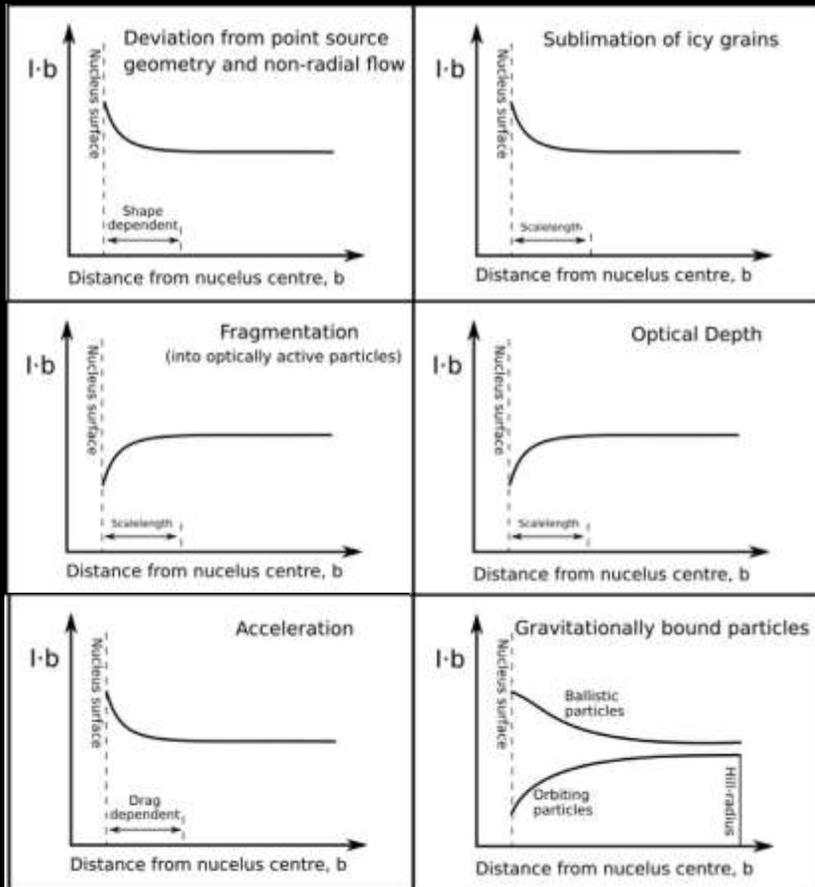
Activity has many faces:
diffuse (?), jet-like, transient, outbursts, sunset.



Observation geometry matters!

Shi et al., 2018

Processes of inner dust-coma (below 12 km) complex, thus poorly understood.

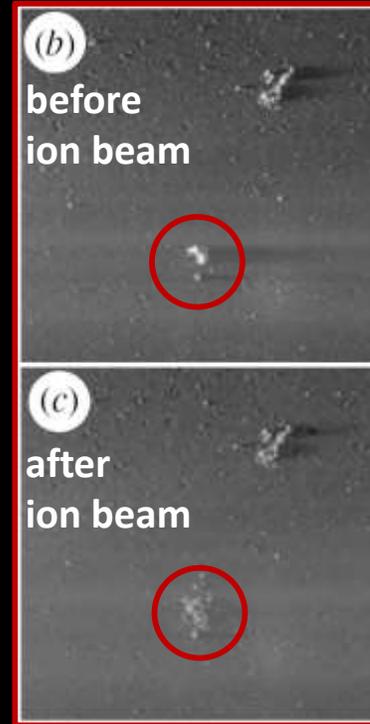
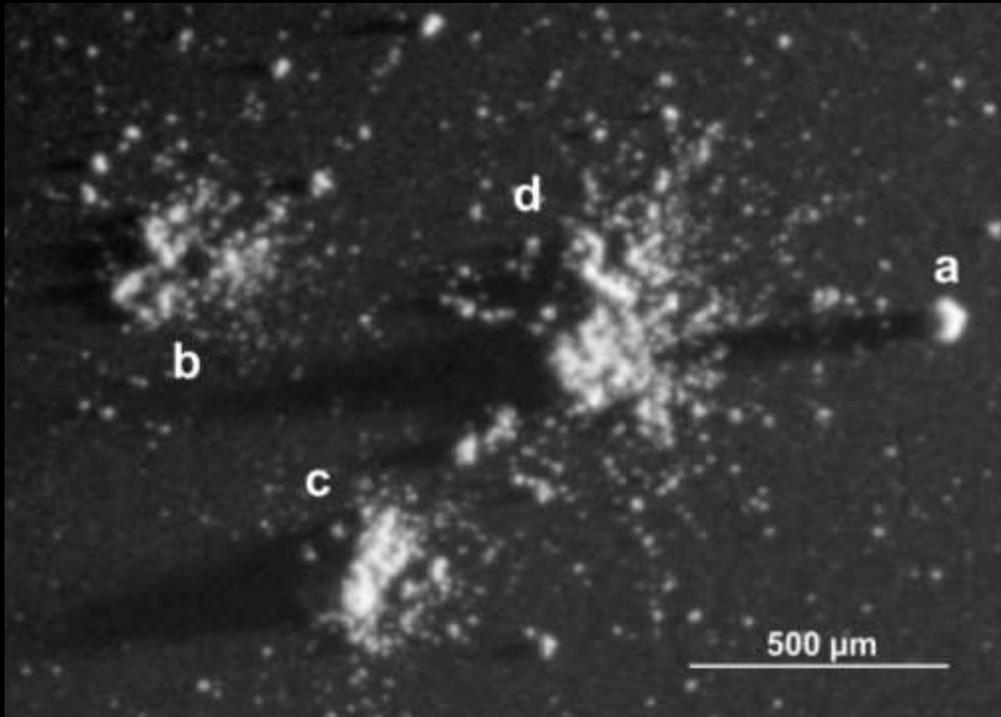


$$f(b; u, v, w, z) = u/b^v + wb + z$$

The brightness (“dust density”) as a function of impact parameter (“projected radius”) follows a “1/r” profile above 11.89 km.

Dust properties are consistent with agglomeration.

Langevin et al., 2016



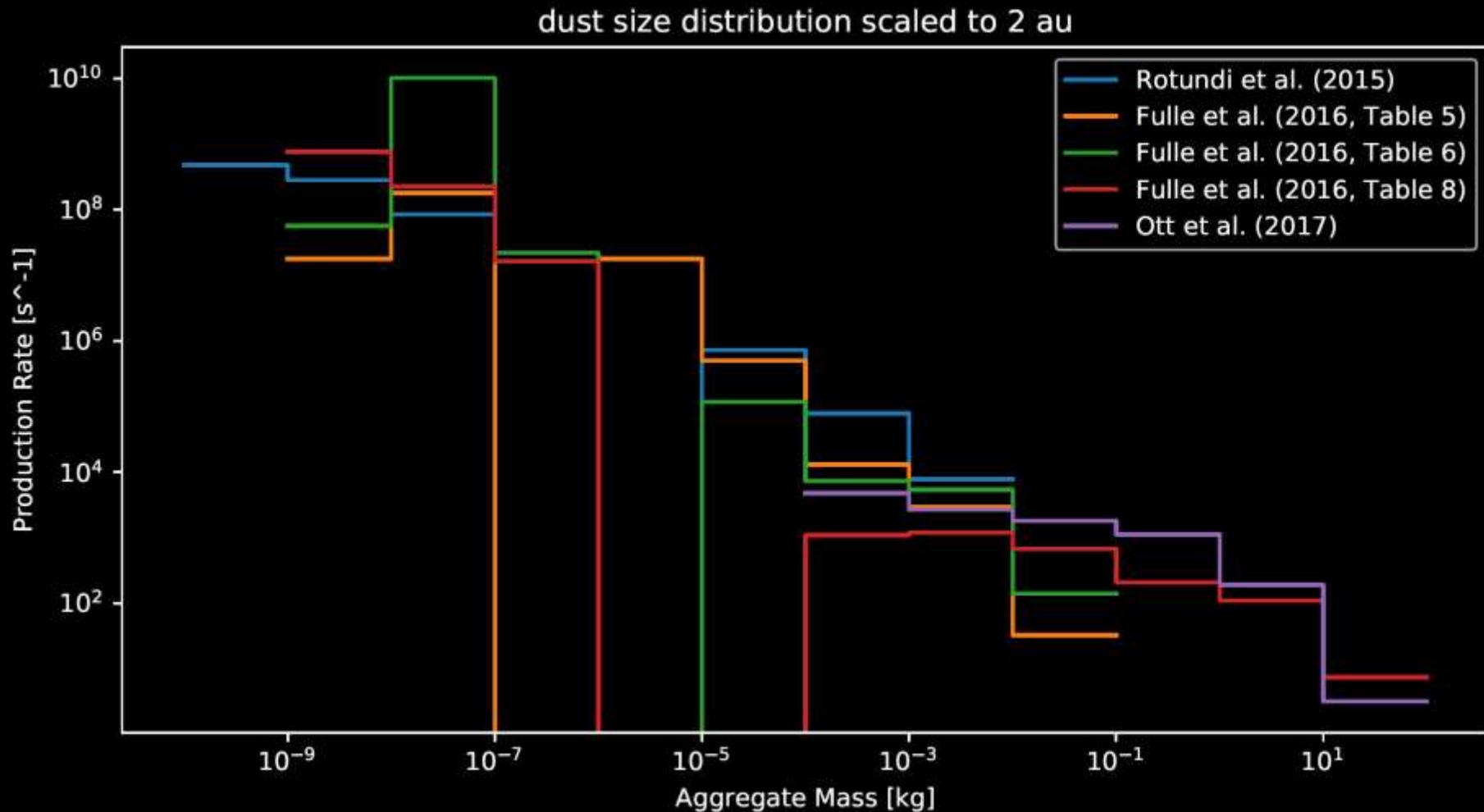
Hilchenbach et al., 2017

COSIMA: Loose structures, can be broken up by electrostatic charging (using SIMS ion beam).

