

# Cold plasma rotation in the tornado-like prominence of July 13, 2014: a real motion or an illusive effect?

J. Rybák<sup>1</sup>, J. Ambróz<sup>1</sup>, P. Gömöry<sup>1</sup>, J. Kavka<sup>1</sup>, J. Koza<sup>1</sup>, M. Kozák<sup>1</sup>, A. Kučera<sup>1</sup>, P. Schwartz<sup>1</sup>, S. Tomczyk<sup>2</sup>, S. Sewell<sup>2</sup>, G. Capobianco<sup>3</sup>, S. Fineschi<sup>3</sup>, M. Temmer<sup>4</sup>, V. Posch<sup>4</sup>

<sup>1</sup> – Astronomical Institute, SAS, Tatranska Lomnica (Slovakia), <sup>2</sup> – High Altitude Observatory, NCAR, Boulder (USA),  
<sup>3</sup> – INAF - Astrophysical Observatory of Torino, Pino Torinese (Italy), <sup>4</sup> – Institute of Physics, University of Graz, Graz (Austria)

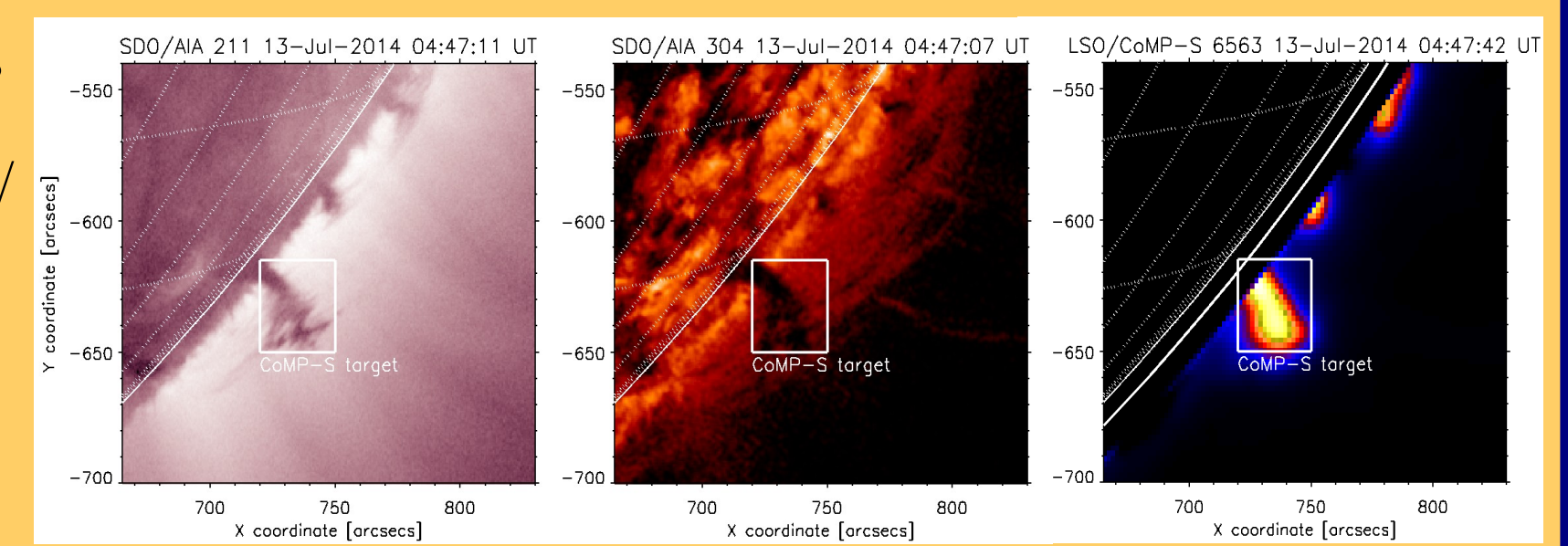
**Abstract:** We analyze in a case study the tornado-like prominence of July 13, 2014 which shows changing position in the SDO/AIA imaging at EUV wavelengths, using the H $\alpha$  2D spectroscopy data acquired with the CoMP-S instrument (Lomnický Stit Observatory - LSO, AISAS, Slovakia). The aim of the study is to address the question whether this structure is a real tornado (prominence leg plasma rotating around central axis) or we just observe illusive signatures of an apparent rotational motion, like oscillation. Our case study results indicate that: a/ the detected Doppler shifts do not show a permanent blue/red-shift pattern along the vertical axis of the structure during the whole 45 min observing time interval, b/ the present variations of the Doppler shifts ( $\pm 4$  km/s) are not in general clearly correlated with the H $\alpha$  integral line emission of the structure, c/ the Doppler shift variations do not show any regular oscillatory behavior. These results lead to conclusion that the Doppler shifts of this particular tornado-like structure cannot be interpreted as real cold plasma rotation around the vertical axis of the structure. However the purely imaging SDO/AIA observations show clear illusive vortical motions in this tornado-like structure. We suggest that the 'vortical illusion' (Panascenzo et al., 2014) - a combination of the counter-streaming flows in the prominence threads and possible radiative transfer effects - are causing the apparent rotational motion of this tornado-like structure.