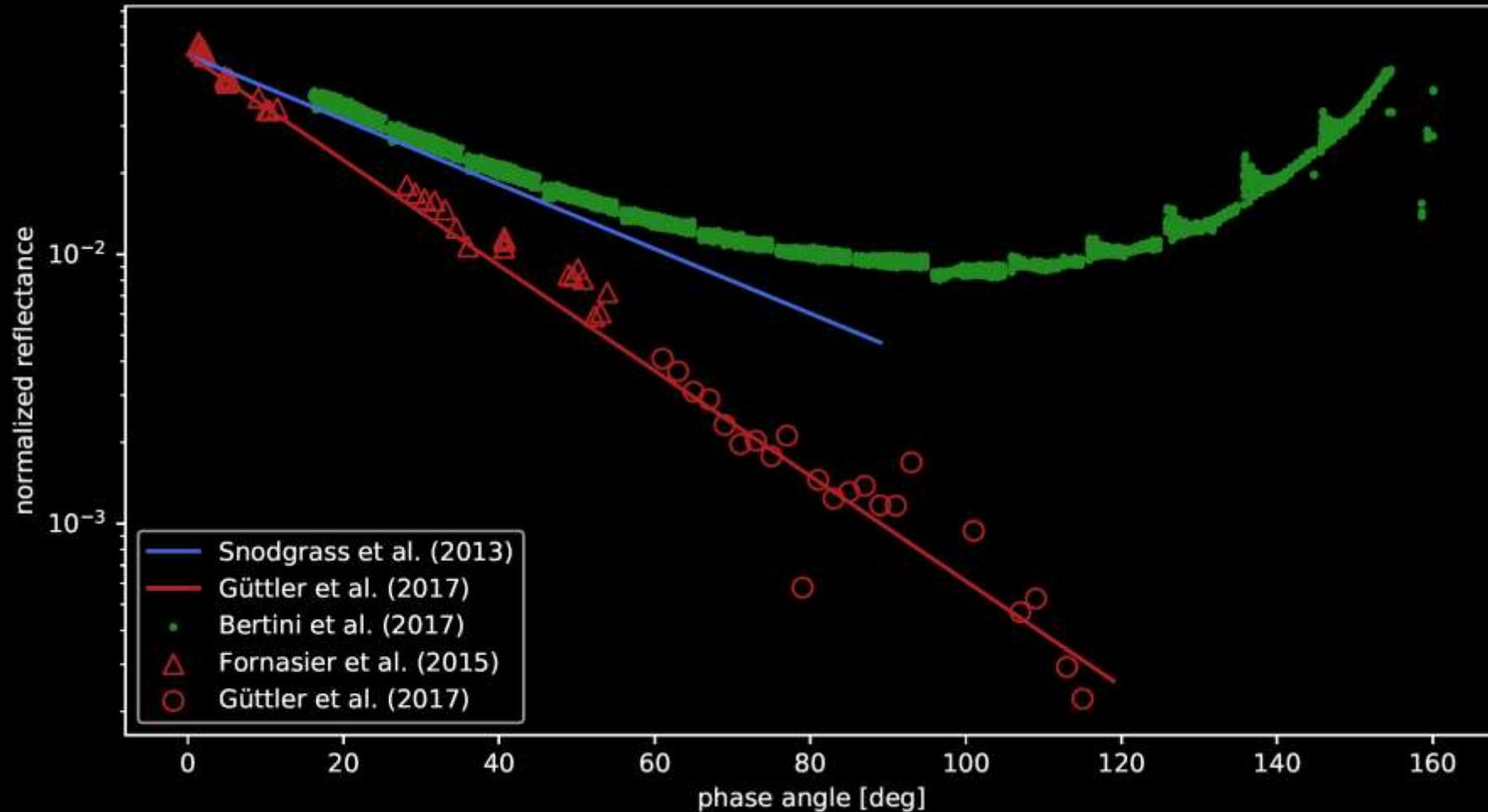
MIDAS: Smallest observed structures in the 0.1 μm range. Consistent with several model approaches (VIRTIS, MIRO, etc.)

Mannel et al., in prep.

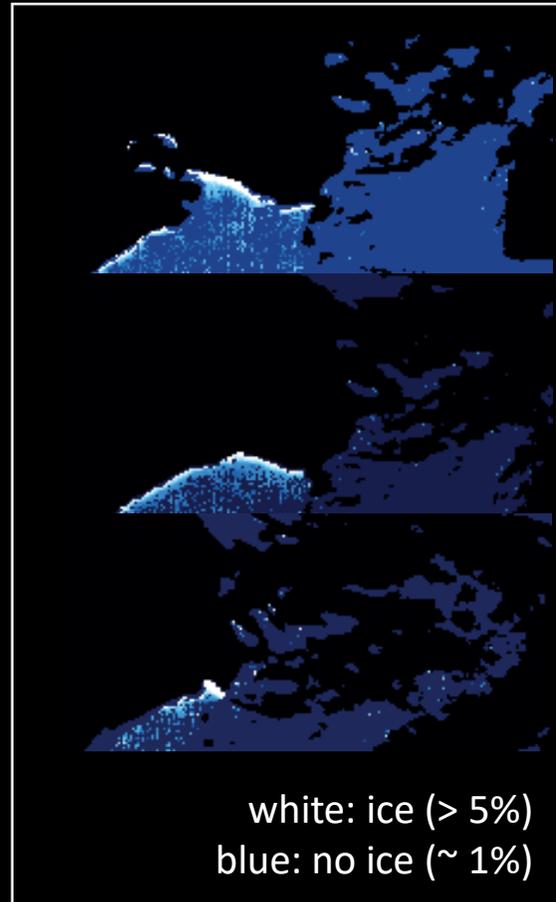
The dust size distribution is variable and not fully understood.



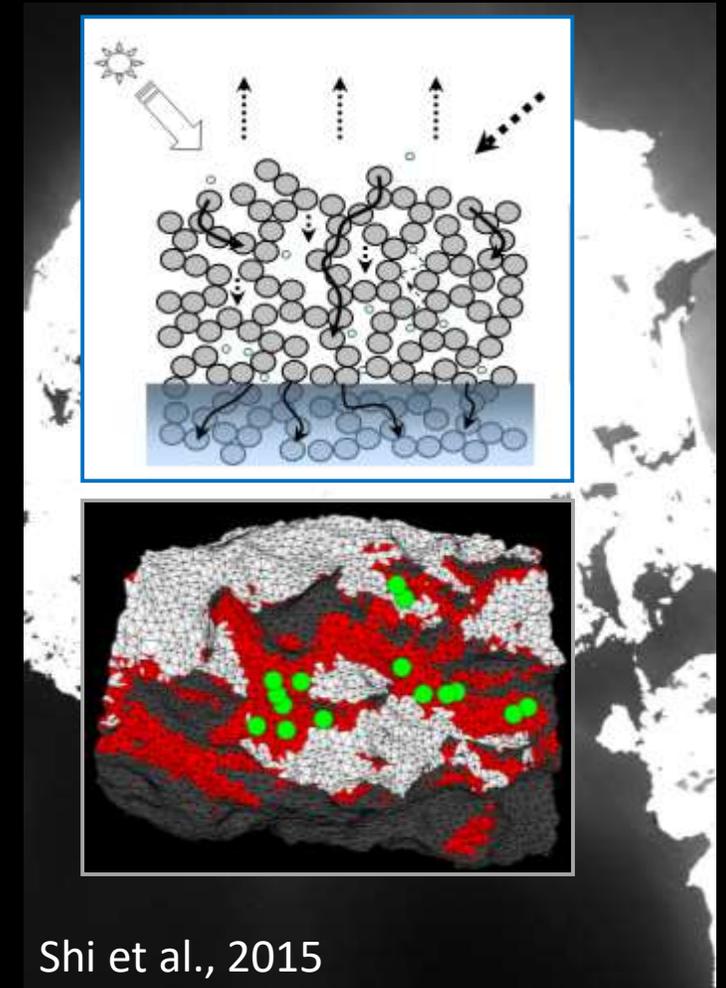
The dust phase function affects most size measurements.



Many sub-surface processes are understood but model dependent.