## Introduction

Recent observations of the SDO/AIA instrument (Lemen et al., 2012) have revitalized phenomenon of solar tornado-like prominences (Pettit, 1932). These tornado-like structures, seen as vertical prominence legs above the solar limb, are showing changing positions in the SDO/AIA imaging suggesting a rotational motion around their axis (Panascenzo et al., 2014; Mgebrishvili et al., 2015; Levens et al., 2016; Schmieder et al., 2017). Such tornado rotation around its axis was spectroscopically confirmed from hot plasma emission surrounding the structure (Su et al. 2014; Levens et al. 2015) using the Hinode/EIS spectrometer (Kosugi et al., 2007) though Young et al. (2012) pointed out that such rotation could be caused by an instrumental effect. On the other hand, the cool material ( $T \sim 10^4$ K) emission does not present clear signatures of rotation in two investigated cases so far (Orozco Suarez et al., 2012; Schmieder et al., 2017). Panascenzo et al. (2014) explained the apparent vortical motion in prominences - only observed in projection at the limb - as counterstreaming flows giving the illusion of rotation. We try to address the question, using the LSO/CoMP-S spectroscopic observations of a similar structure, whether it is a real tornado or we just observe illusive signatures of an apparent rotational motion.

## The target tornado-like structure: SDO/AIA and LSO/CoMP-S data

- July 13, 2014, south from AR 12110, above the SW limb: [+735", -635"]
- visible as a dark structure in all SDO/AIA UV channels (except 9.4 nm)
- the dark structure is rooted vertically in front of the limb (Fig. 1)
- plane-of-sky motion in several parts of the structure (Fig.3, upper panels)
- SDO/AIA dark structure ~ CoMP-S H $\alpha$  integral line emission structure: a clear coincidences of patterns



**Fig. 1.** The target tornado-like structure of 2014/07/13 (04:47 UT, [+735", -635"], position angle 235°): the SDO/AIA 21.1 nm and 30.4 nm channels (left and middle panels) and the LSO/CoMP-S H $\alpha$  integral line emission (right panel).

## LSO instrumentation & CoMP-S observations

### Instrumentation:

- ZEISS 200/3000 Lyot coronagraph (Lexa, 1963)
- CoMP-S instrument: tunable 4-stage Lyot filter with Stokes polarimeter (Kučera et al., 2011)

### CoMP-S Observations:

- date/time/observer: 2014/07/13: 04:47–05:32 UT, J. Kavka
- FoV: 800" x 600", position angle 235°
- spectral line: H $\alpha$  656.28 nm (9 points, 0.025 nm step, 98 scans)
- filter passband width: FWHM = 0.081 nm (measured)
- exposure time: 50 ms, scan time: 22 s
- detector: PCO sCMOS camera: 2560\*2160 6.5  $\mu$ m pixels