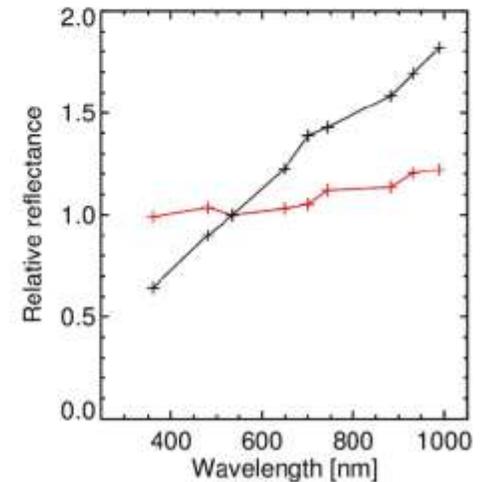
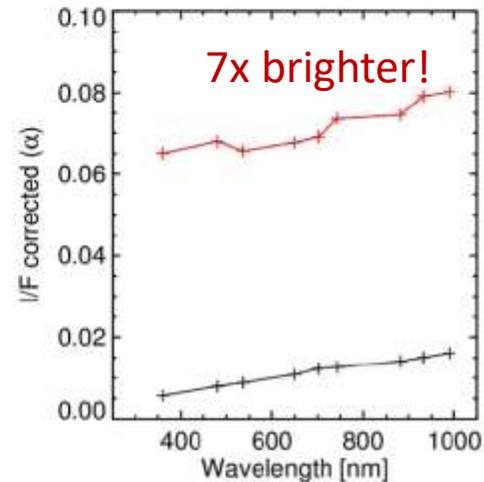
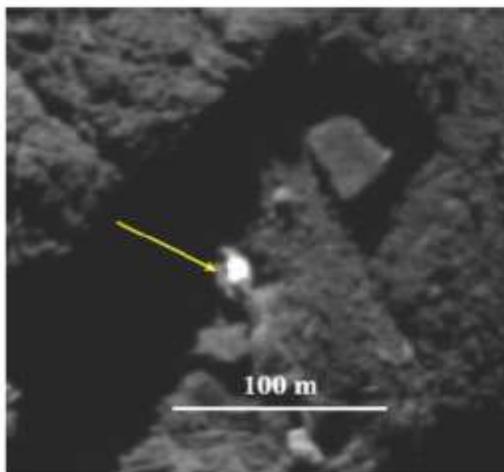
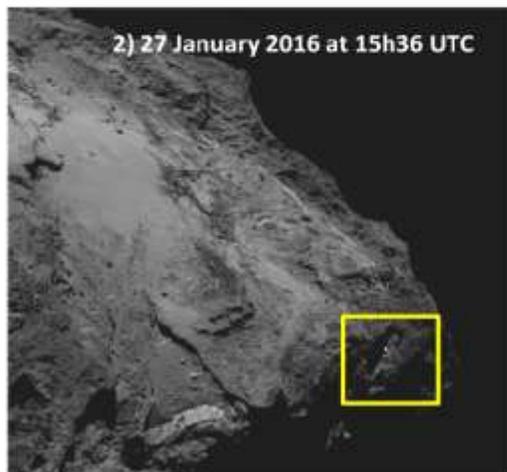
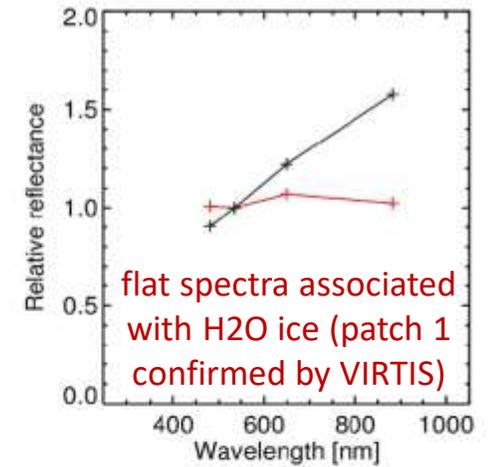
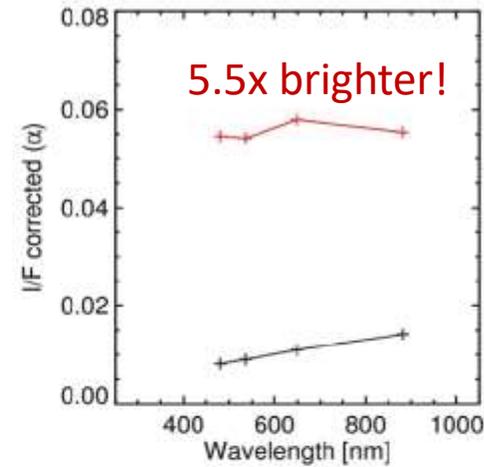
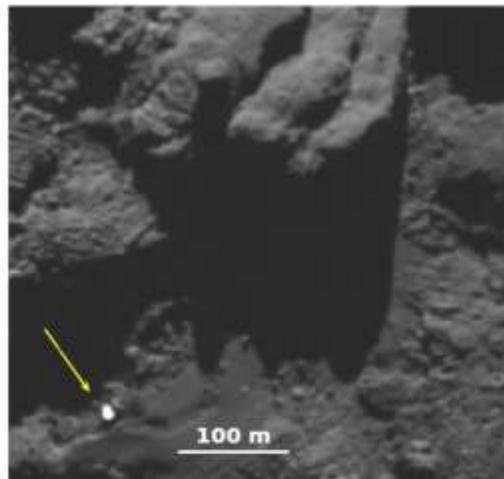


De Sanctis et al., 2015

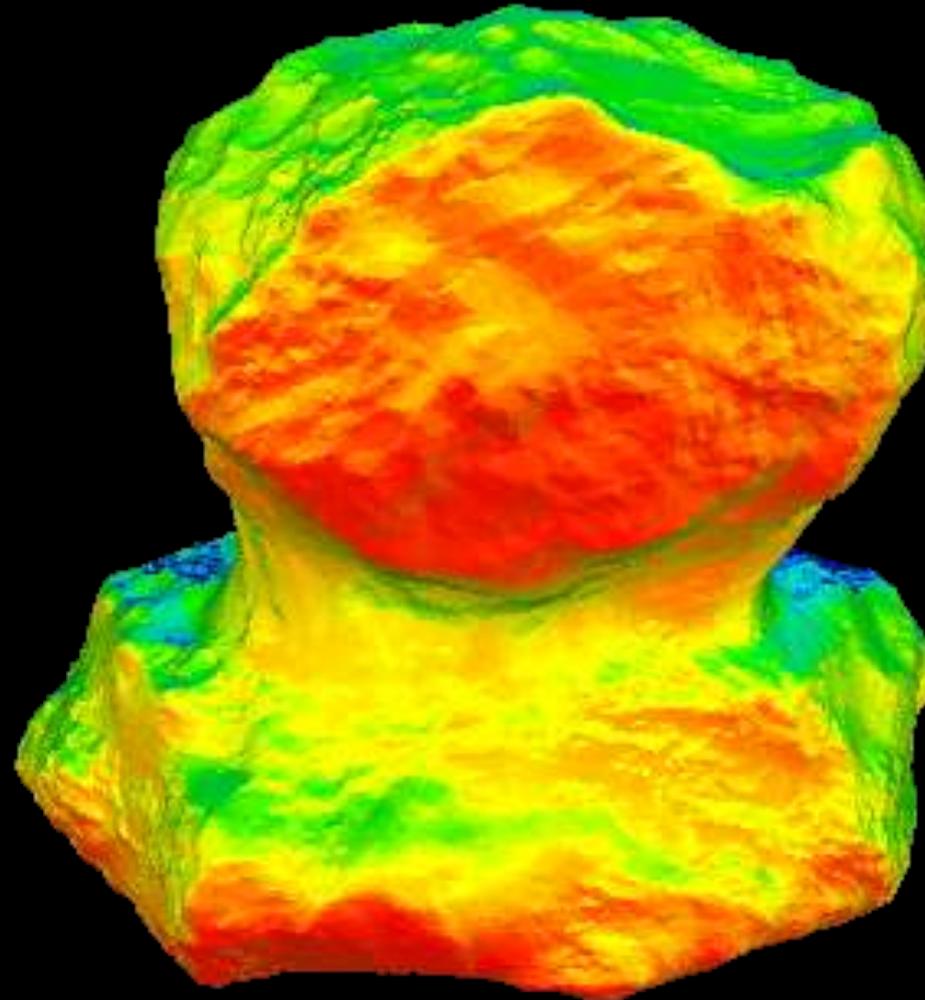
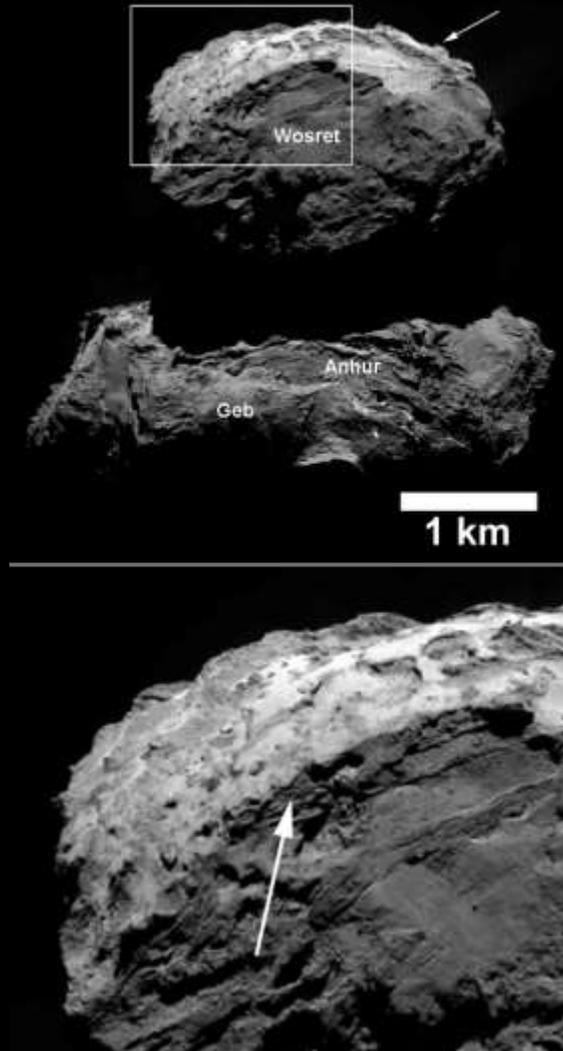