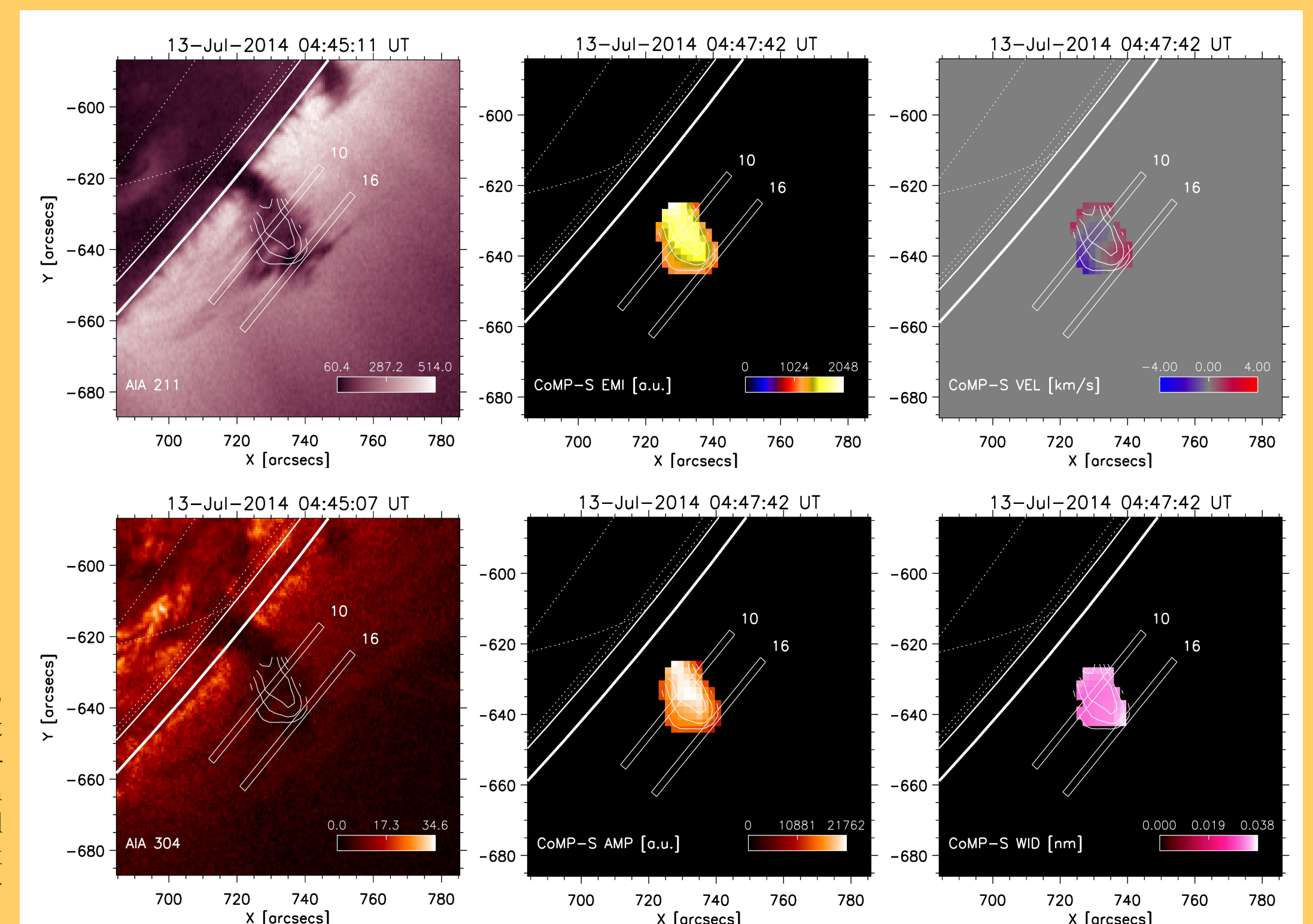
### CoMP-S Data reduction:

- photometric data reduction
- Stokes I parameter data only
- 4x4 pixel binning for the resulting 1.33"/px spatial sampling
- sky background and the instrumental scattered light removed
- post-facto image jitter correction
- rest line wavelength correction
- Gaussian profile fitting to the individual H $\alpha$  spectral line profiles
- temporal (3-scan) smoothing of the Gaussian fit parameters
- sub-arcsecond co-alignment to the SDO/AIA coordinates