Shi et al., 2015

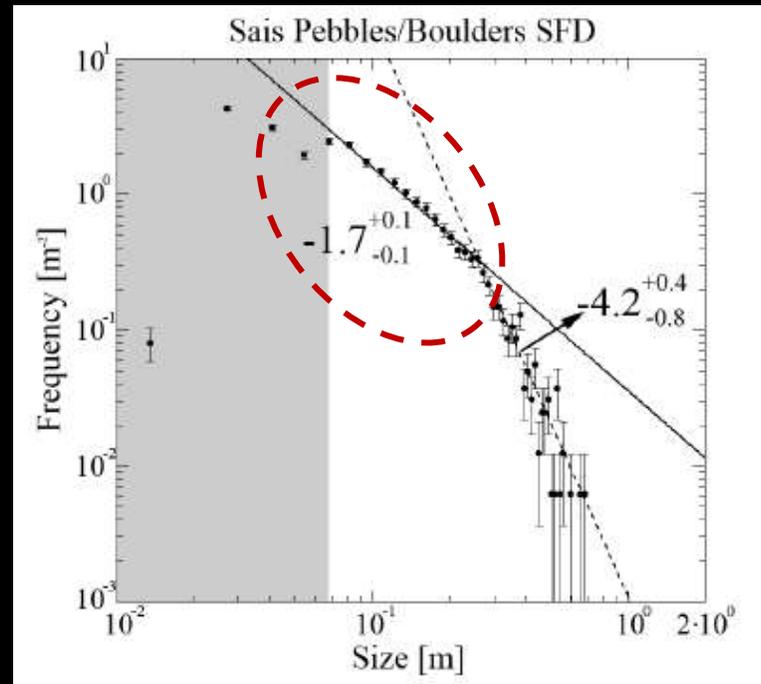
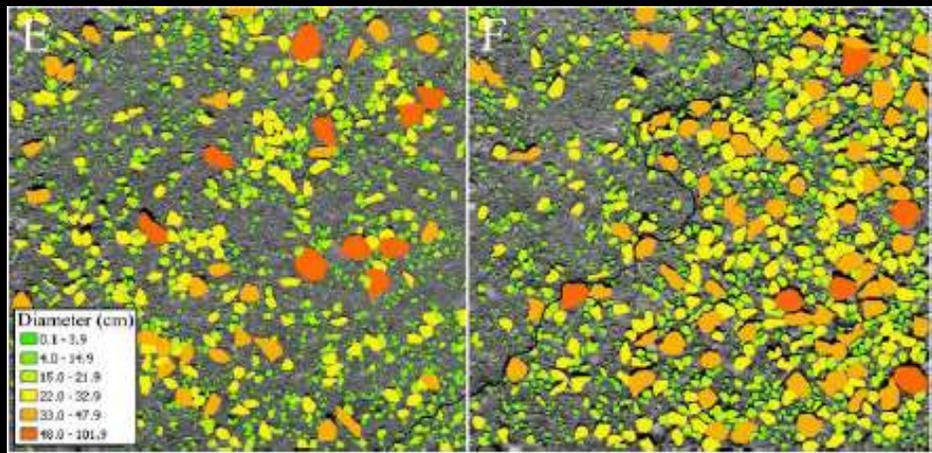
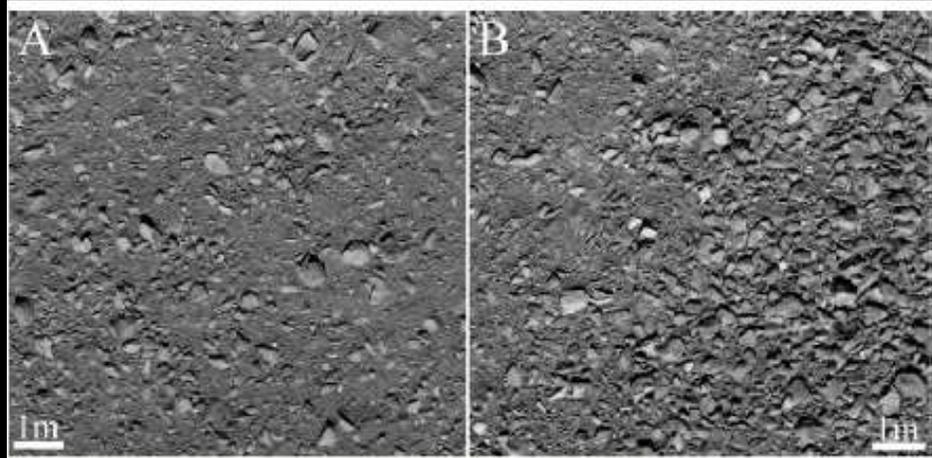
Bright patches indicating inhomogeneity?



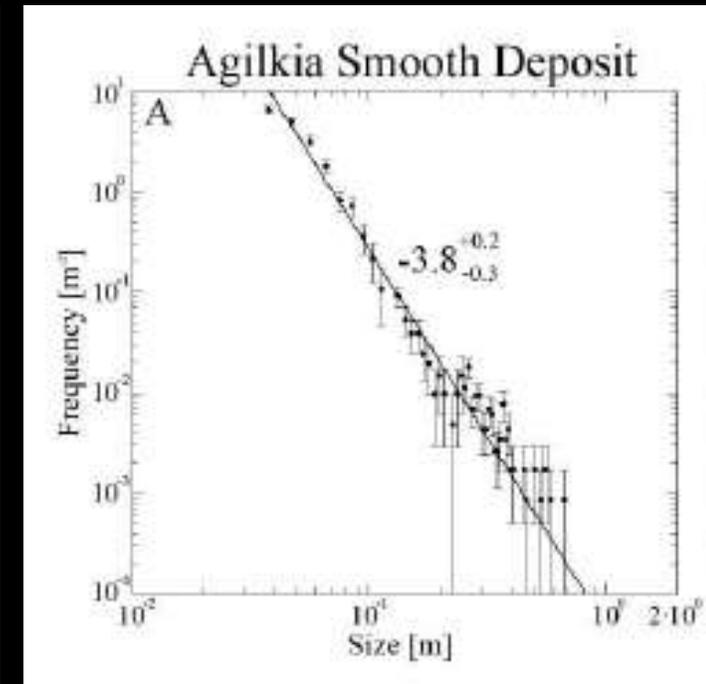
Back fall is much more important than previously acknowledged.



Which material is falling back? Size segregation?

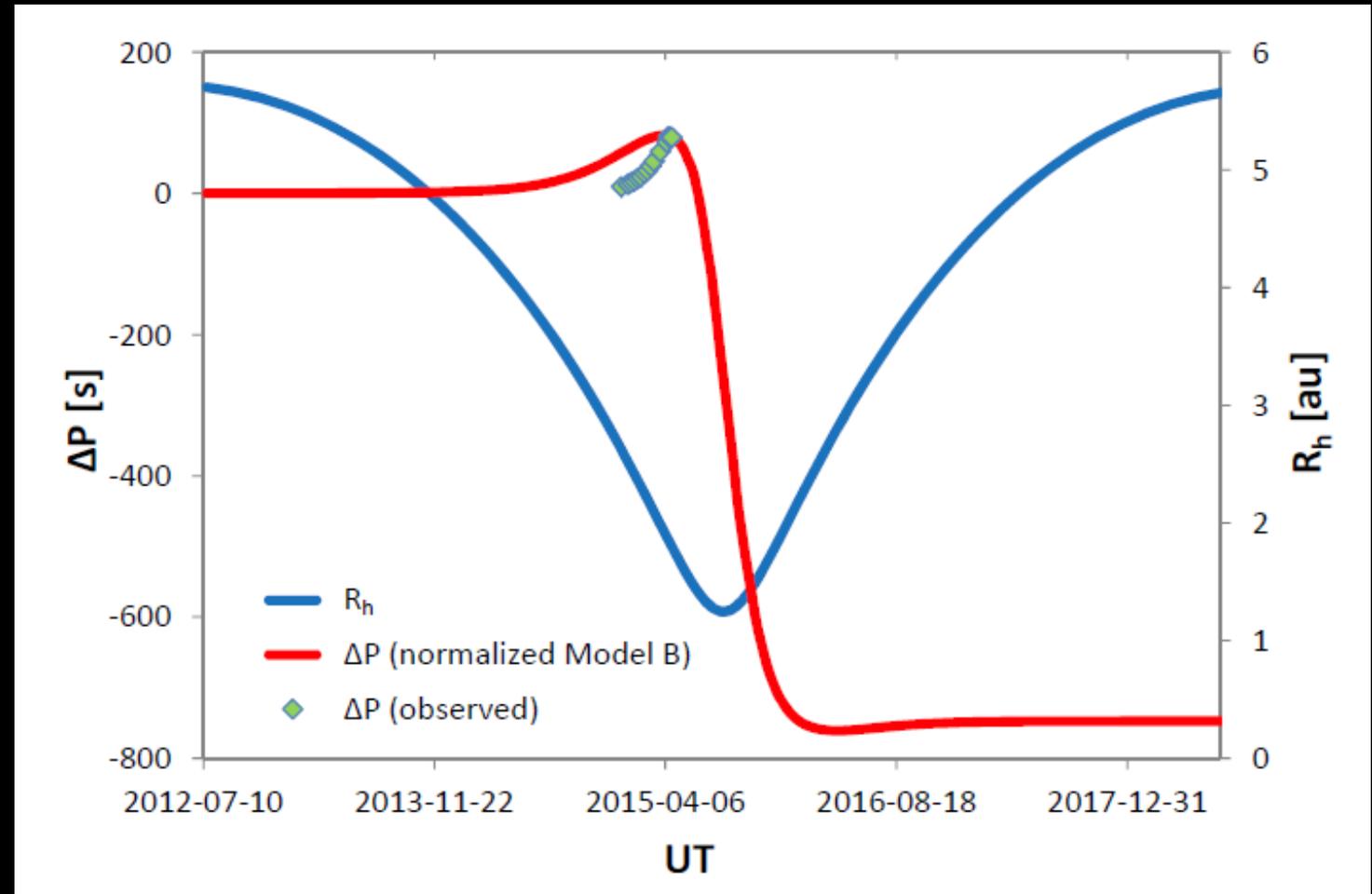
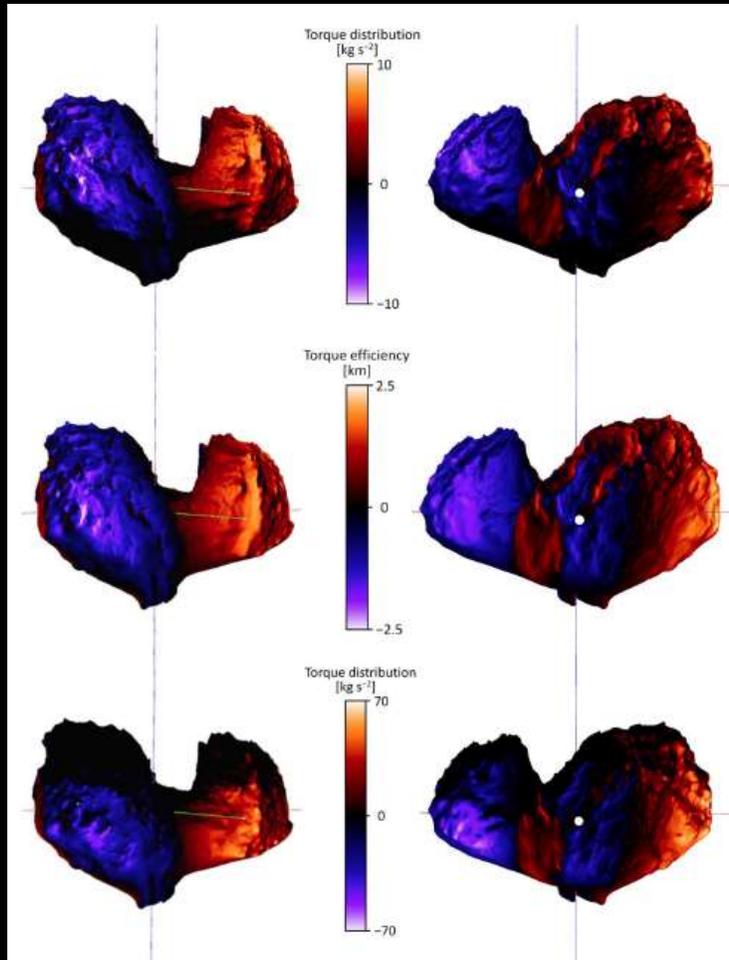


Rosetta landing site Sais

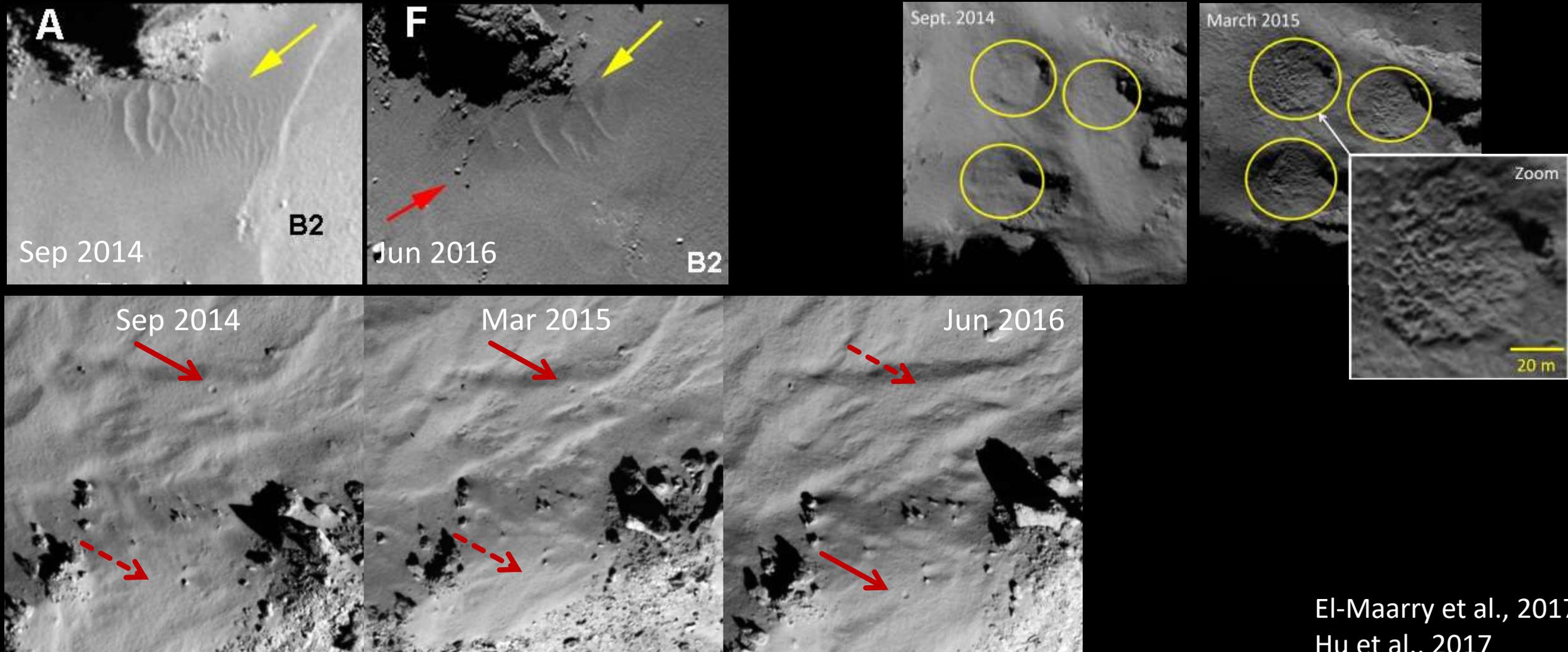


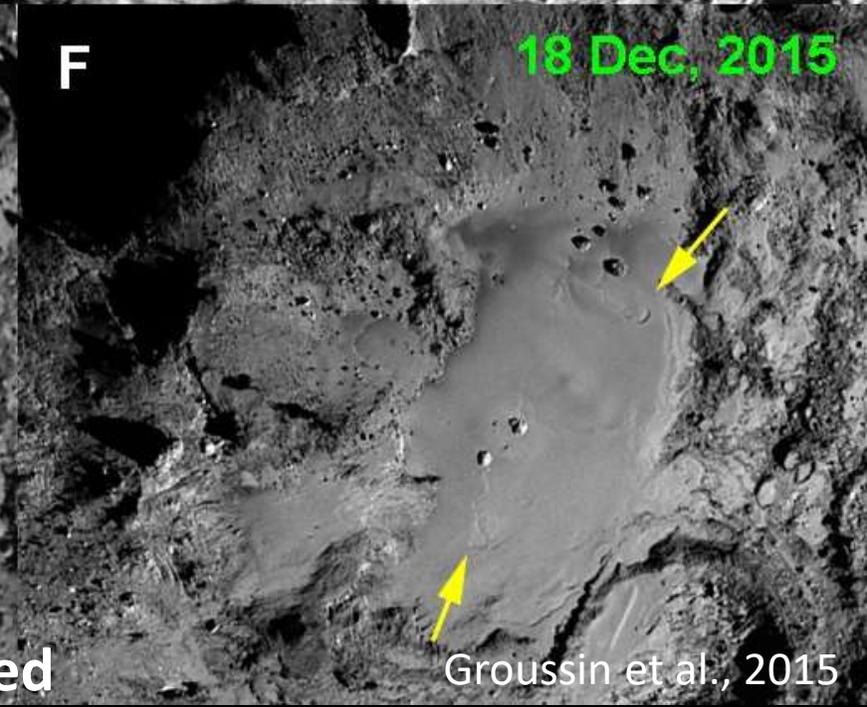
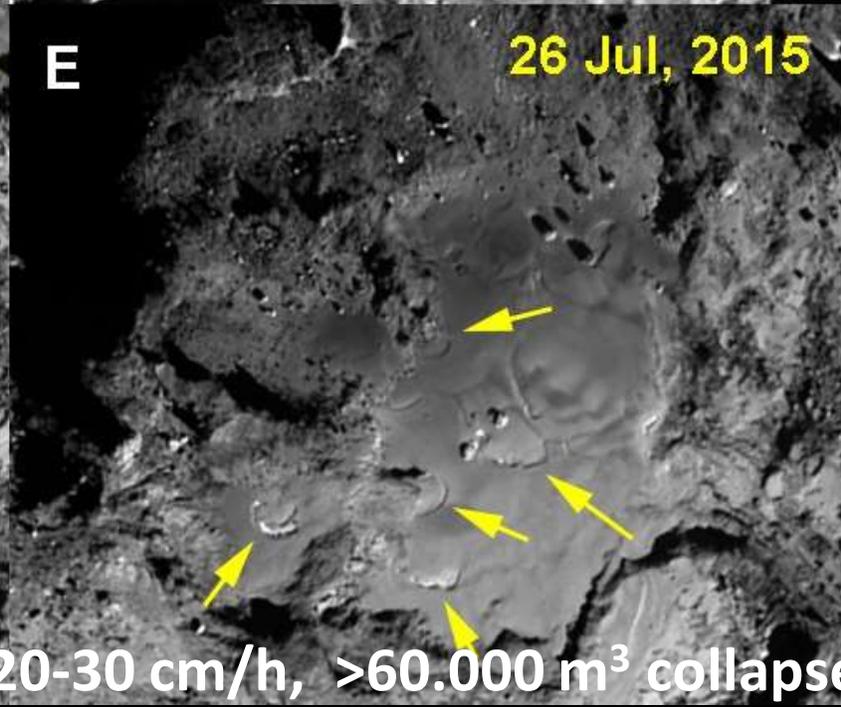
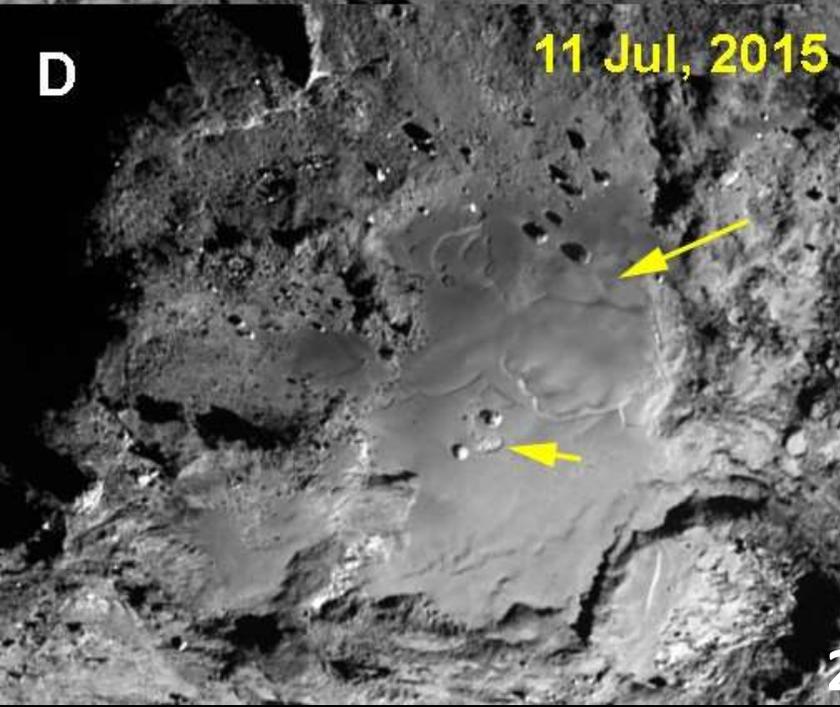
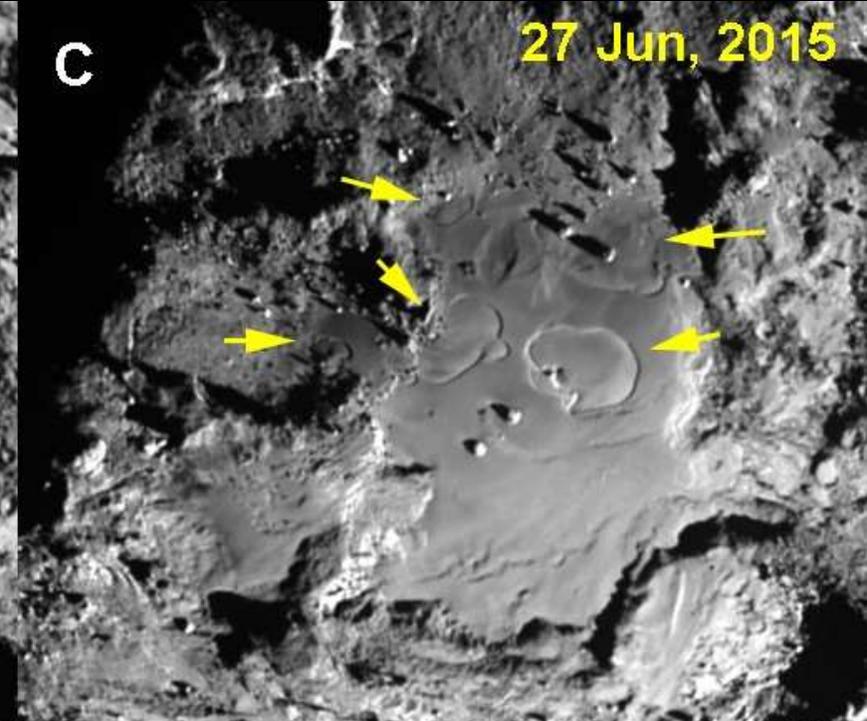