## The tornado-like structure zoom

- A zoom area just around the SDO/AIA image of the tornado-like structure itself (Fig. 2)
- the CoMP-S data mask: a radial overoccluding for 13" above the solar limb
- a manually introduced H $\alpha$  integral line emission mask ( $> 900$  a.u.) leading to the CoMP-S structure body covering the whole SDO/AIA 21.1 nm channel structure body (Fig.2)
- the masked zoom area filled with 20 adjoining 2" wide and 52" long pseudo-slits (parallel to the local limb) for an individual time-space analysis of the target tornado-like structure evolution
- two pseudo-slits (#10 and #16) selected in the target structure for an illustrative display of the time-space cuts (Fig.3)

**Fig. 2.** The target tornado-like structure at the beginning of the CoMP-S measurements: the SDO/AIA 21.1 nm and 30.4 nm channels (left panels) and the LSO/CoMP-S H $\alpha$  line Gaussian fit parameters: EMI – the integral line emission, VEL – the Doppler shift, AMP – the Gaussian amplitude, WID – the line width (middle and right panels). The H $\alpha$  total line emission mask is applied (white contours in all panels). The thick white line marks the Lyot coronagraph diaphragm edge. Two narrow rectangles give positions of the pseudo-slits #10 and #16.

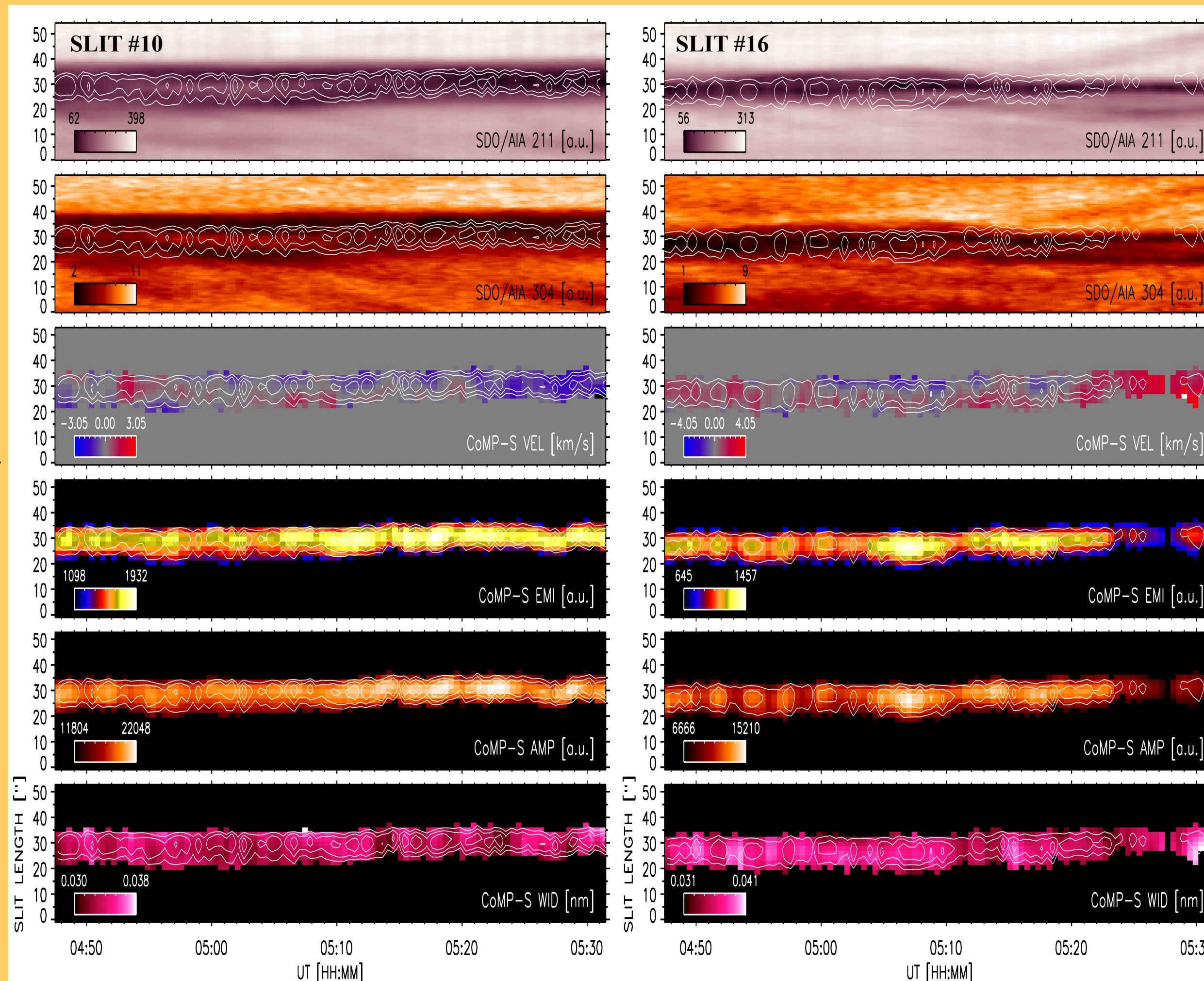