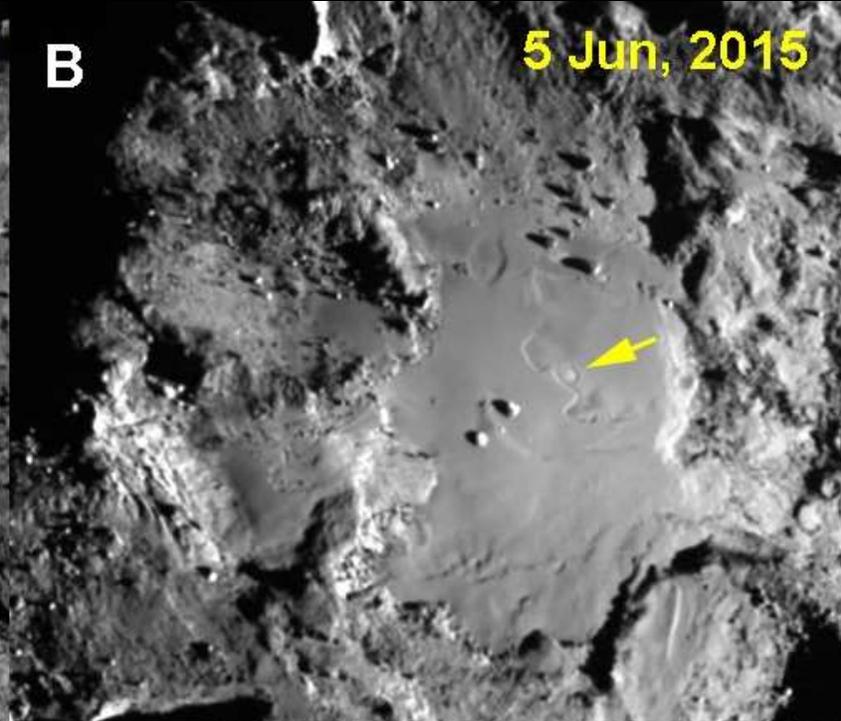
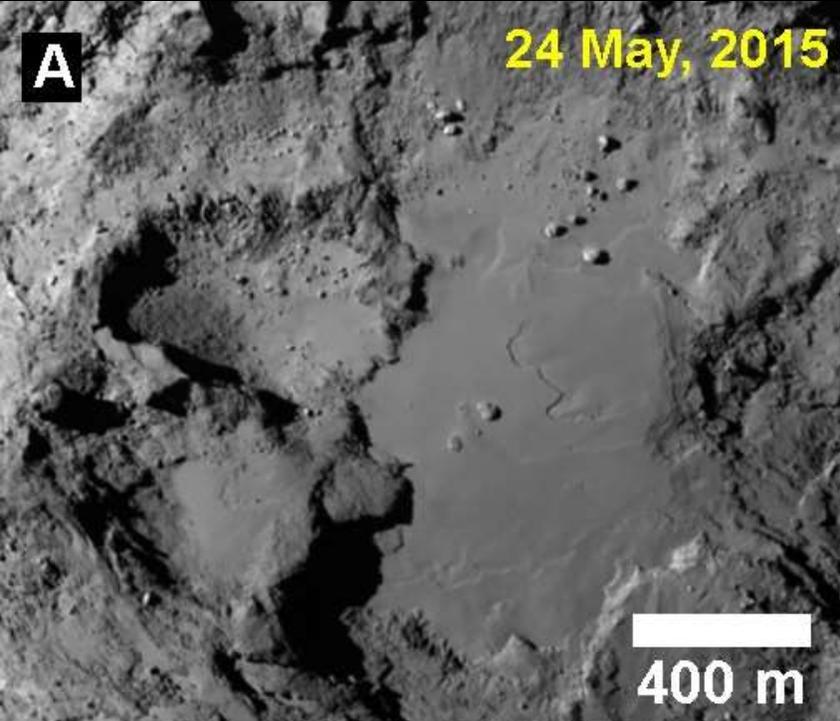
Philae touchdown site Agilkia

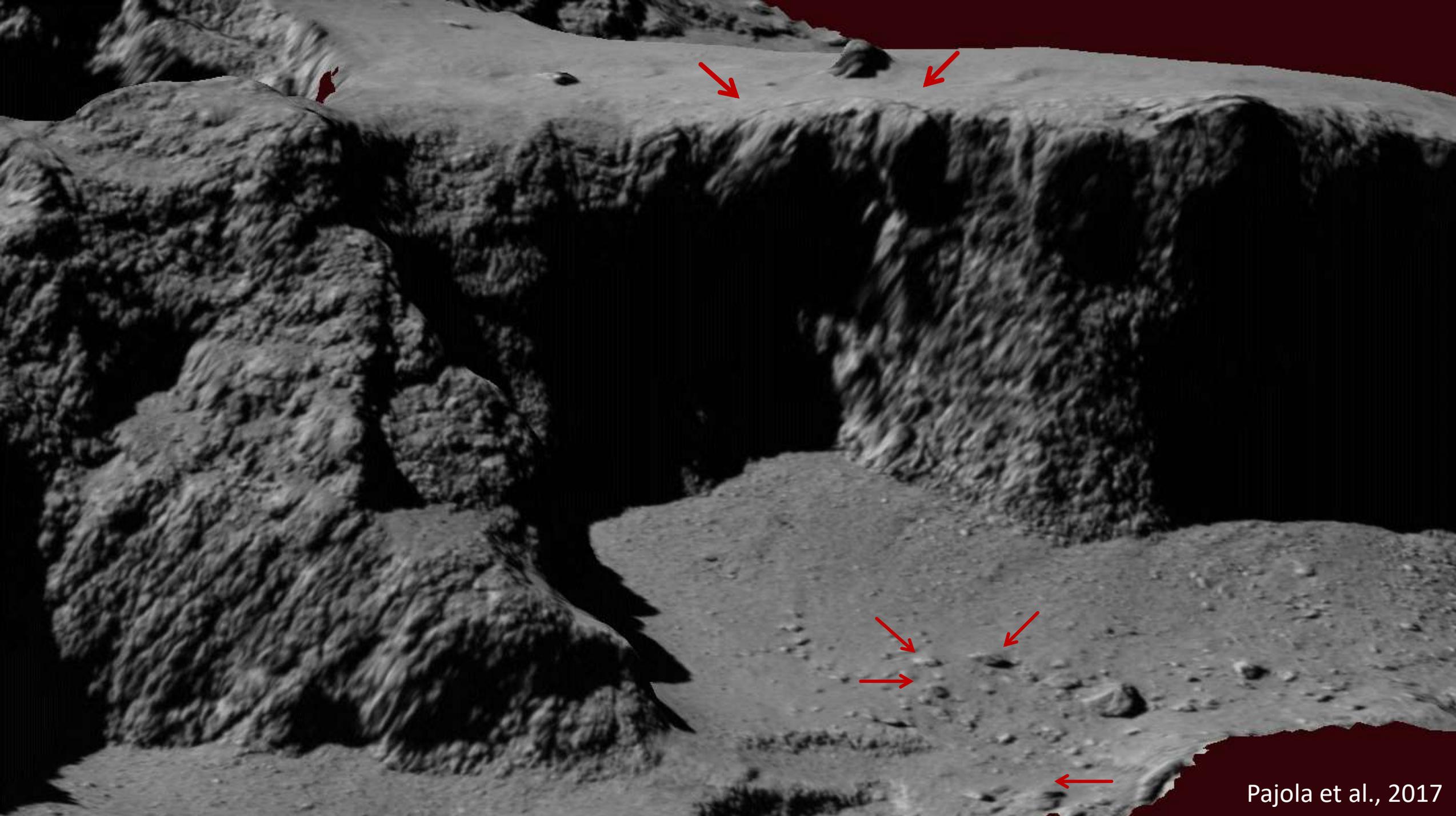
Non-gravitational forces were studied in greater detail than ever before.



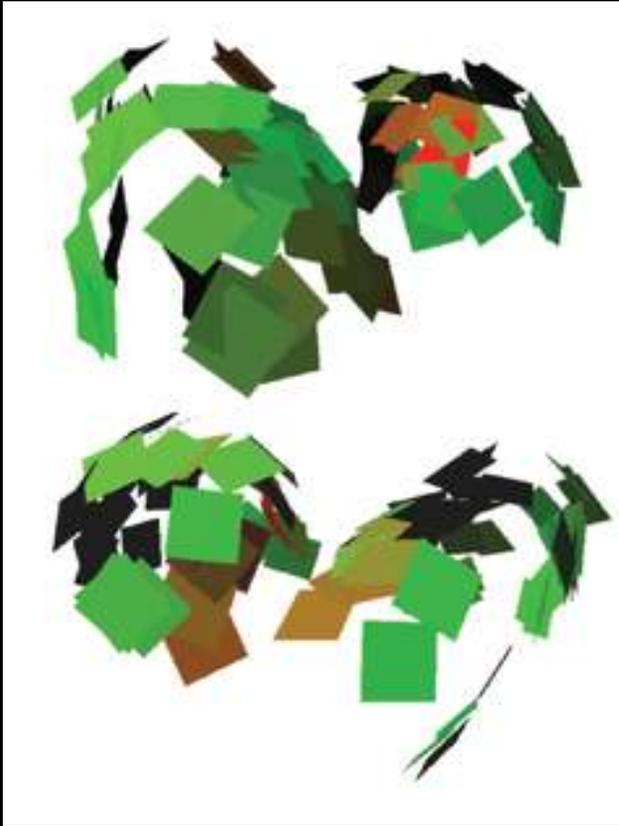
Surface changes are mostly localized and as obscure as activity itself.



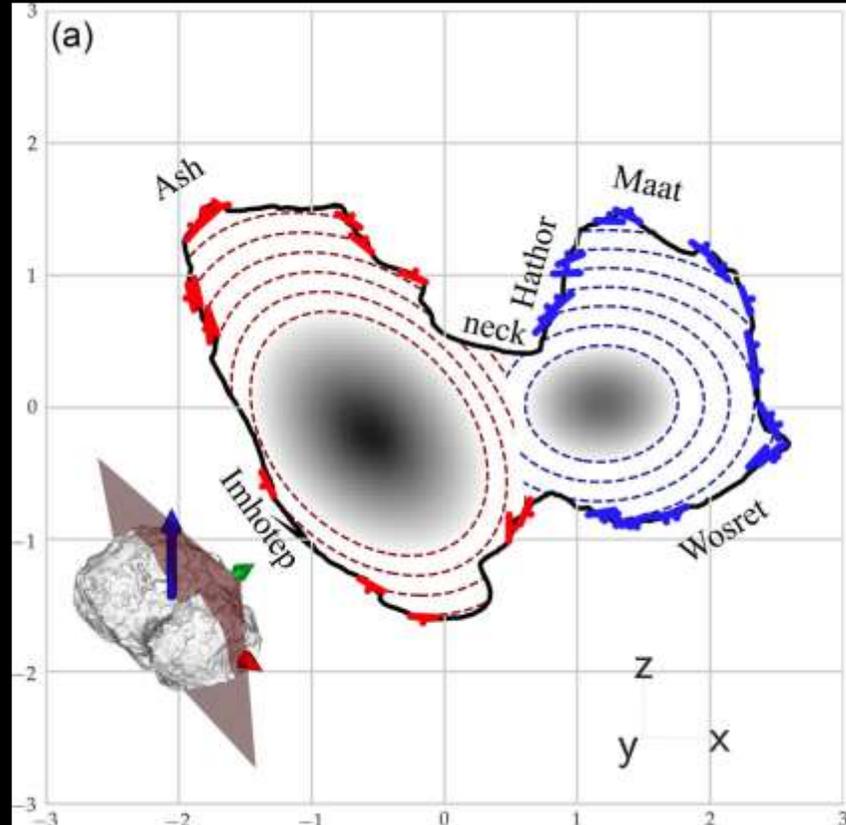




Layering is a mystery – several ideas and possible implications.