## Motions in the target tornado-like structure

- Two representative pseudo-slits (#10 and #16) selected for presentation of results in a core and at the top of the CoMP-S structure body (Fig.3, left and right panels, respectively)
- clear apparent vortical motions in the SDO/AIA intensity movies<sup>1</sup> lead to a typical time-space behaviour of this tornado-like structure: variable/inclined tracks in the intensity variations (Fig.3, the top panels)
- the H $\alpha$  line emission location is changing relatively to the AIA 21.1 and 30.4 nm dark structure (Fig. 3, overlays in the top panels)
- the Doppler shifts **do not show a permanent blue/red-shift pattern** along the vertical axis of the structure during the whole observing time interval although there are short time intervals ( $\sim 10$  minutes) of the opposite Doppler shifts across the structure
- in general, the Doppler shifts variations are not clearly matching the H $\alpha$  integral line emission or the EUV AIA intensity behaviour
- the Doppler shift variations do not show any regular oscillatory behavior in time

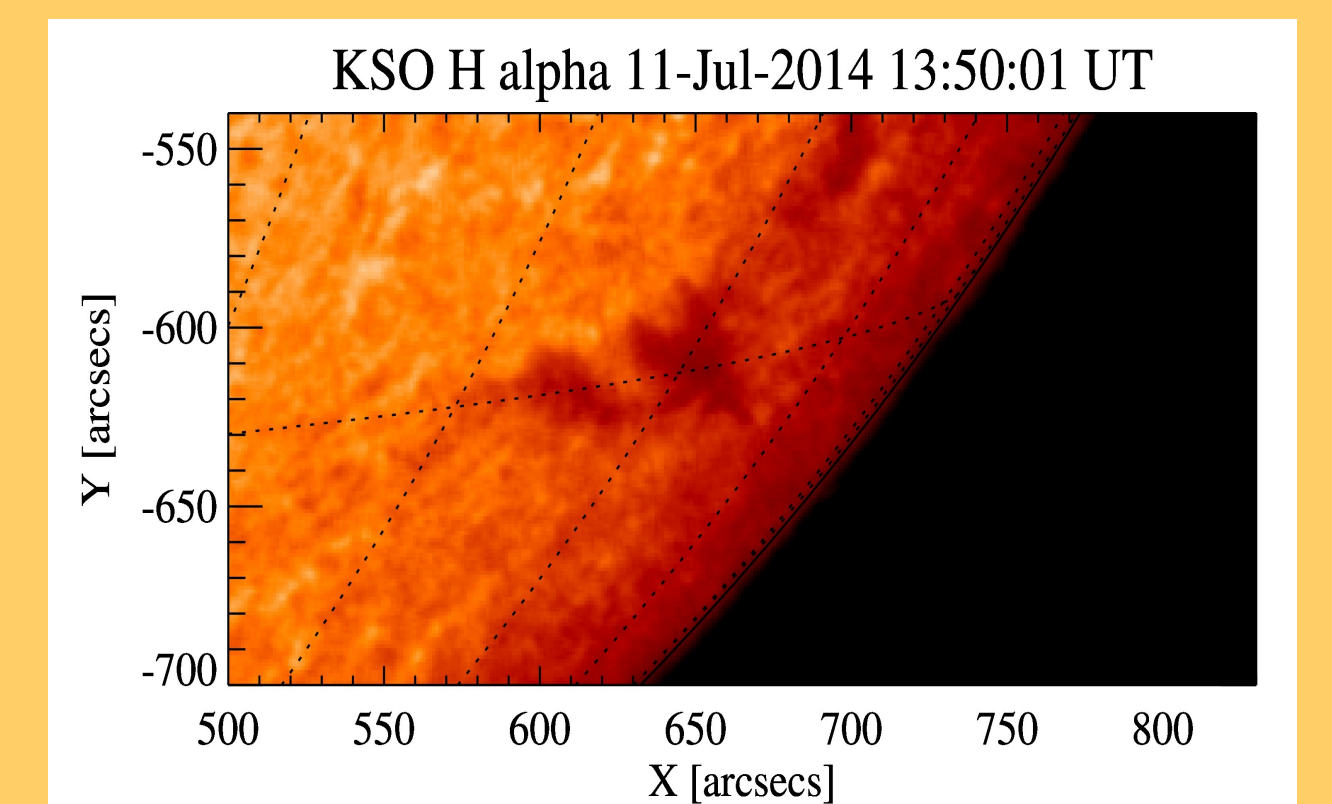
<sup>1</sup> - [https://www.astro.sk/~choe/open/comp-s\\_ho\\_obs\\_prog\\_002/event\\_47\\_20140713/sdo\\_aia/](https://www.astro.sk/~choe/open/comp-s_ho_obs_prog_002/event_47_20140713/sdo_aia/)