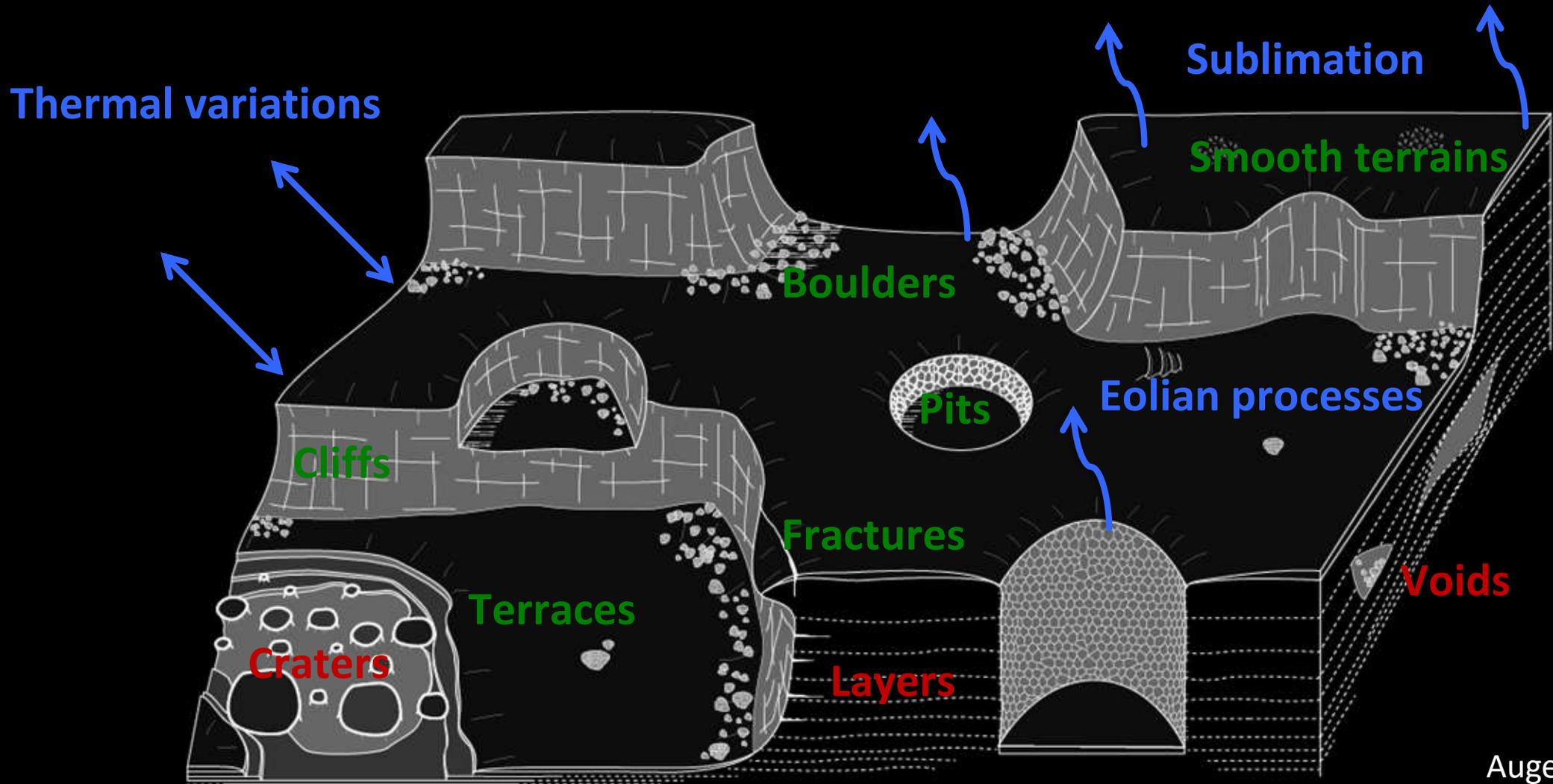
Massironi et al., 2015



Penasa et al., 2017

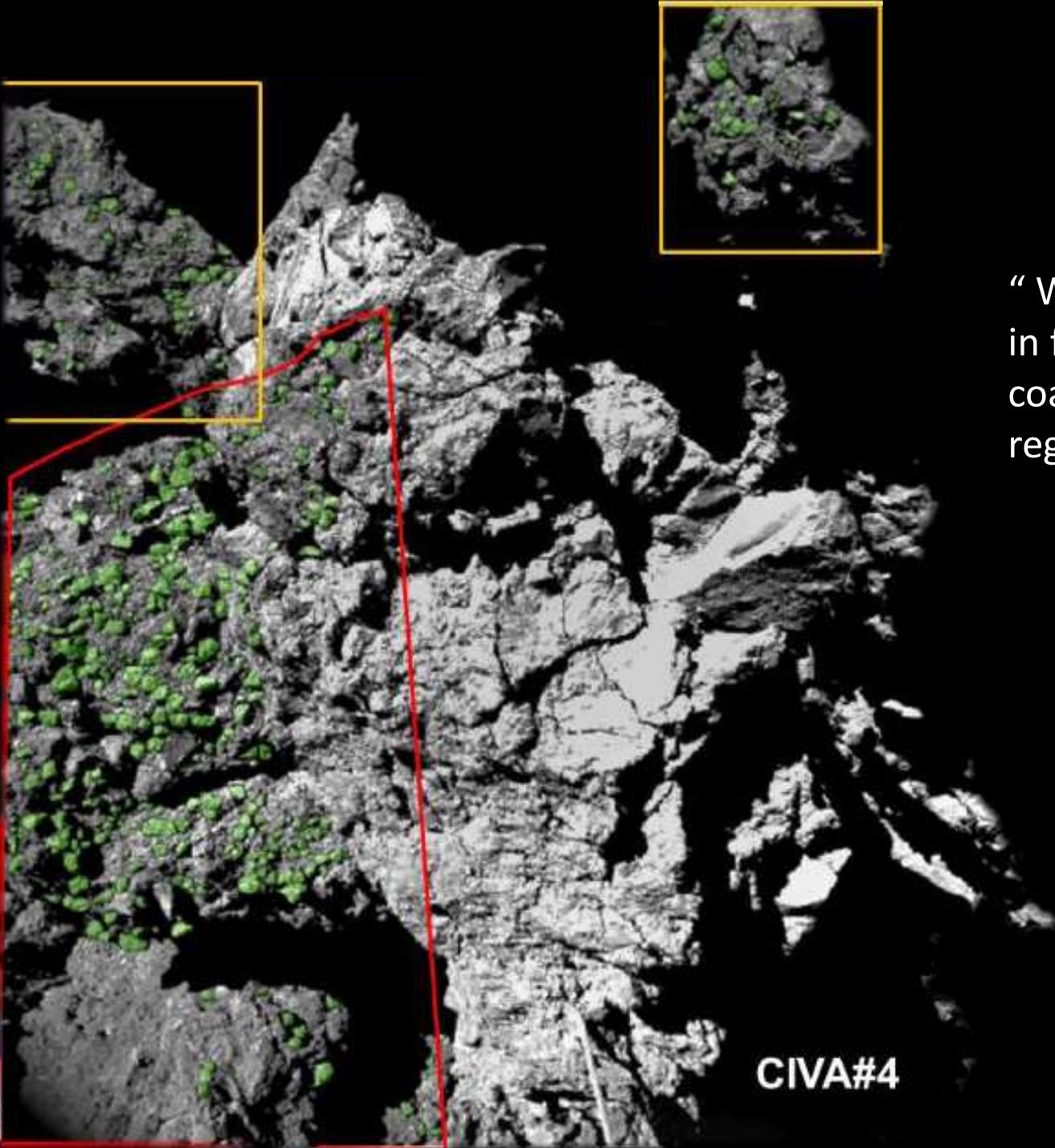
Franceschi et al., in prep.

Surface processes and features are manifold in appearance and nature.

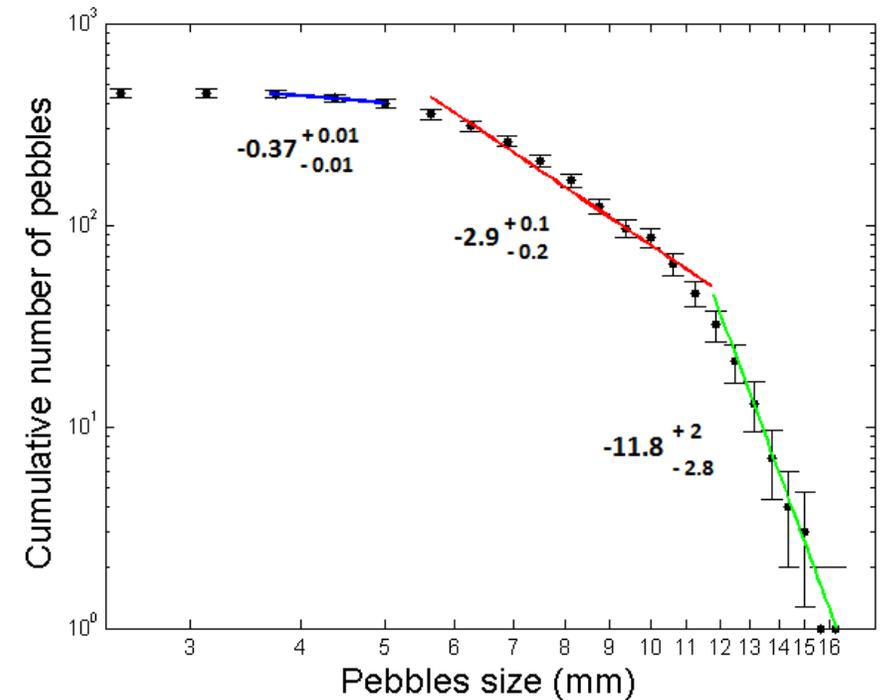


Small-scale structures

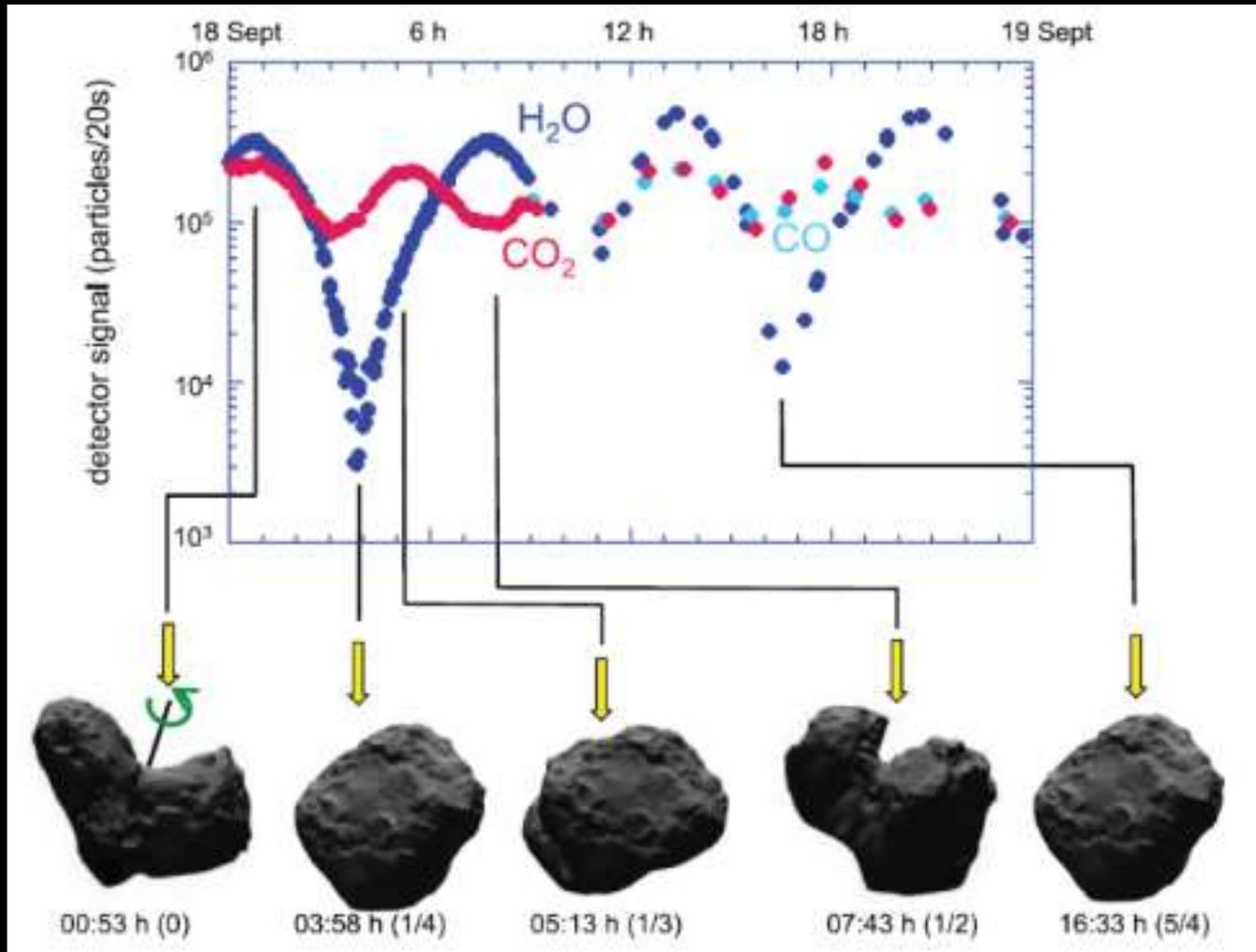
“ We thus speculate that the Abydos landscape could be in favour of pebble accretion model instead of runaway coagulation model with a formation location in the outer region of the Solar System. “



Poulet et al., 2016



Core temperature is low, exact value not clear.



- H_2O and CO_2 not correlated
- Physically unmixed
- Sublimation temperatures of CO , O_2 , N_2 , and Ar ices all in the range $22 < T_s < 25K$
- If these ices are also unmixed, the maximum core temperature could be in the 25 K range
- Further details in Gasc et al. (2017)

Hässig et al.