**Fig. 3.** The pseudo-slits #10 and #16: the time-space plots of the SDO/AIA 21.1 nm and 30.4 nm channel intensities (upper two panels) and the LSO/CoMP-S H $\alpha$  line Gaussian fit parameters: EMI – the integral line emission, VEL – the Doppler shifts, AMP – the line amplitude, WID – the line width (bottom panels). White overlaying contours: behaviour of the H $\alpha$  integral line emission.



## Conclusions

Our results lead to the conclusion that the Doppler shifts of the particular target tornado-like structure cannot be interpreted as real cold plasma rotation around the vertical axis of the structure. However, the purely imaging SDO/AIA observations show clear illusive vortical motions in this tornado-like structure. This combination of results derived from the cool plasma spectroscopy and the SDO/AIA imaging bears a resemblance to results presented already for a similar structure by Schmieder et al. (2017). We suggest that the "vortical illusion" (Panascenzo et al., 2014) - a combination of the counter-streaming flows in the prominence threads - and possible radiative transfer effects in the cold plasma blobs in the prominence (Fig.4) - are causing the apparent rotational motion of this tornado-like structure.



**Fig. 4.** The filament corresponding to the target tornado-like structure: the KSO H $\alpha$  image, taken more than 1.5 days before the presented SDO/AIA imaging and the CoMP-S spectroscopic observations. showing the individual filament blobs located mostly in the E-W direction.

## References

- Lexa, J., 1963, *BAC* 14, 107
- Kučera, A., Ambróz, J., Gömöry, et al., 2011, *CAOSP* 40, 135
- Kosugi, T., Matsuzaki, K., Sakao, T., et al., 2007, *Sol. Phys.* 243, 3
- Lemen, J.R., Title, A.M., Akin, D.J., et al., 2012, *Sol. Phys.* 275, 17
- Levens, P.J., Labrosse, N., Fletcher, L., et al., 2015, *A&A* 582, A27
- Levens, P.J., Schmieder, B., Labrosse, N., et al., 2016, *ApJ* 818, 31
- Mgebrishvili, I., Zaqarashvili, T. V., Kukhianidze, V., et al., 2015, *A&A* 810, 89
- Orozco Suarez, D., Asensio Ramos, A., & Trujillo Bueno, J., 2012, *ApJ* 761, L25
- Panascenzo, O., Martin, S.F., Velli, M., 2014, *Sol. Phys.* 289, 603
- Pettit, 1932, *ApJ* 76, 9
- Schmieder, B., Mein, P., Mein, N., et al., 2017, *A&A* 597, A109
- Su, Y., Gömöry, P., Veronig, A., et al., 2014, *ApJ* 785, L2
- Young, P. R., O'Dwyer, B., Mason, H.E., 2012, *ApJ* 744, 11

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